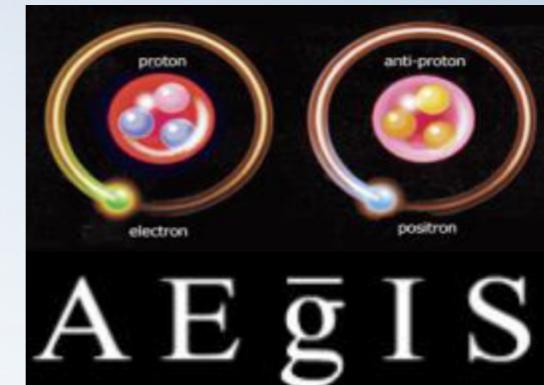




Hyperfine Structure of Antihydrogen

ERC Advanced Grant

PI: Prof. Dr. Eberhard Widmann



Antimatter gravity measurement in a beam - the AEg IS experiment

E. Widmann

Stefan Meyer Institute for Subatomic Physics, Vienna
on behalf of the AEg IS collaboration



INPC2016
Adelaide, Australia
15. Sep. 2016

A E ē I S collaboration



ÖAW



Stefan Meyer Institute



CERN



Czech Technical University



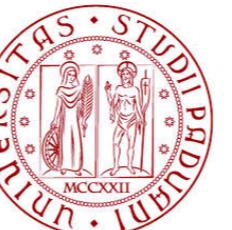
ETH Zurich



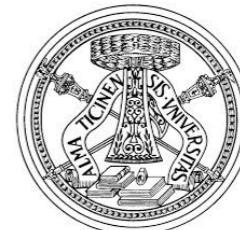
University of Genova



University of Milano



University of Padova



University of Pavia

И
И
Н
Р
И
Р

Institute of Nuclear Research of the Russian Academy of Science



Max-Planck Institute
Heidelberg



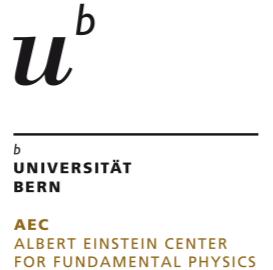
Politecnico di
Milano



University College London



University of Bergen



University of Bern



University of Brescia



Heidelberg University



University of Lyon 1



University of Oslo



University of
Paris Sud



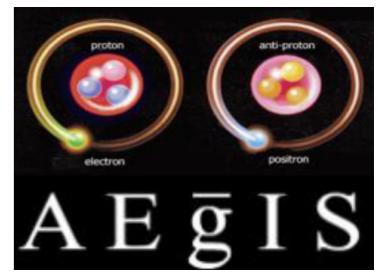
University of
Trento



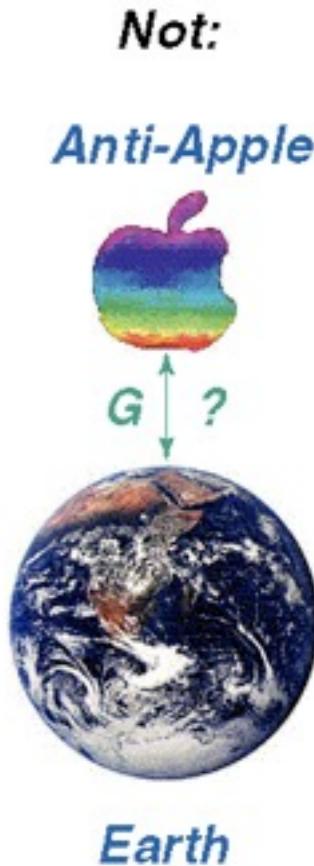
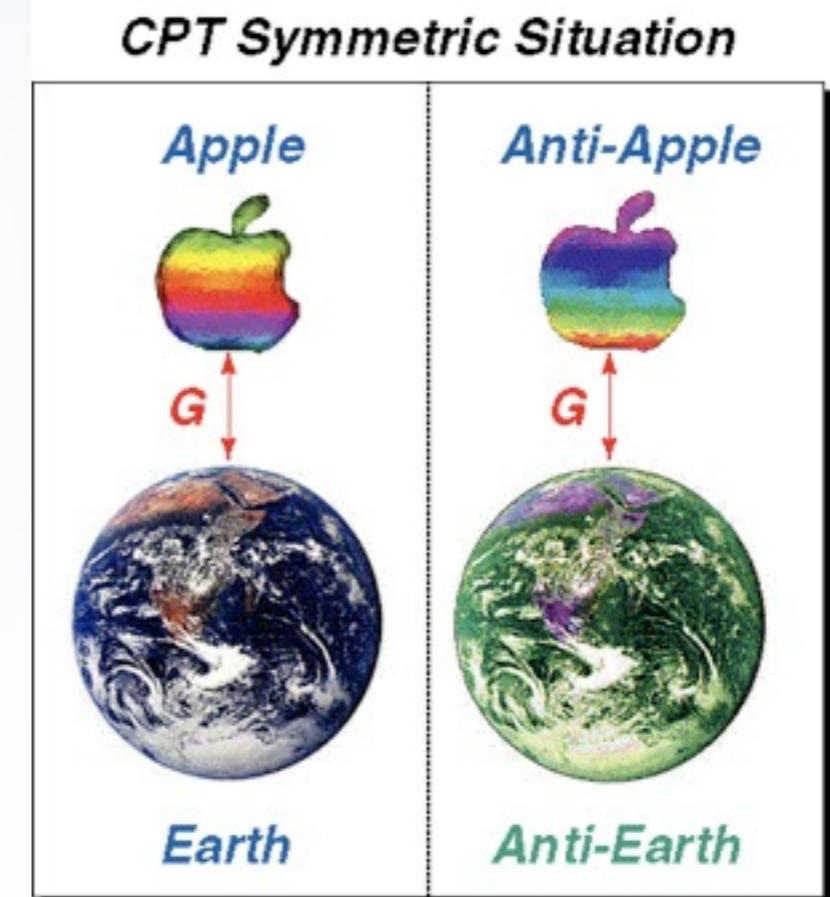
INFN sections of:
Genova, Milano,
Padova, Pavia,
Trento



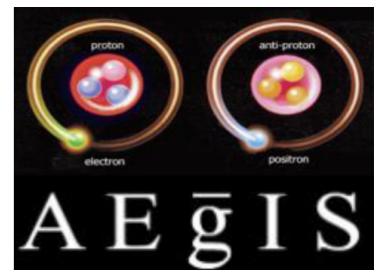
Gravitational Acceleration of Antimatter



- No direct test of CPT
 - Weak equivalence principle
- no precise experimental test available
- Highest precision reachable with neutral antimatter
- AEg IS
 - Antimatter Experiment - gravity, Interferometry, Spectroscopy



Antimatter and gravity

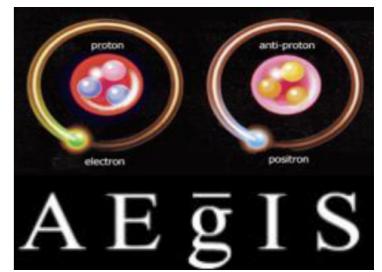


- Antigravity: $g_{\text{matter}} = -g_{\text{antimatter}}$
 - separation of matter and antimatter in Universe
- Quantumgravity
 - Graviton ($S=2$) → adds Gravivector ($S=1$), Graviscalar ($S=0$)
 - simplest case: static potential

$$V = -Gm_1 m_2 / r$$

- a: Gravivector, b: Graviscalar
- – attractive (matter-matter), +: repulsive: matter-antimatter
- matter experiments: $|a-b|$
- antimatter: $a+b$

AEg IS Goal

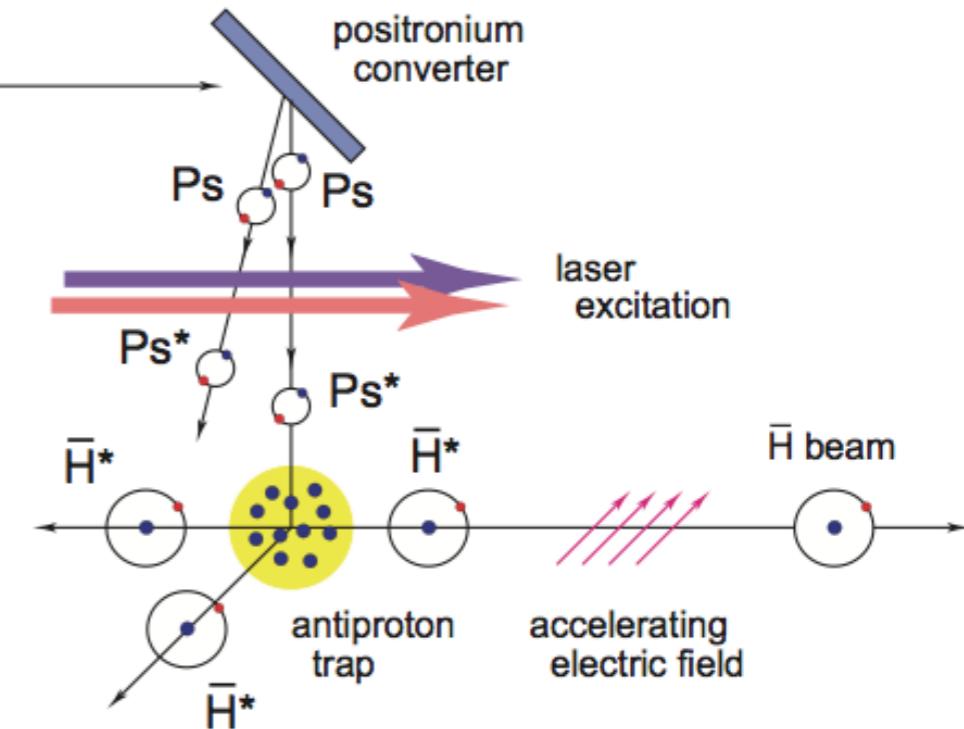


- First goal is to measure g to $\mathcal{O}(1\text{-}10\%)$
 - in the free fall of cold H atoms
- Test of WEP on antimatter in the Earth's gravitational field
 - First direct measurement for antimatter - no assumptions
 - Large amount of antimatter atoms needed
- AEg IS method
 - Produce cold Rydberg H *
 - Accelerate H * to produce neutral beam
 - Measure gravitational deflection of H beam using a Moiré deflectometer

Production of H atoms

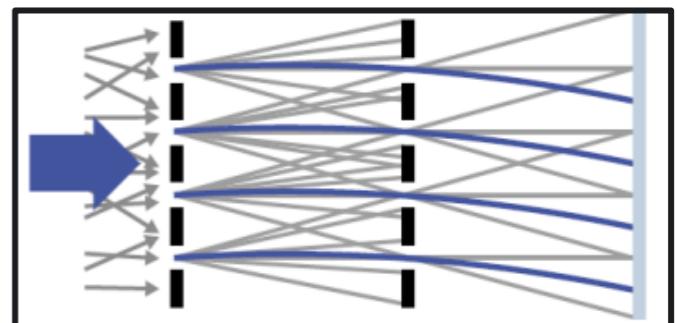
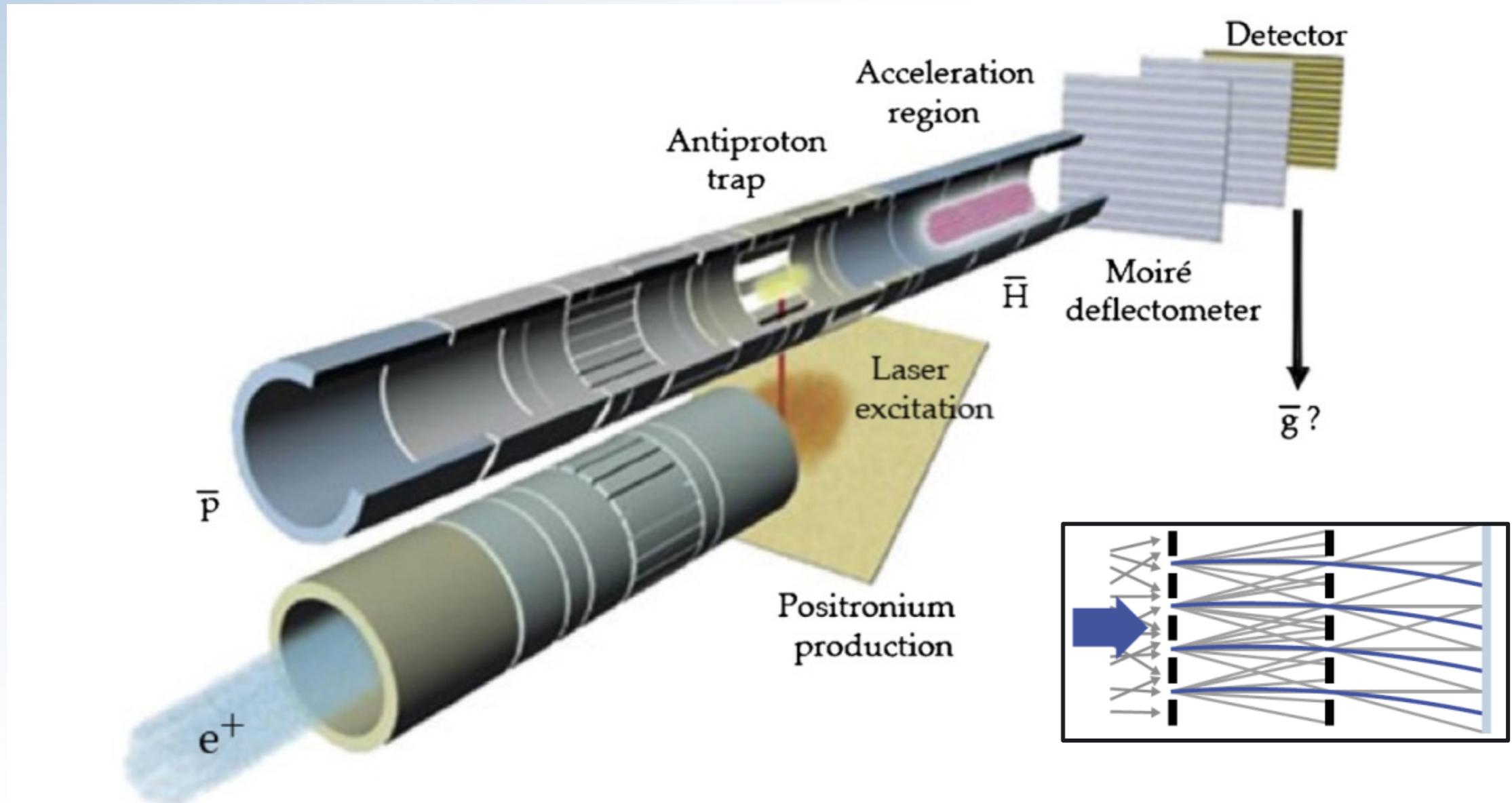
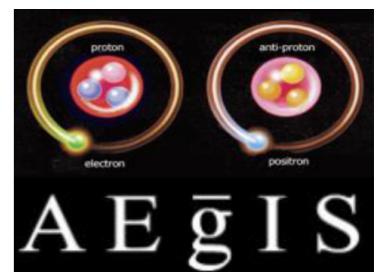


- Charge exchange with Ps^*
 - Capture p from CERN-AD
 - Cool p
 - Ps production by e^+ on SiO_2
 - Ps laser excitation to Rydberg state
 - Interaction Ps* with p cloud
 - Recombination: resonant charge exchange

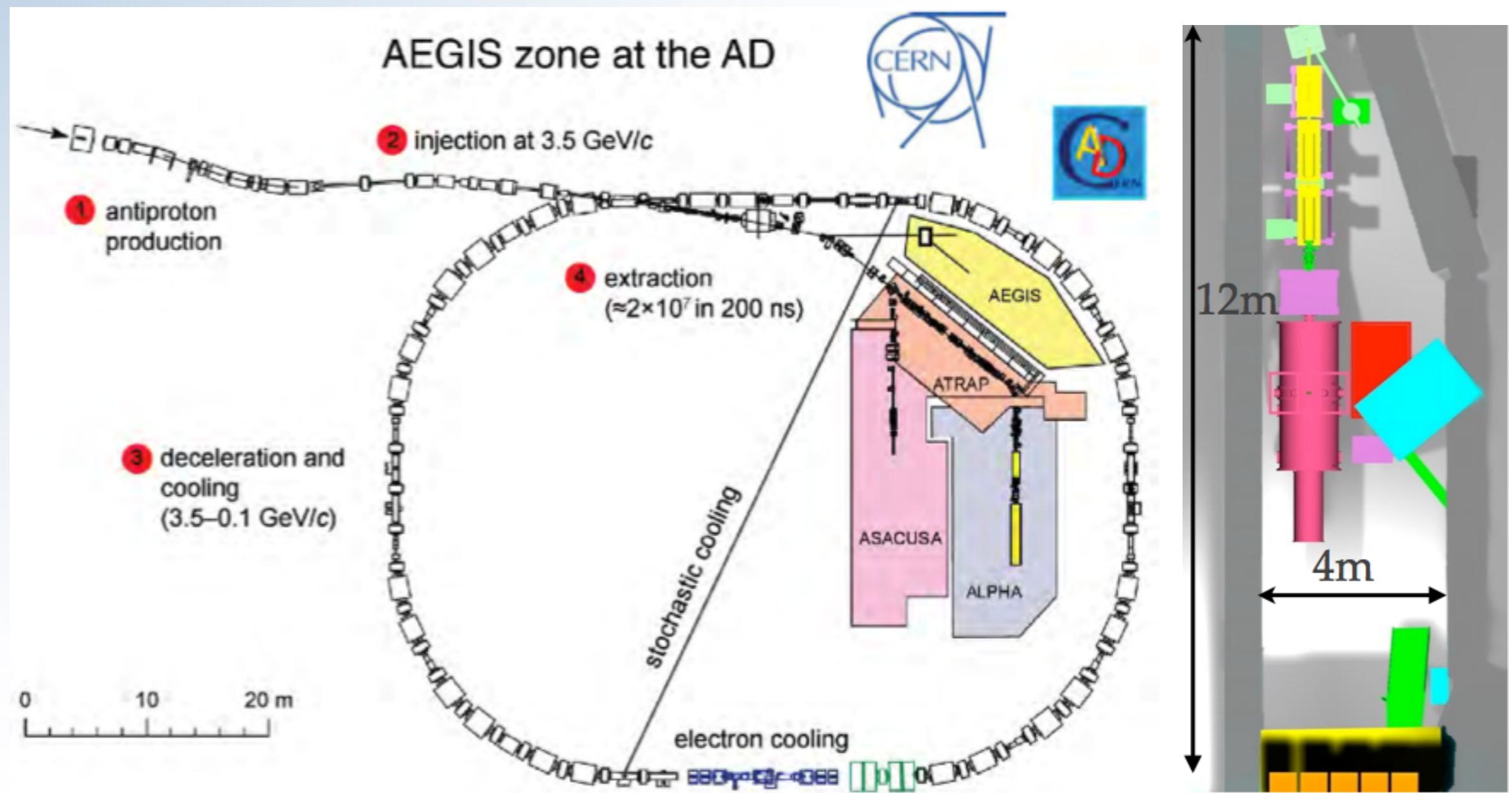
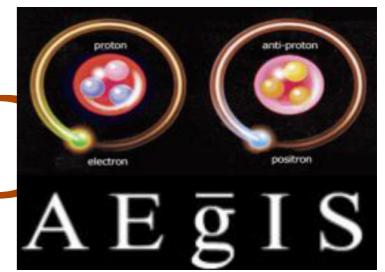


enhanced formation rate:
 $\sigma \sim n^4$
 $\text{Ps}(n_1) \rightarrow (n_2)$

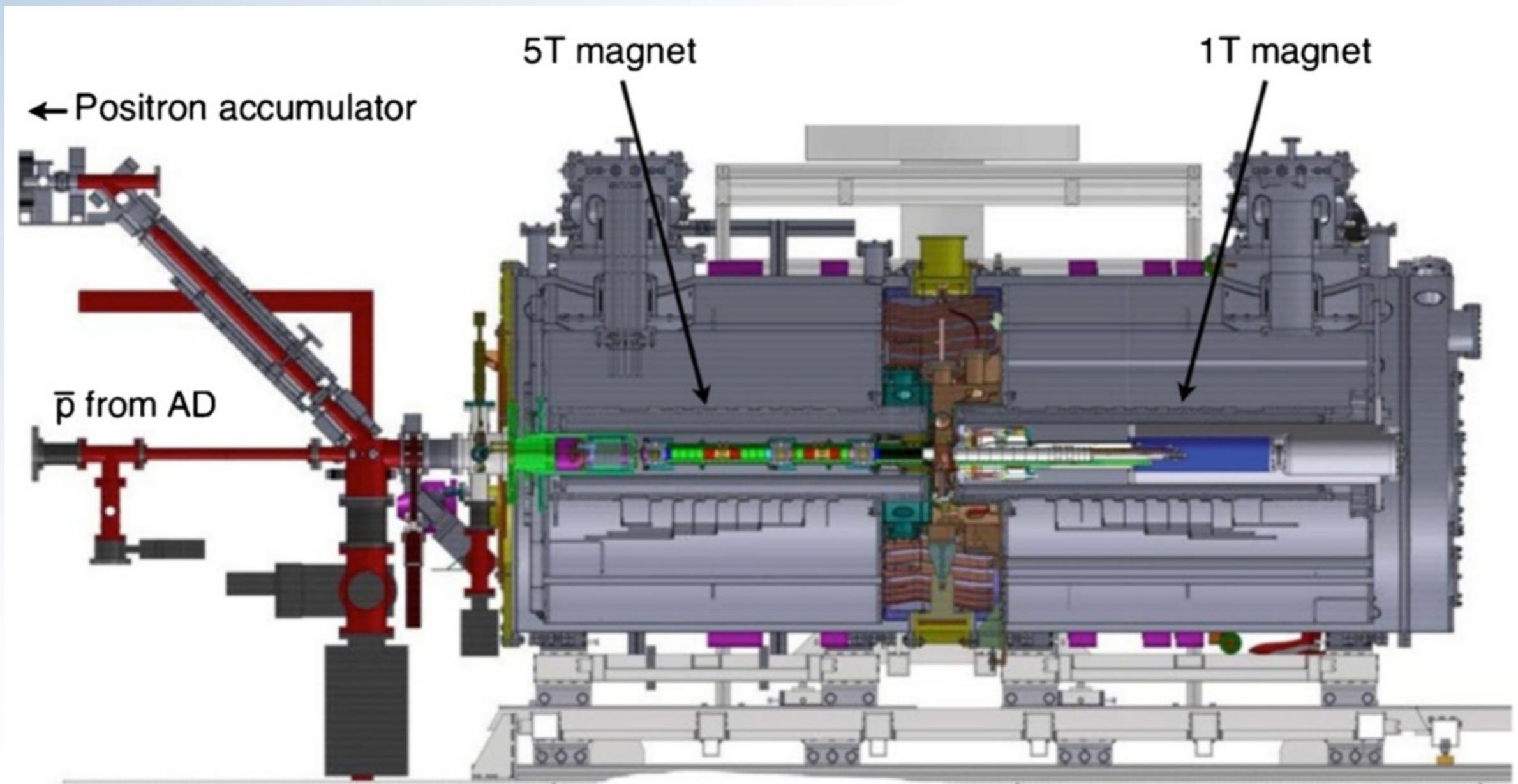
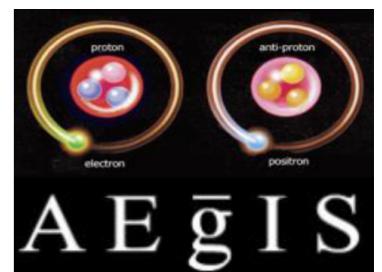
AEGIS principle



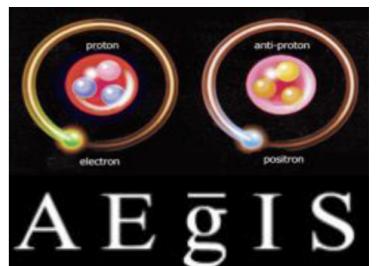
Experimental setup: CERN-AD



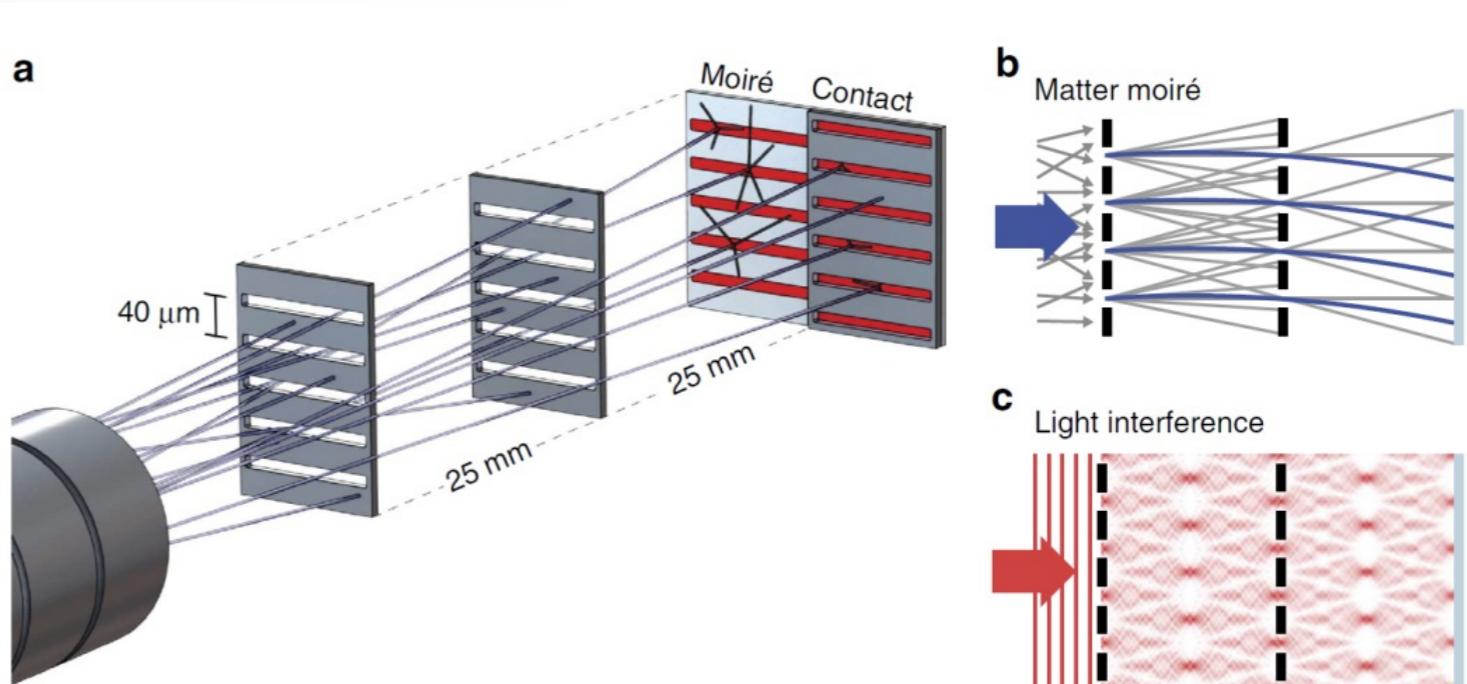
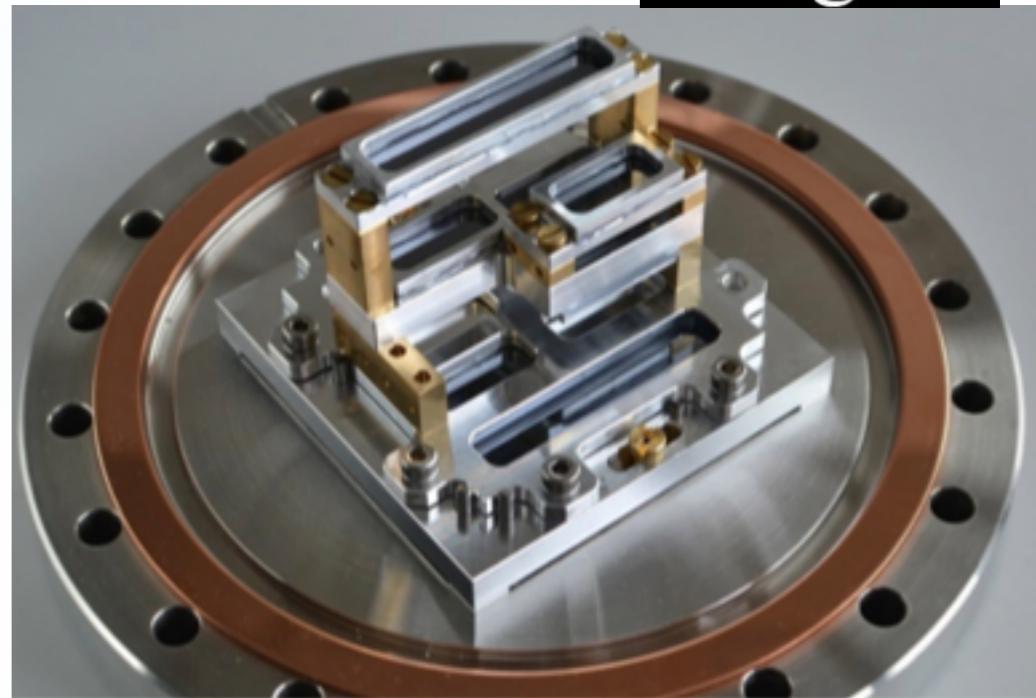
AEG IS Setup



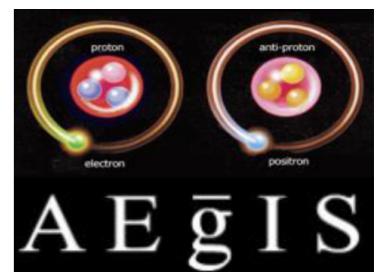
Gravity measurement - proof of principle



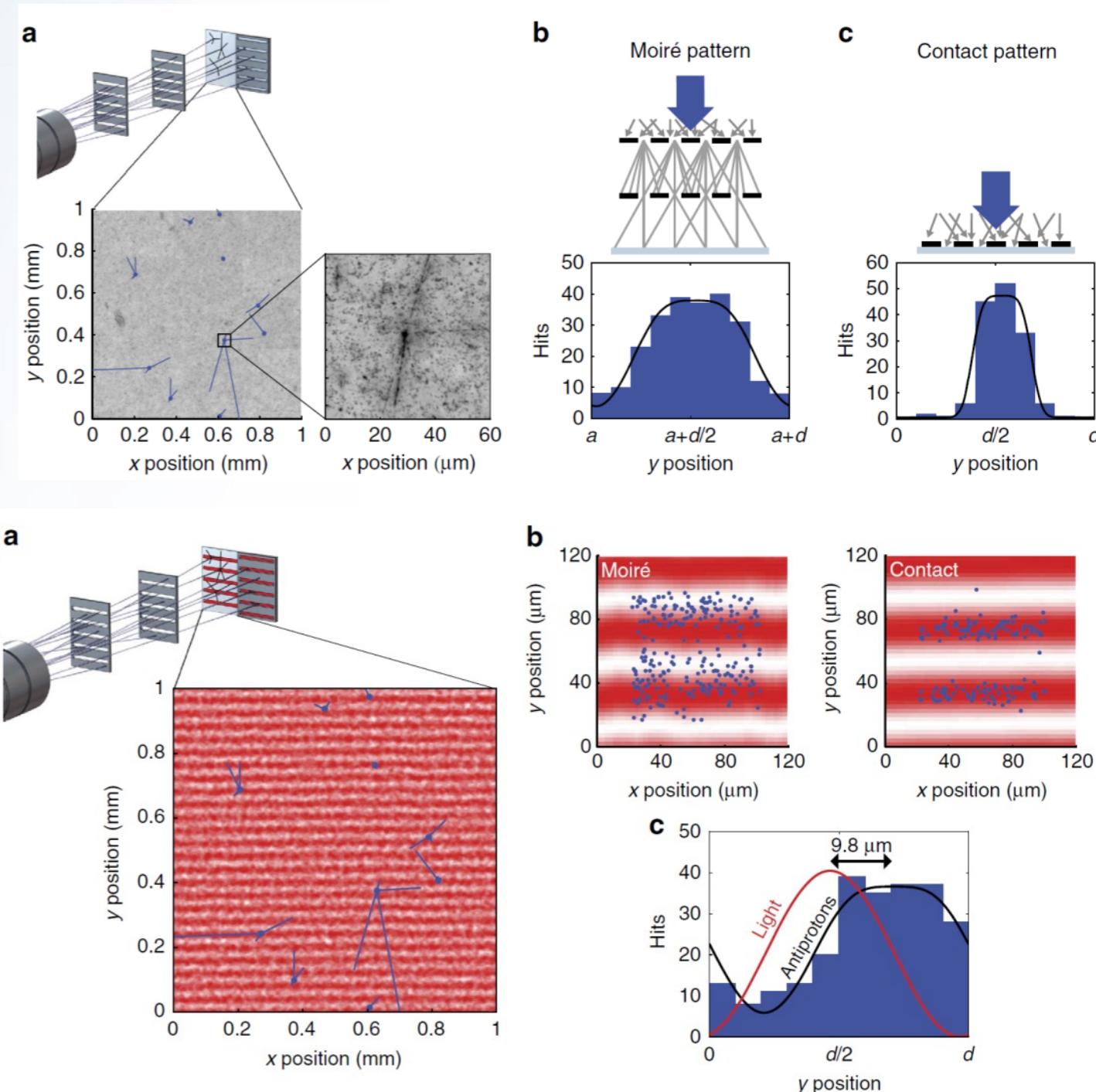
- Mini-moiré deflectometer
 - distance 25 mm
 - slit 12 μm , pitch 40 μm , 100 μm thick
 - p beam $E \sim (100 \pm 150)$ keV traversing 1T magnet
 - light reference:
Talbot-Lau
 - emulsion detector



p deflectometer result



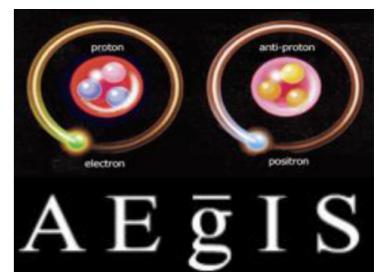
- Pattern observed
 - Shift between p and light observed
- $\Delta y = 9.8 \pm 0.9(\text{stat}) \pm 6.4(\text{syst}) \mu\text{m}$
- consistent with residual B, E fields
 - sensitivity of μm reached
 - H beam case
 - velocity $\times 10^{-4}$
 - distance $\times 40$
 - Force 10^{-10}



Aghion, S. et al. Nature Communications 5, 4538 (2014)

Ps production: SSPALS measurement

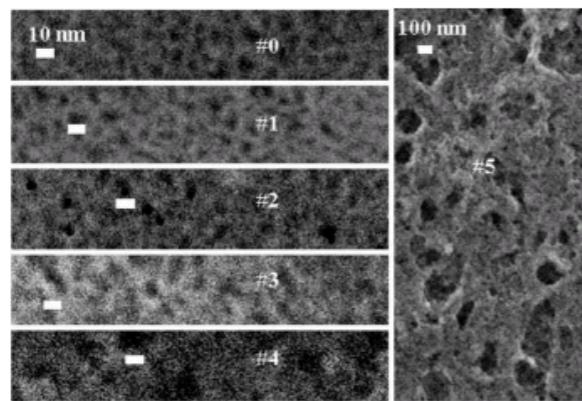
Cassidy D B et al., NIMB 2007, 580, 1338



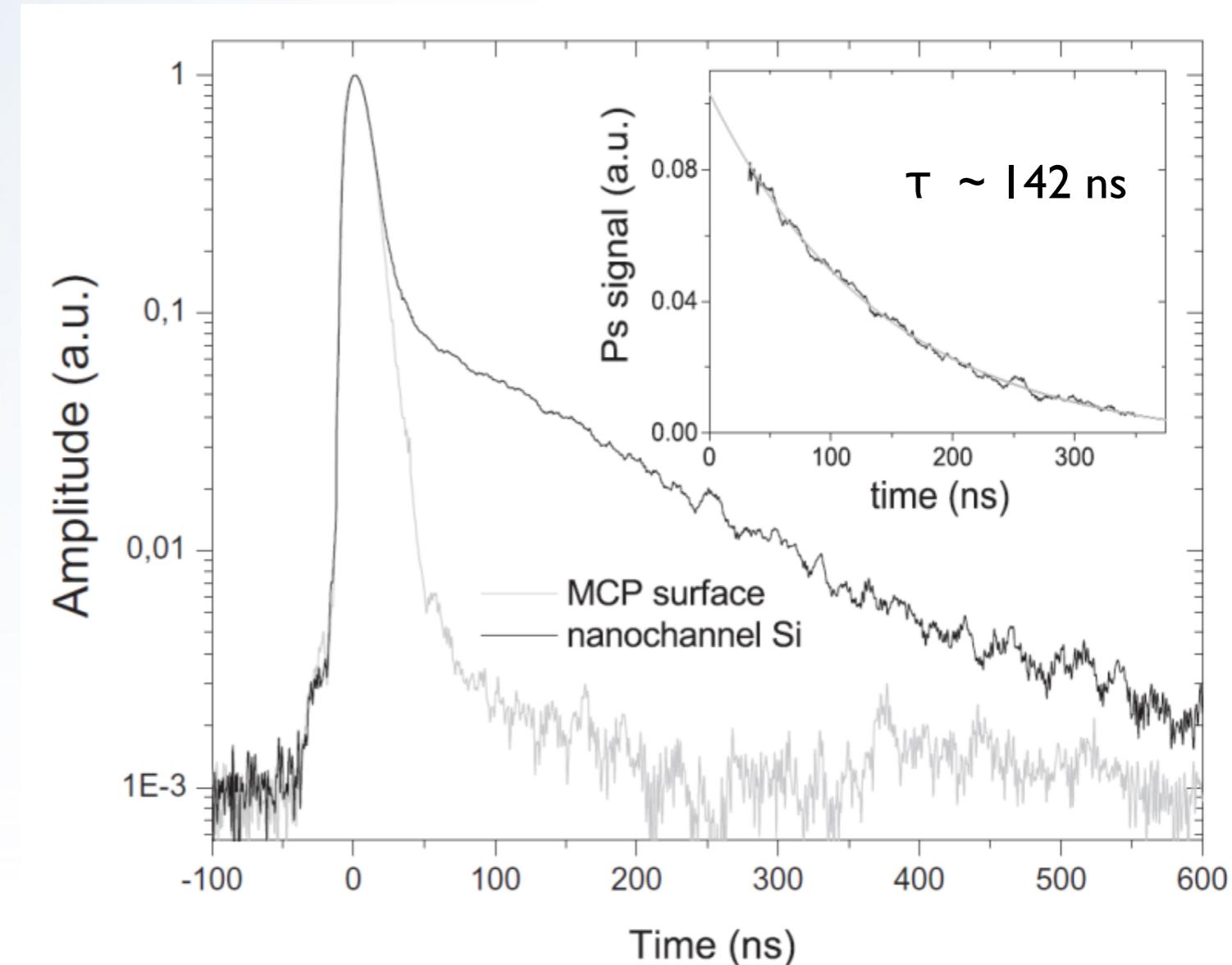
$\sim 4 \times 10^6$ Ps atoms in vacuum

PbWO₄ detector

e+/Ps converter



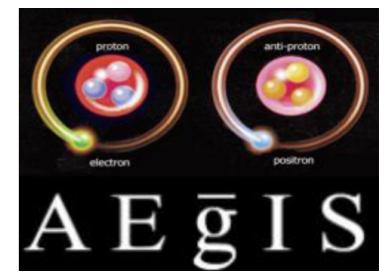
Nanochanneled Si



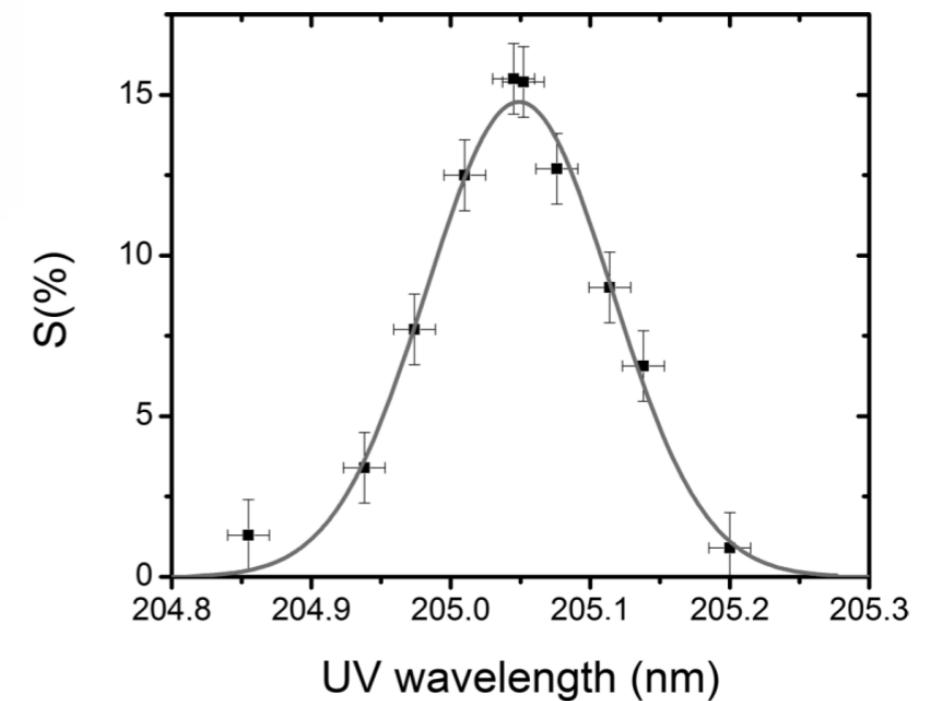
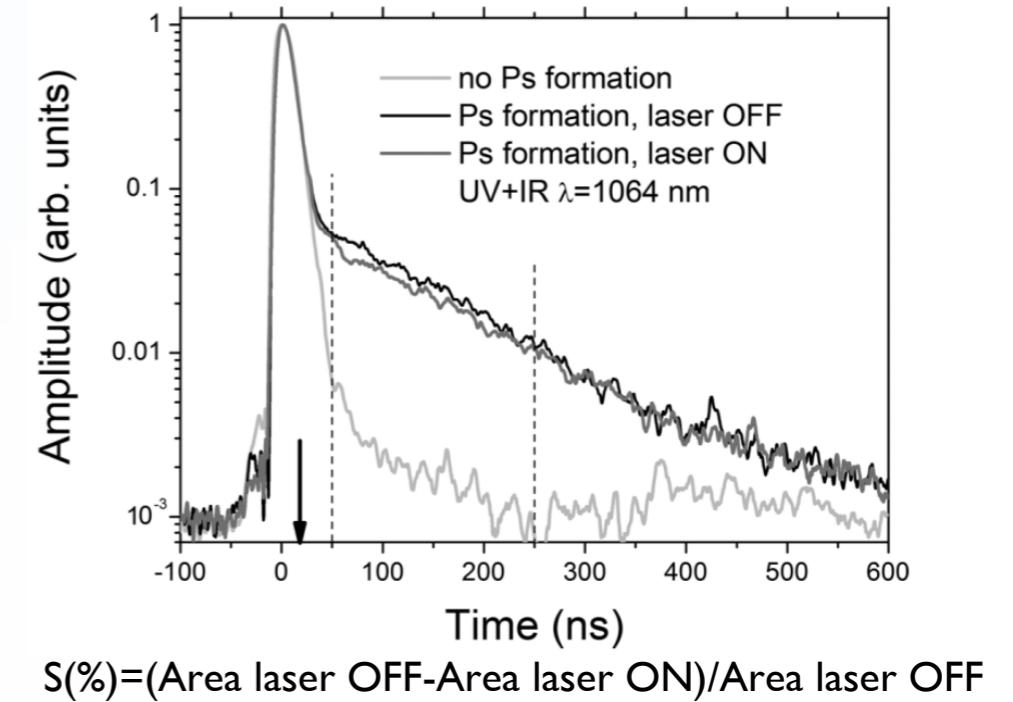
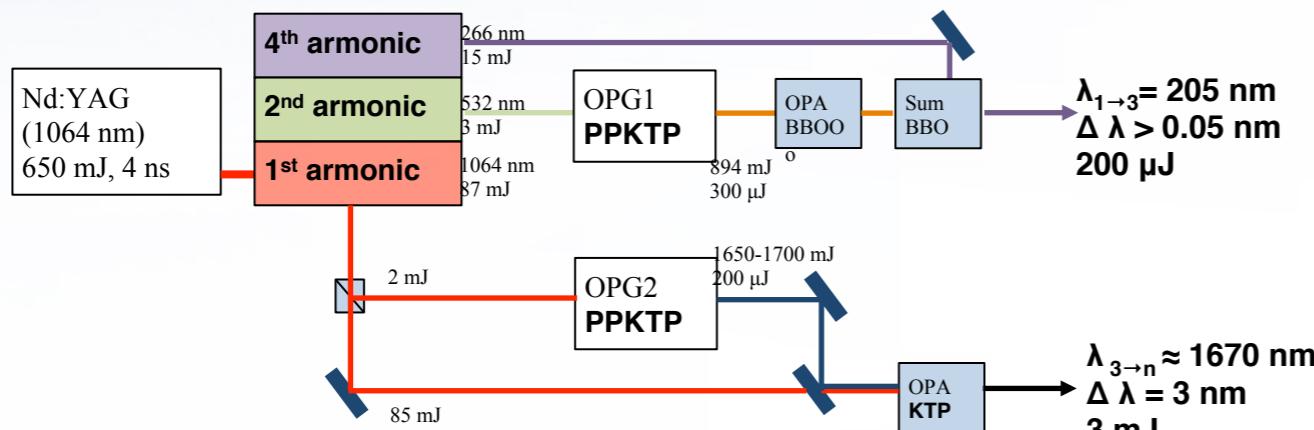
Mariazzi S et al., Phys. Rev. B 2010, 81, 235481

Aghion S et al., NIM B, 362 86–92 (2015)

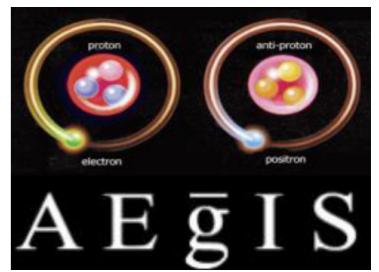
Demonstration of Ps n=3 laser excitation



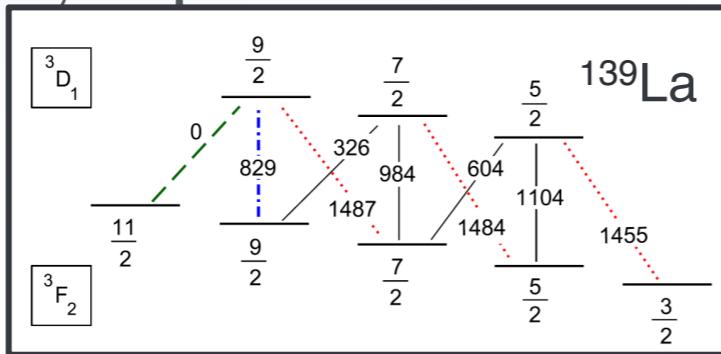
- 3P excitation line centered at 205.05 ± 0.02 nm
- excitation-ionization efficiency ~15% (determined by ionization by IR pulse)
- Evidence for n=15-18 Rydberg excitation by 2-step laser excitation



Cooling of p

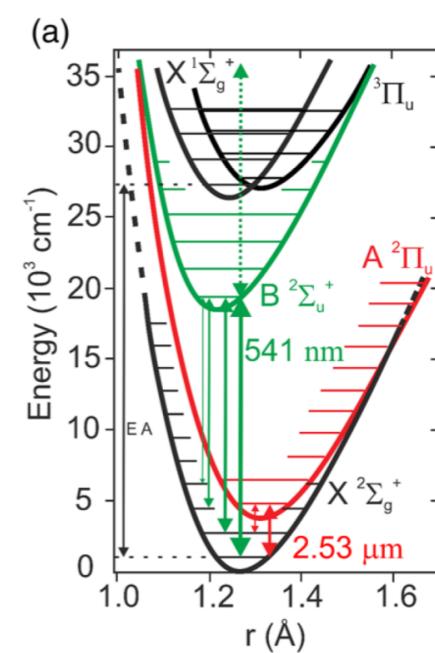


- Cold p to maximize flux
 - Final goal: $T \sim 100$ mK, 1st phase $T \sim 7$ K
- Cooling mechanisms
 - Radiative electron cooling
 - Evaporative / adiabatic p cooling
 - Sympathetic resistive cooling of p with e⁻ cooled by resistive cooling
 - Sympathetic laser cooling with negative ions

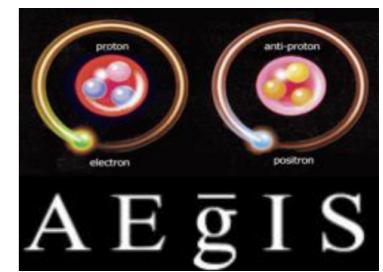


E. Jordan *et al.*,
PRL **115** 113001 (2015)

- Sympathetic cooling with $C\downarrow 2\uparrow -$



Summary



- AEg IS aims at first ever direct measurement of deflection of antihydrogen beam by the Earth's gravitational field
 - Test of WEP with antimatter
- Milestone Ps excitation achieved
- Experiment is taking beam
 - 2016 goal: H production at T~7 K (current min. temperature)
 - later stage → 100 mK
 - H beam production
 - Gravity measurement
 - H -HFS spectroscopy



Bundesministerium für
Wissenschaft, Forschung und Wirtschaft



ERC Advanced Grant 291242
HbarHFS
www.antimatter.at
PI EW



E. Widmann



SPARES





AEGIS results



ALPHA GRAVITY MEASUREMENT

- release trapped \bar{H}
- too hot for gravitational force
 - limits on ratio inertial and gravitational mass
- Gravity
 - no systematics
 - $m_{\text{grav}}/m_{\text{inertial}} < 75$
 - with systematics
 - $m_{\text{grav}}/m_{\text{inertial}} < 110$
- Antigravity
 - $m_{\text{grav}}/m_{\text{inertial}} > -65$

ARTICLE

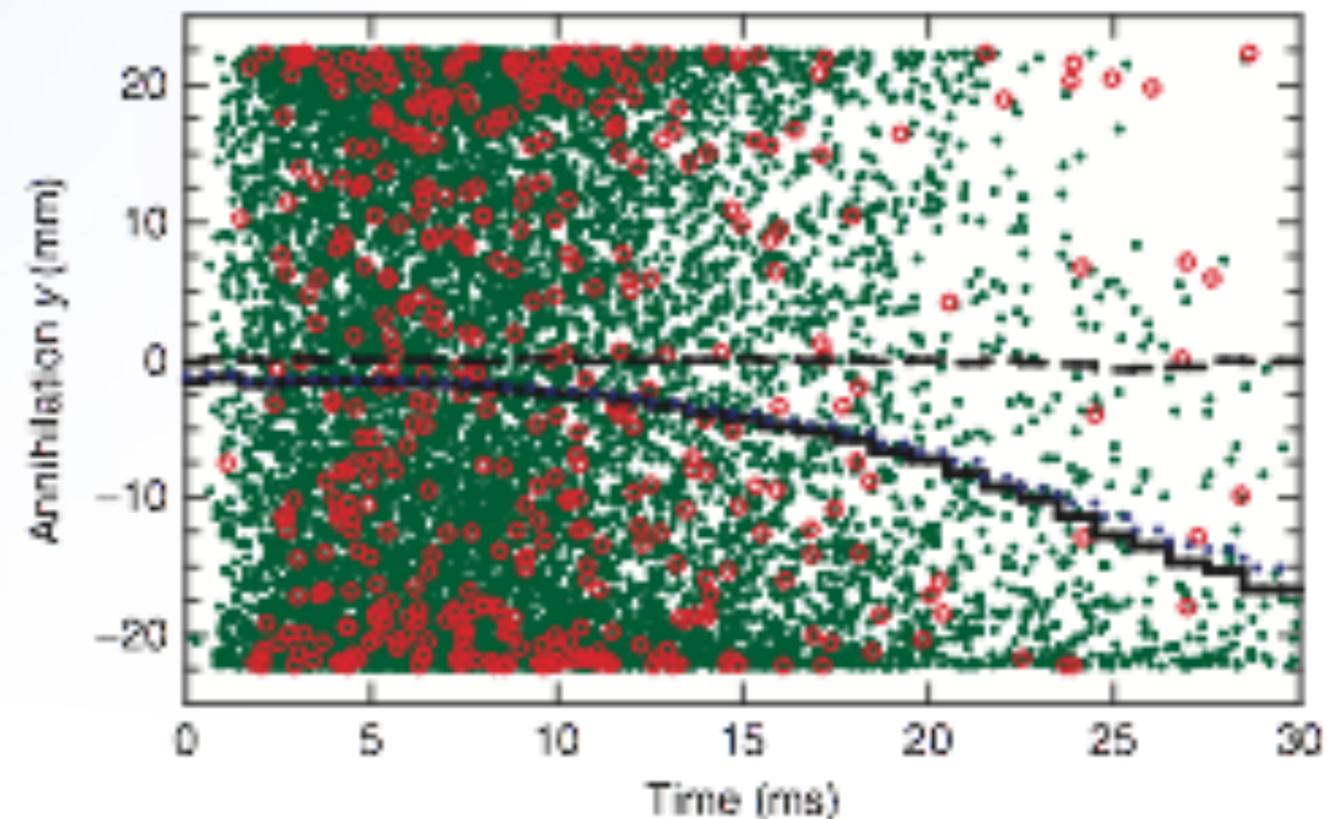
Received 14 Jan 2013 | Accepted 22 Mar 2013 | Published 30 Apr 2013

DOI: 10.1038/ncomms2787

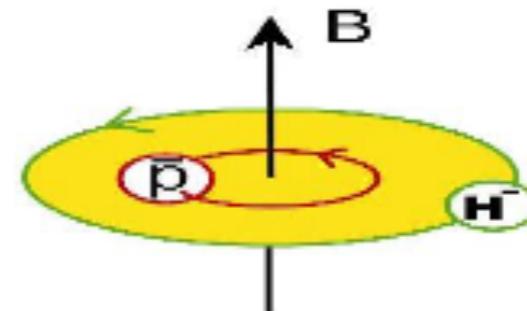
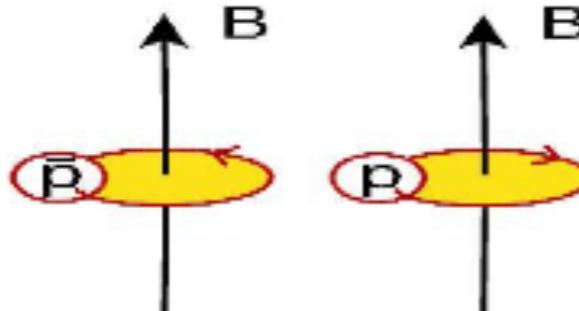
OPEN

Description and first application of a new technique to measure the gravitational mass of antihydrogen

The ALPHA Collaboration* & A.E. Charman¹



GRAVITATIONAL REDSHIFT OF CLOCKS



Gravitational red shift for a clock: $\Delta\omega / \omega = g h / c^2$

→ Antimatter and matter clocks run at different rates
if g is different for antimatter and matter

$$\frac{\Delta\omega_c}{\omega_c} = 3(\kappa - 1) \frac{U}{c^2}$$

for tensor gravity
(would be 1 for scalar gravity)

Hughes and Holzscheiter,
Phys. Rev. Lett. 66, 854 (1991).

grav. pot. energy difference
between empty flat space time
and inside of hypercluster of galaxies

Experiment: TRAP Collaboration, Phys. Rev. Lett. 82, 3198 (1999).

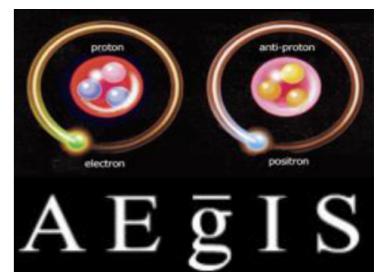
$$\frac{\Delta\omega_c}{\omega_{\partial c}} < 10^{-10} \quad \dashrightarrow \quad \kappa = 1 \pm (< 10^{-6})$$



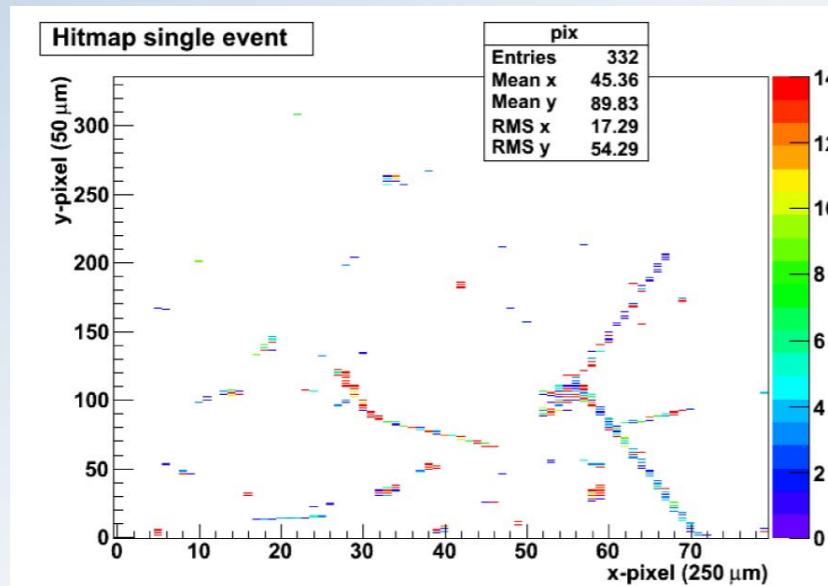
**THANK YOU VERY MUCH FOR YOUR
ATTENTION!**



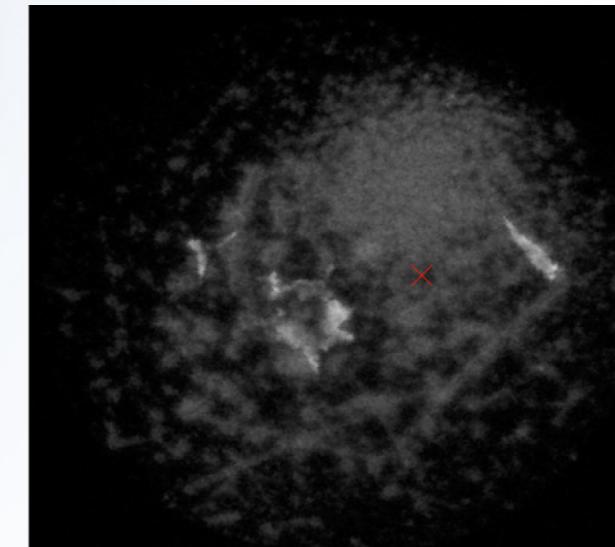
Gravity measurement - H^- detectors



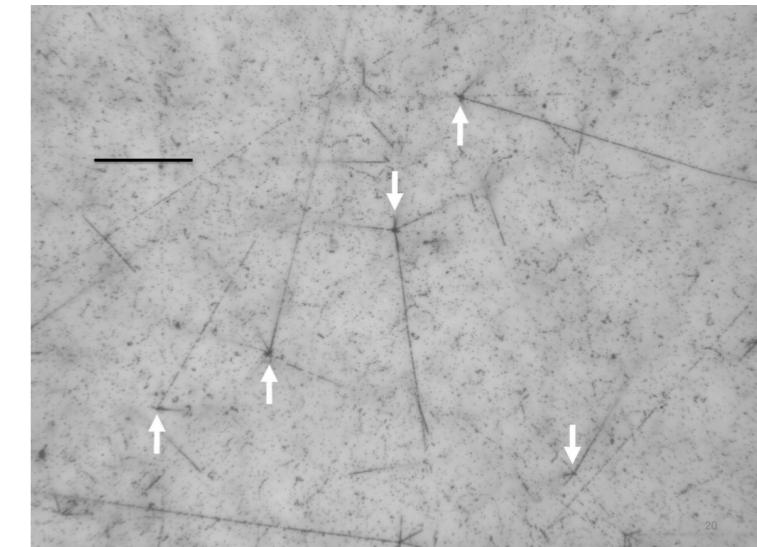
Silicon detectors(strip, pixel)



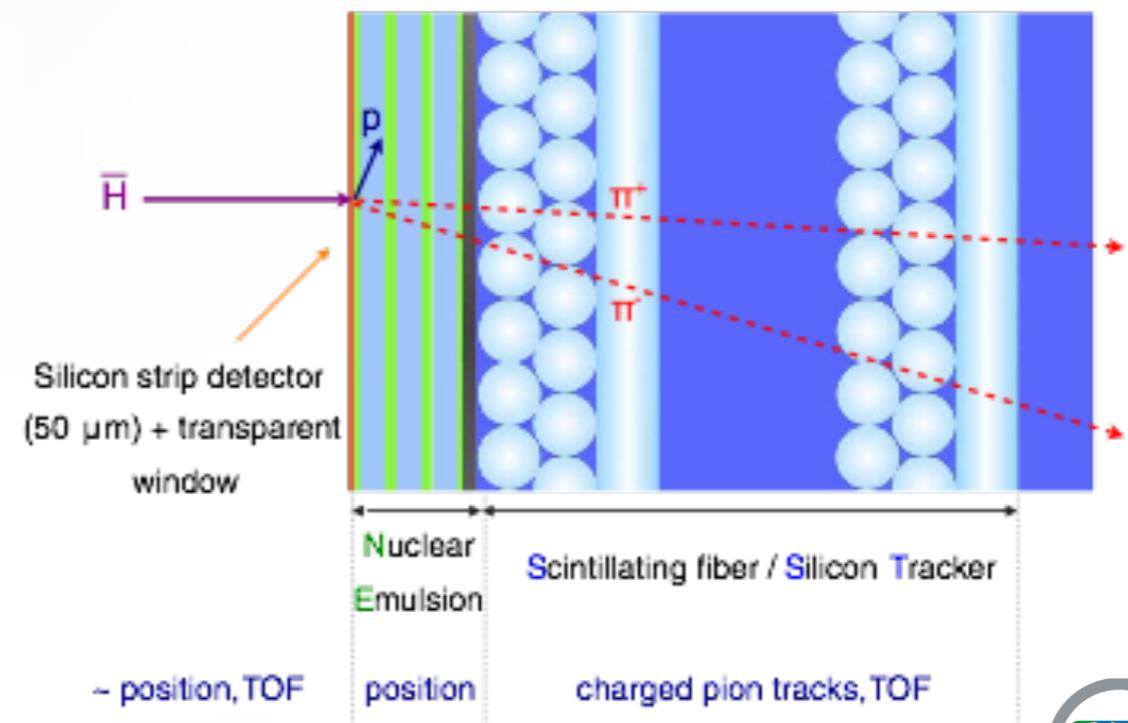
MCP



emulsions

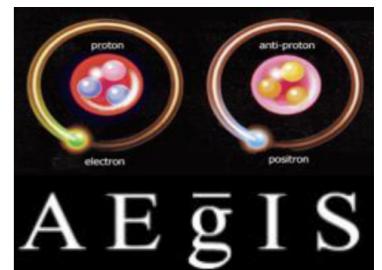


Hybrid detector needed
• timing
• position resolution $\sim 10 \mu m$

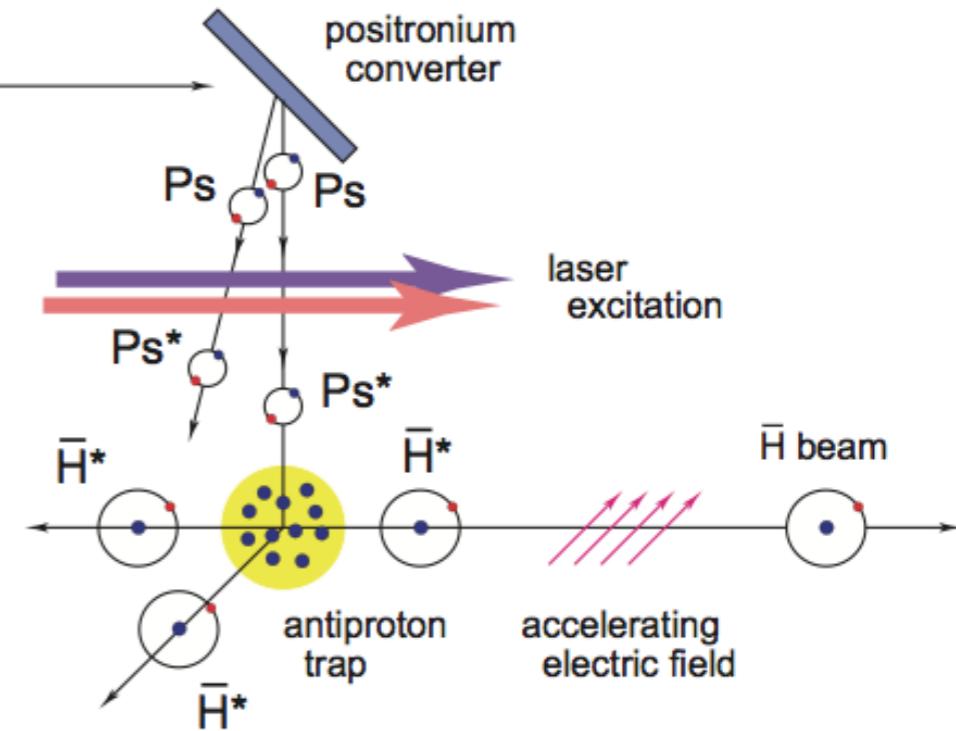
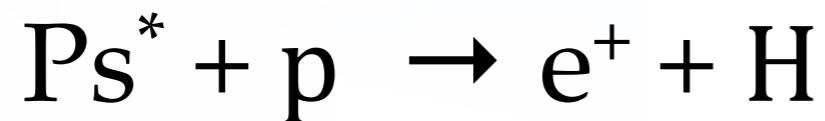


J. Storey et al. Hyp. Int. 0304-3843 (2014)

Production of H atoms



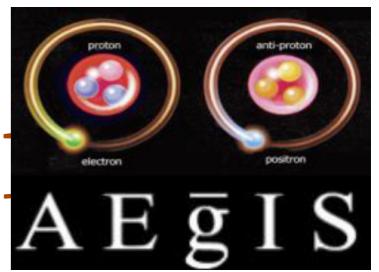
- Charge exchange with Ps^*
- Capture p from CERN-AD
- Cool p
- Ps production by e^+ on SiO_2
- Ps laser excitation to Rydberg state
- Interaction Ps^* with p cloud
- Recombination: resonant charge exchange



enhanced formation rate:

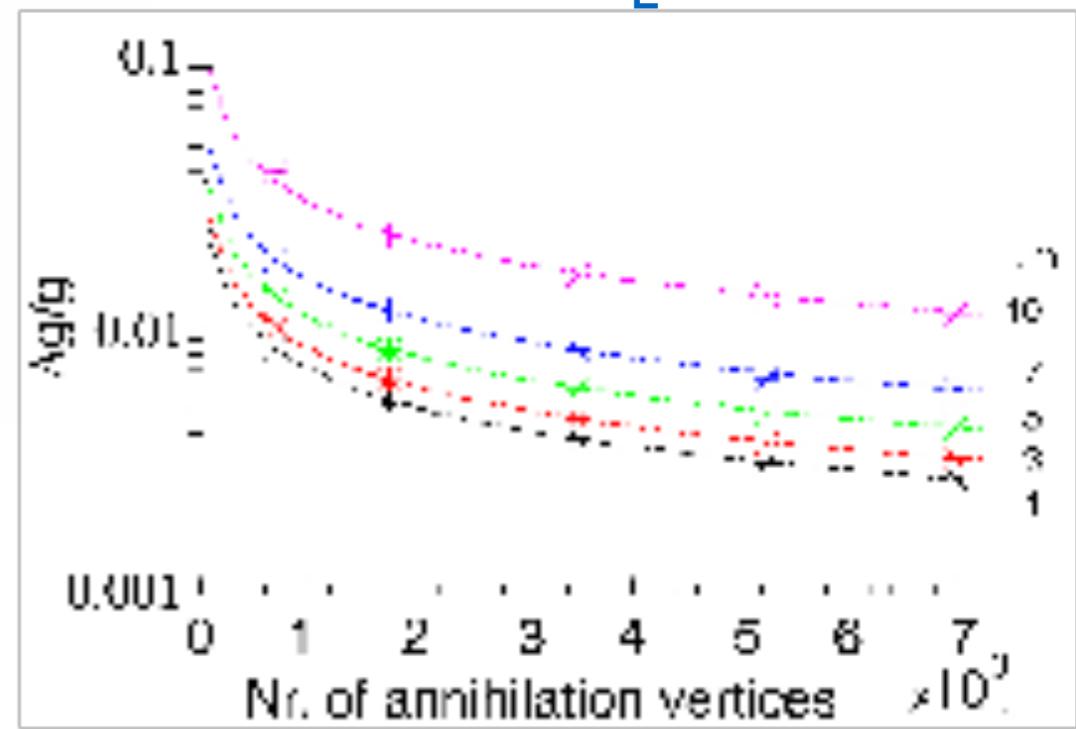
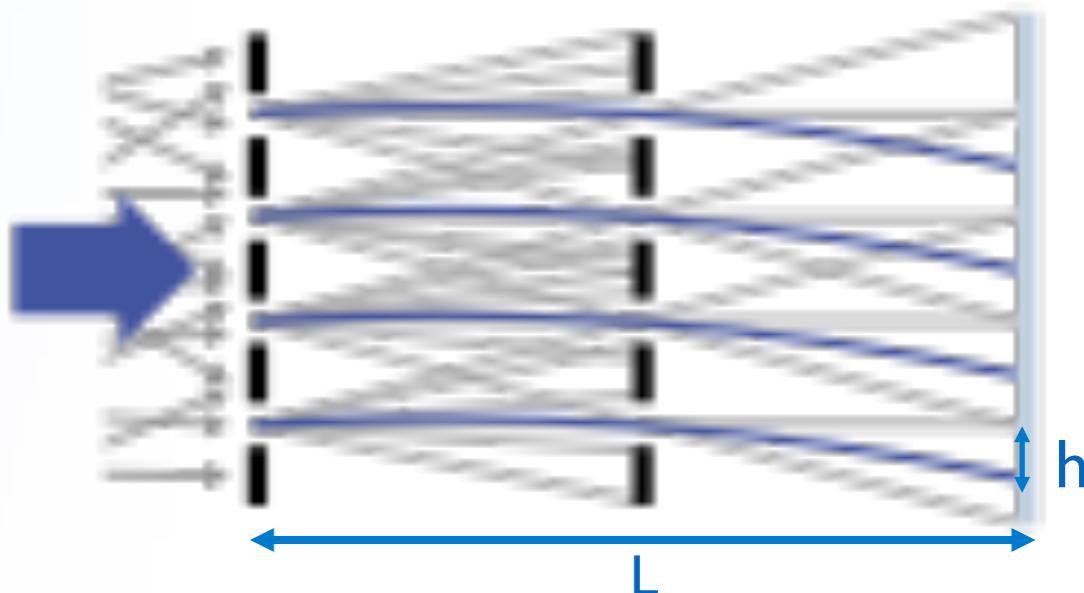
- $\sigma \sim n^4$
- $\text{Ps}(n_1) \rightarrow \text{H } (n_2)$

H beam & measurement principle



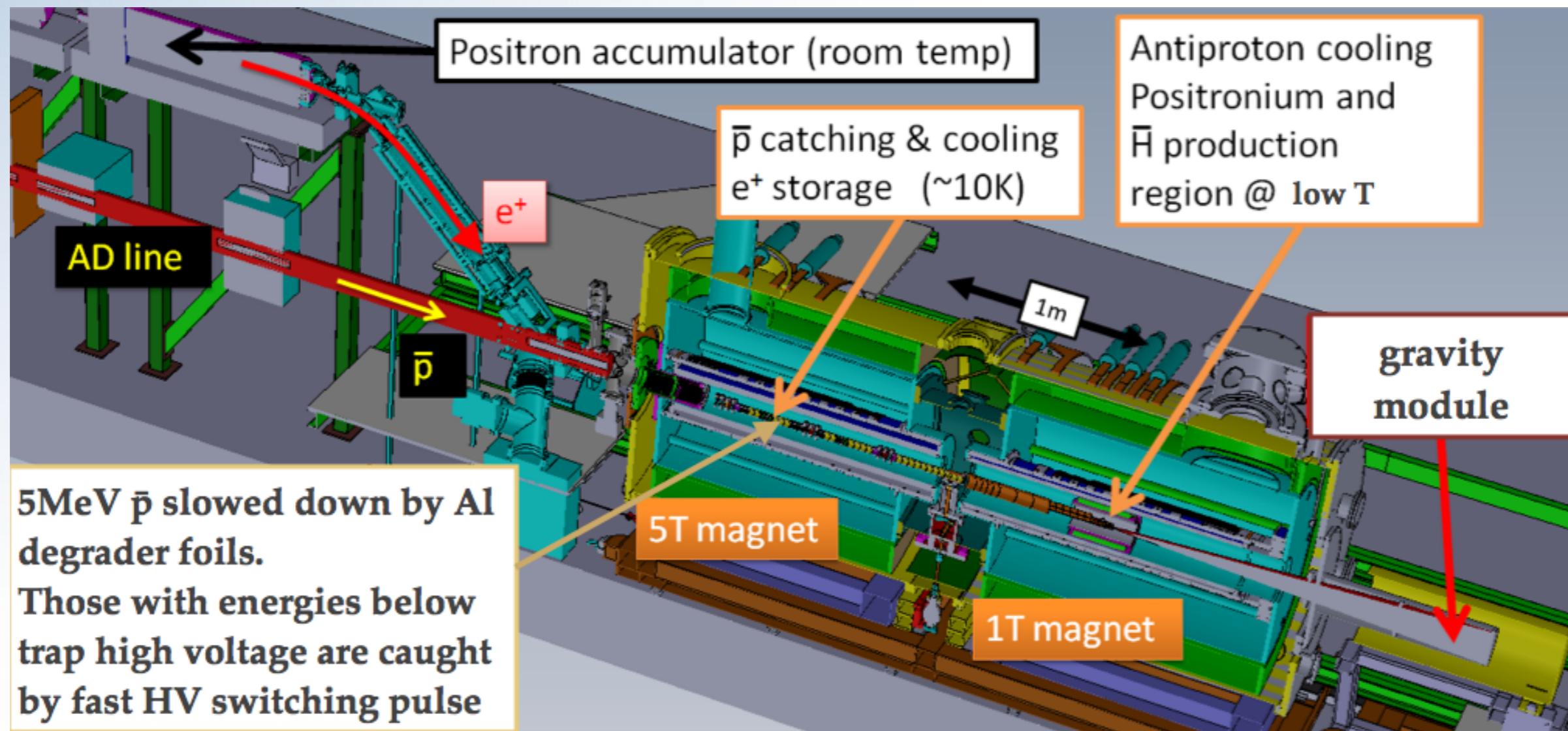
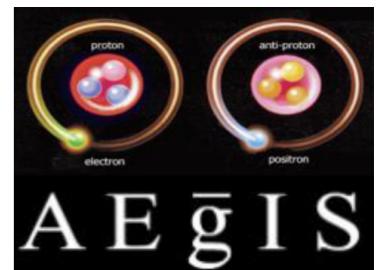
- H beam formation by Stark acceleration
 - $v_h \sim 400 \text{ m/s}$
- Measurement of g
 - Two-grating Moiré deflectometer $\Phi 100\text{mm}$, $L=1\text{m}$, slit $12 \mu\text{m}$, pitch $40 \mu\text{m}$, $100 \mu\text{m}$ thick
 - Position-sensitive annihilation detector $\Delta h \leq 10 \mu\text{m}$ (emulsion)
 - velocity information needed: pulsed beam & detector with time resolution ($\text{TOF} \sim 2.5 \text{ ms}$)

M.K. Oberthaler et al., Phys Rev. A 54, 3165 (1996)



Aghion, S. et al., JINST, 8, P08013 (2013)

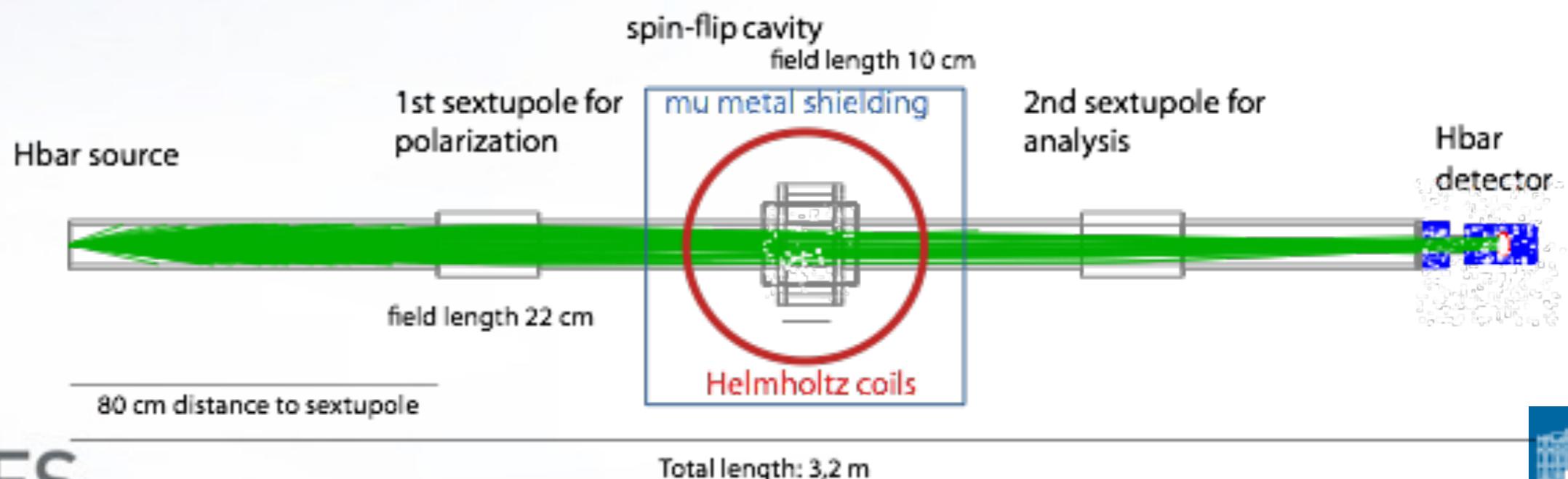
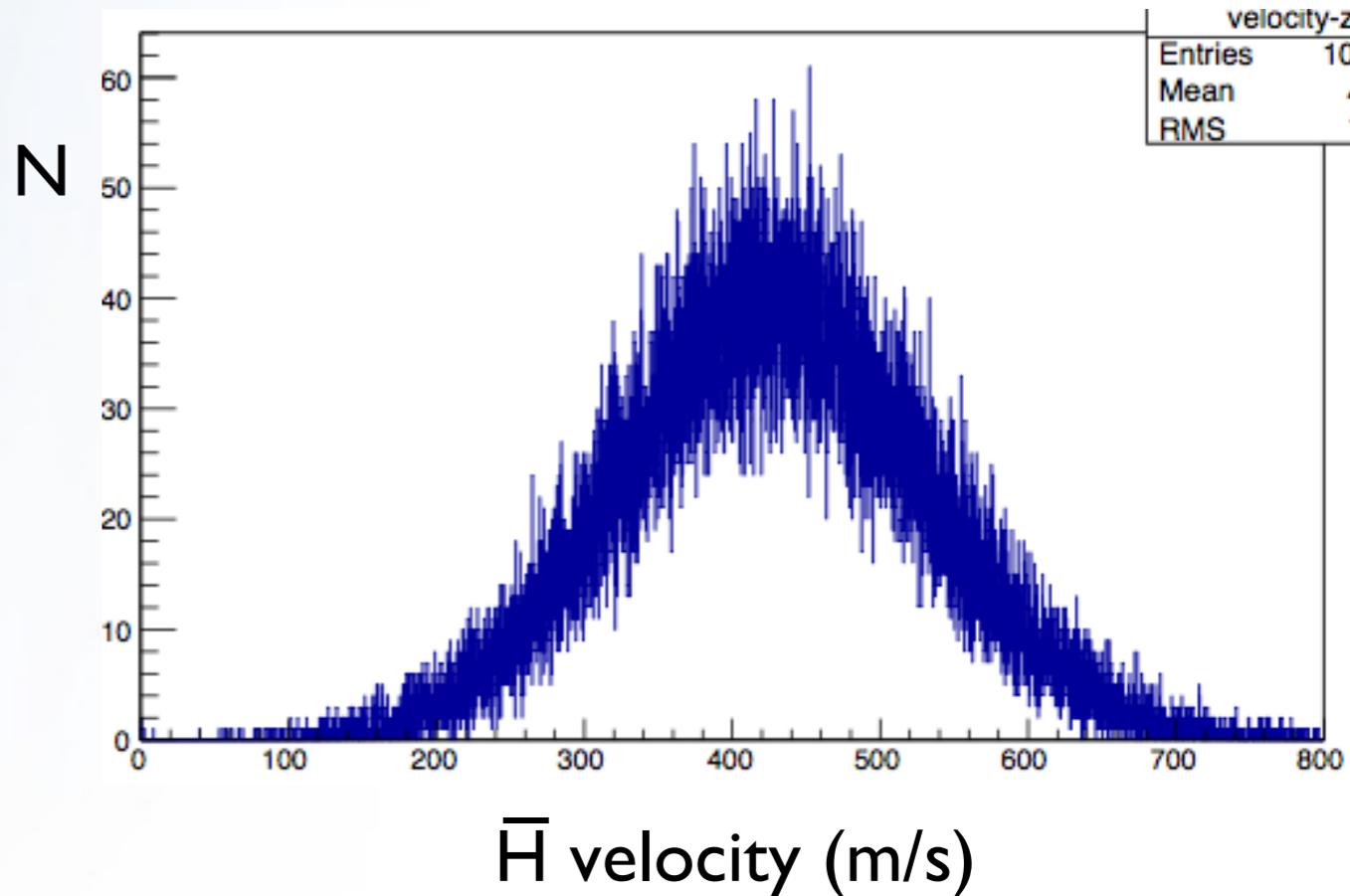
Experimental setup



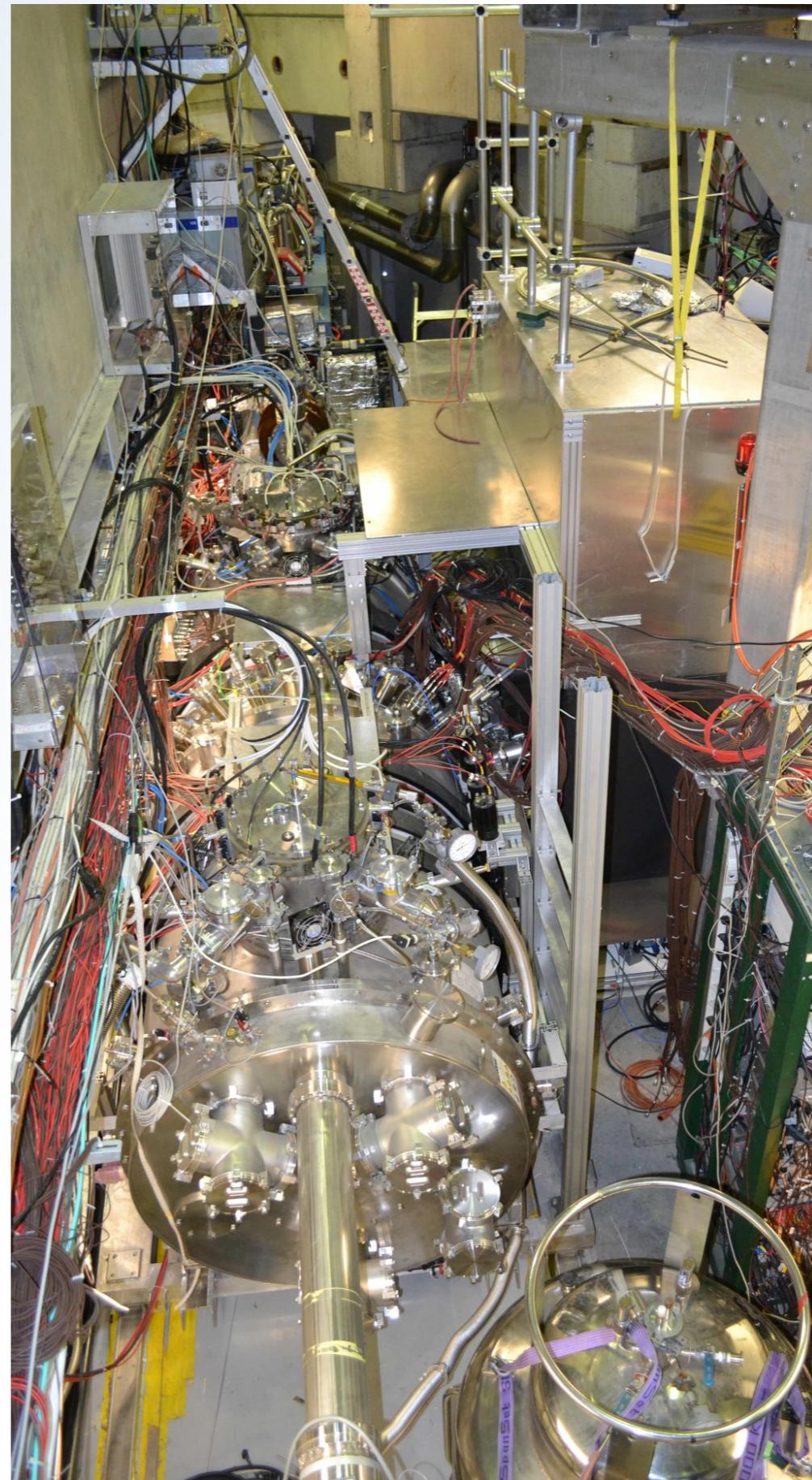
ULTRA-SLOW \bar{H} BEAM



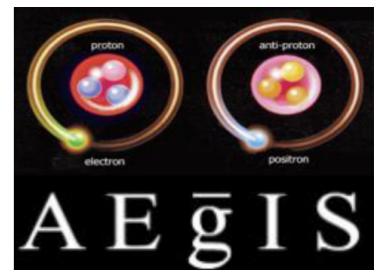
- narrow velocity distribution around 400 m/s
- no polarization: 2nd sextupole needed
- rates similar to CUSP



Experimental area

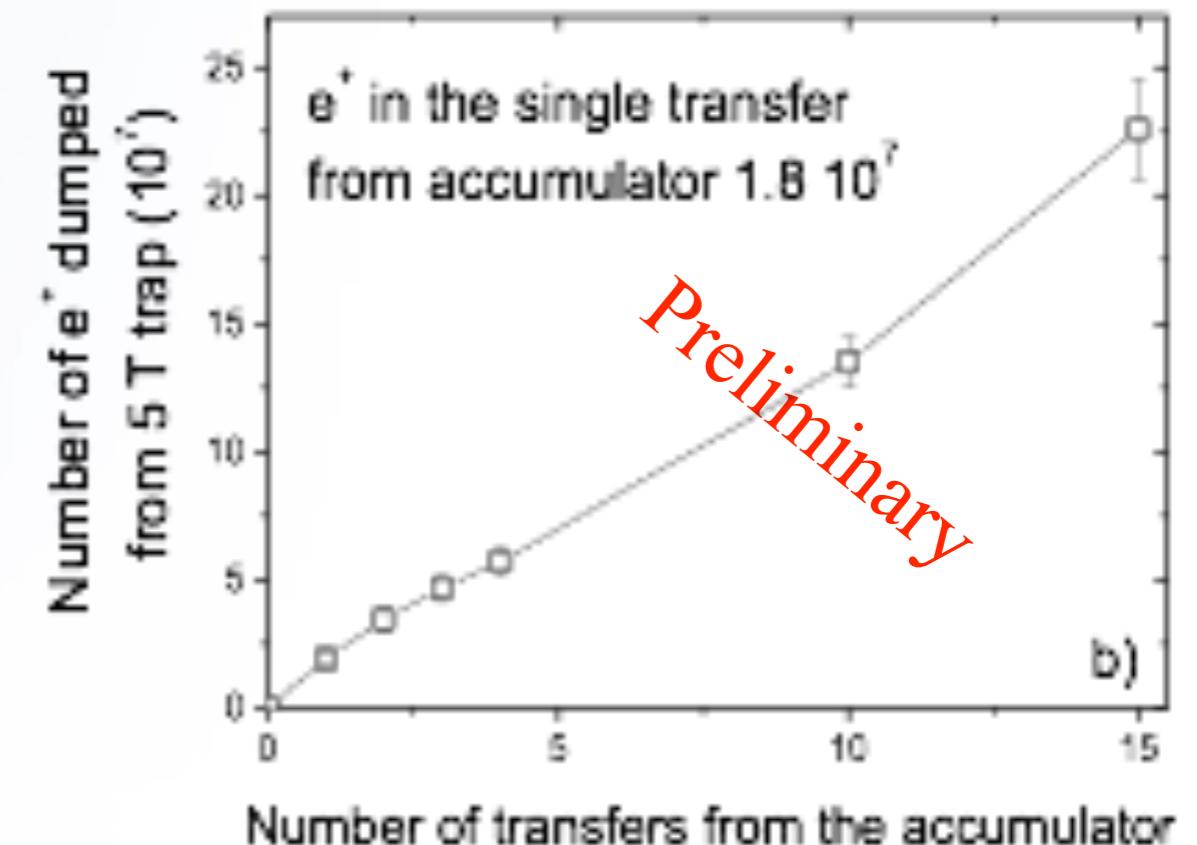
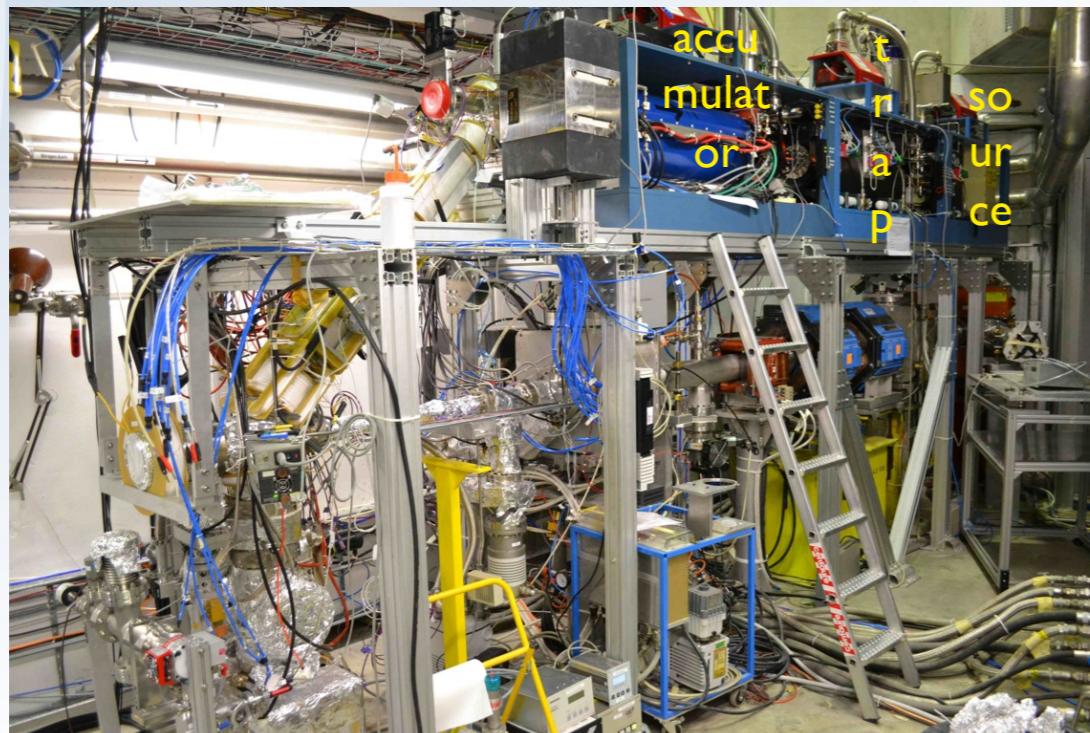


Positron system



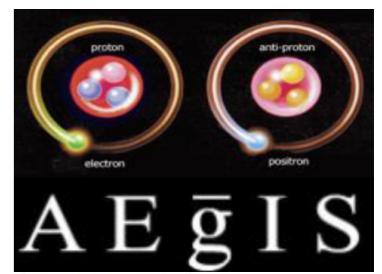
- Surko-type e^+ accumulator

Achieved e^+ numbers in 5

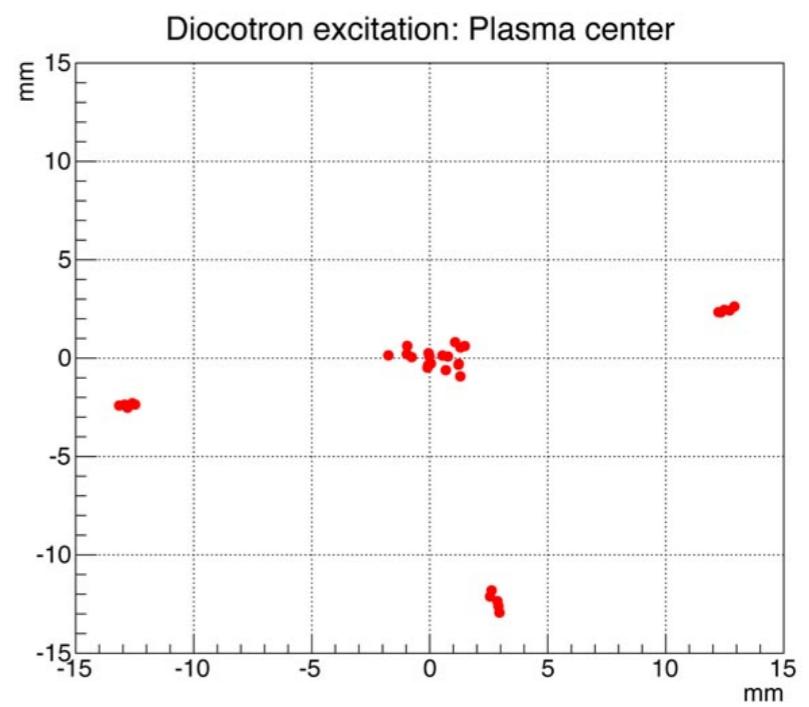
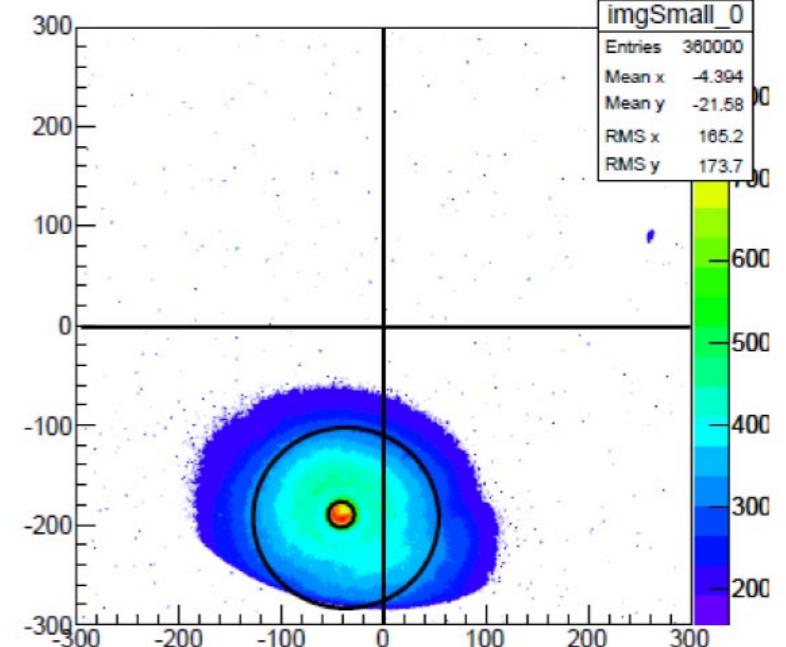
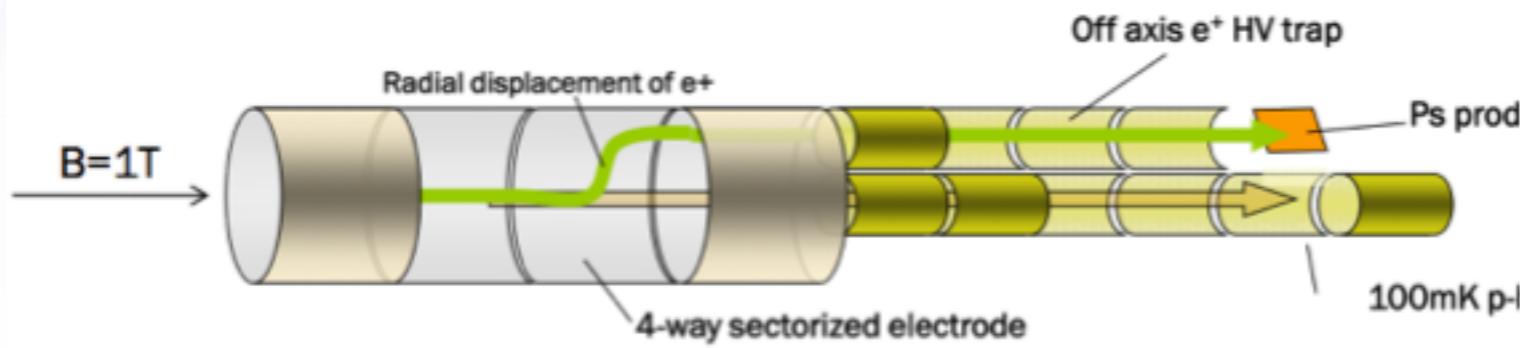


e^+ source strength ~ 10 mCi
total accumulation time ~ 80 min

e⁺ manipulation: tests with e⁻

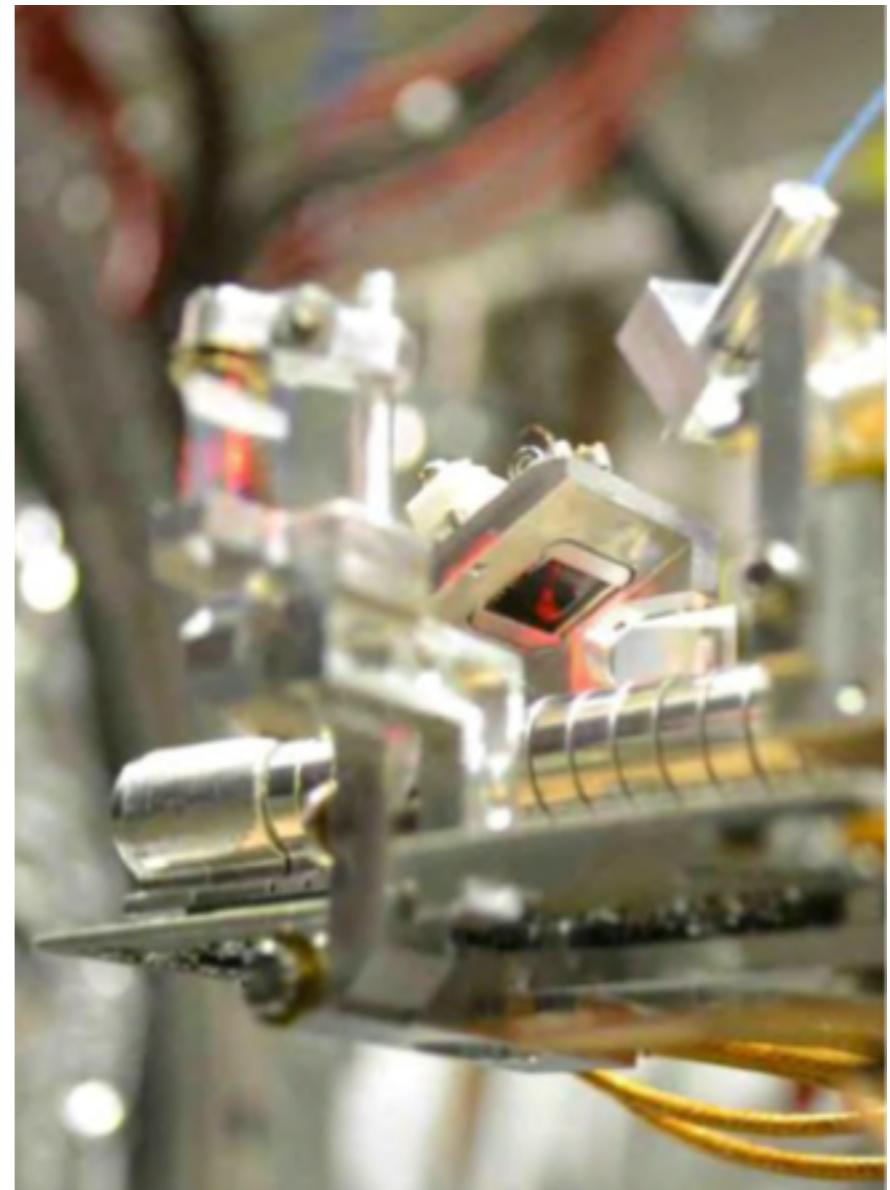
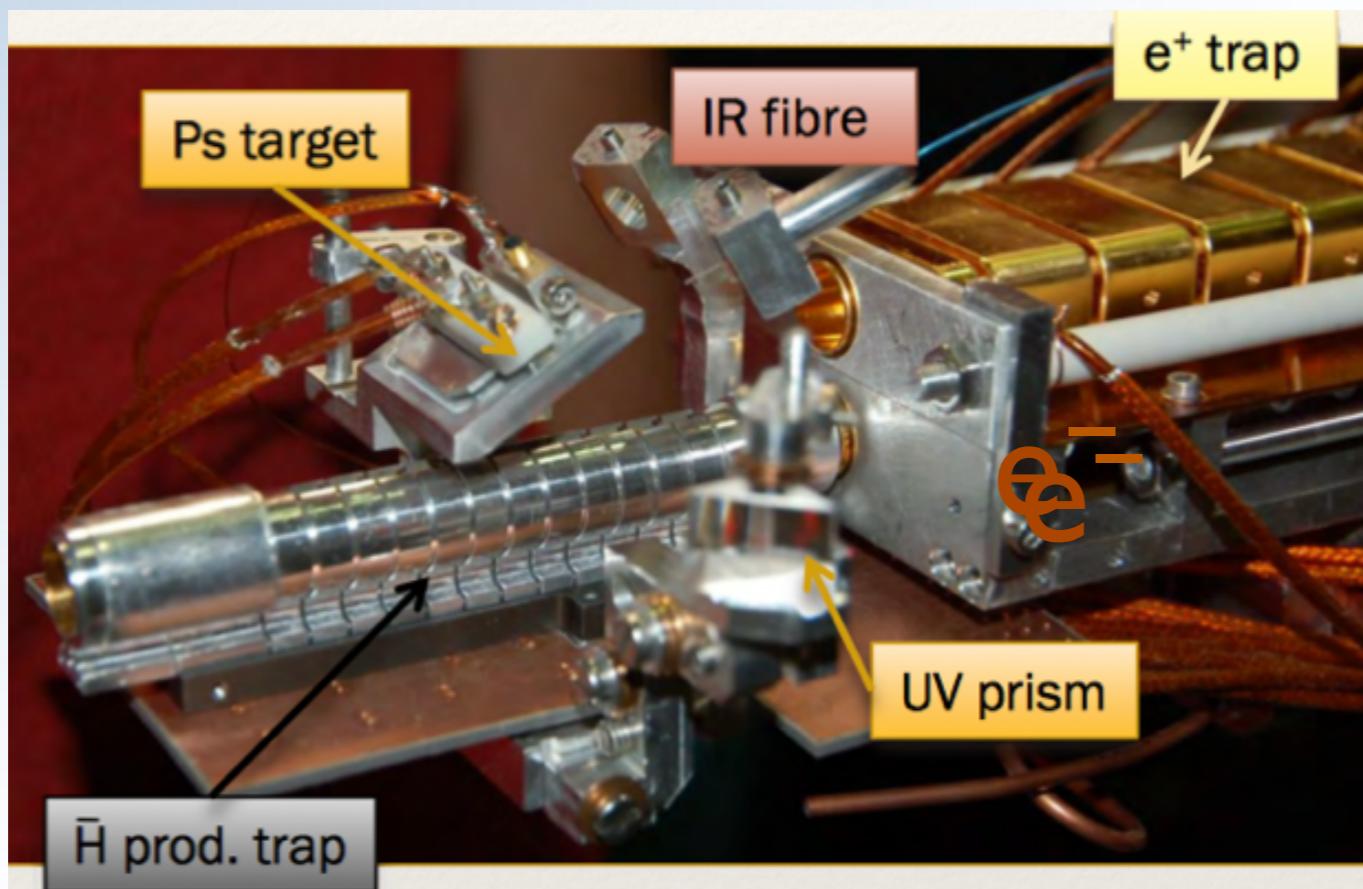
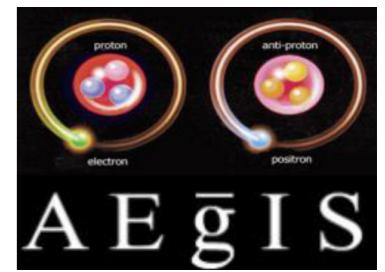


- Rotating wall compression
- Move e⁺ off-axis by exciting diocotron motion



C. Canali et al., Eur. Phys. J. D 65 (3) 499-504 (2011)

Ps and H formation region



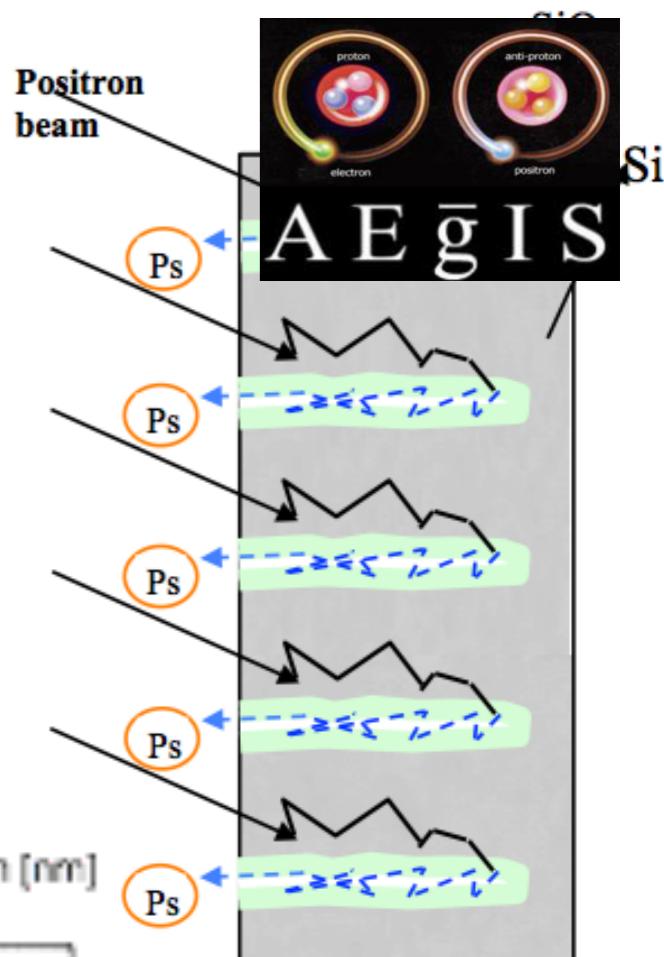
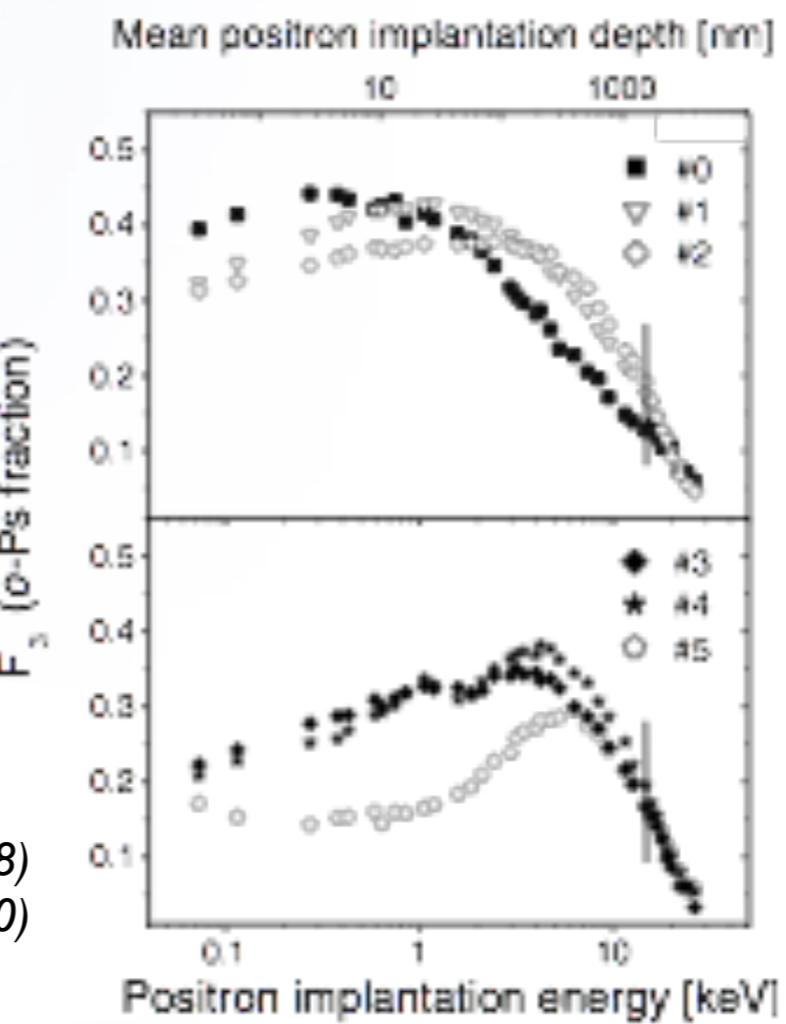
G. Consolati et al., Chem. Soc. Rev. 42, 3821 (2013)

Ps production

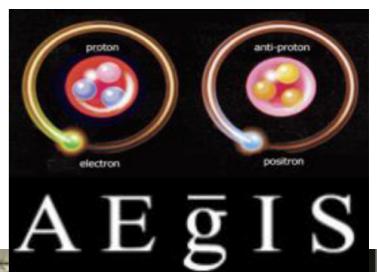
- Implantation of e^+ in nano-porous (8–14 nm Φ) silica
- ~75 K o-Ps needed
- Large implantation depth
- High Ps formation rate
- Possibility to tune nano channel dimensions: tune Ps temperature

S. Mariazzi et al., Phys. Rev. B 78, 085428 (2008)

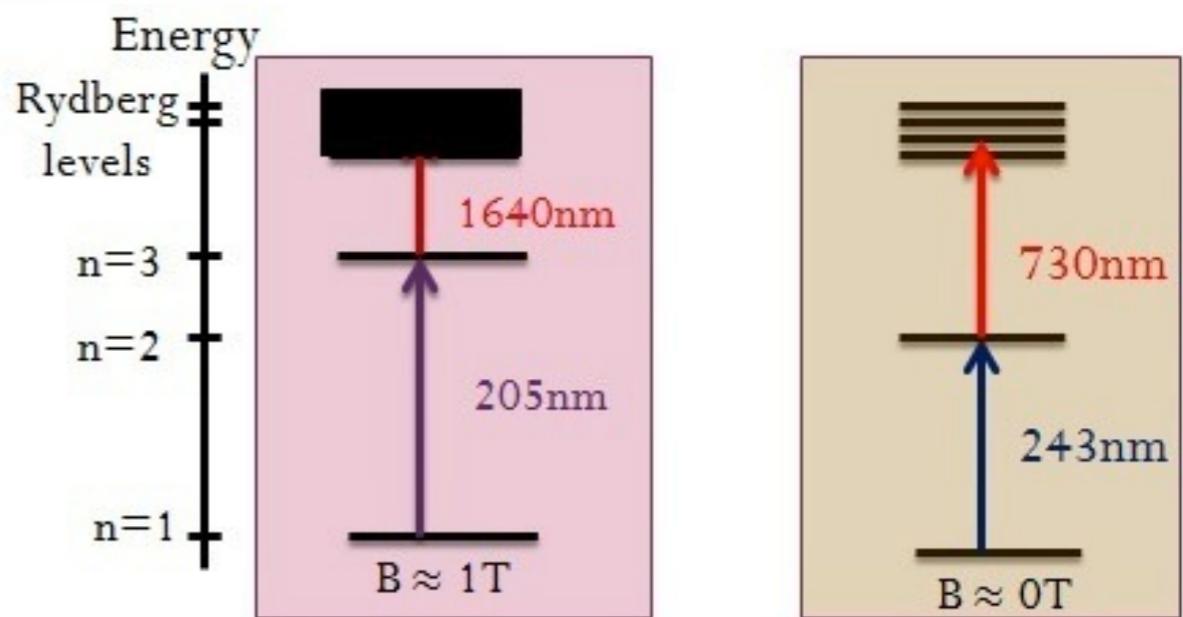
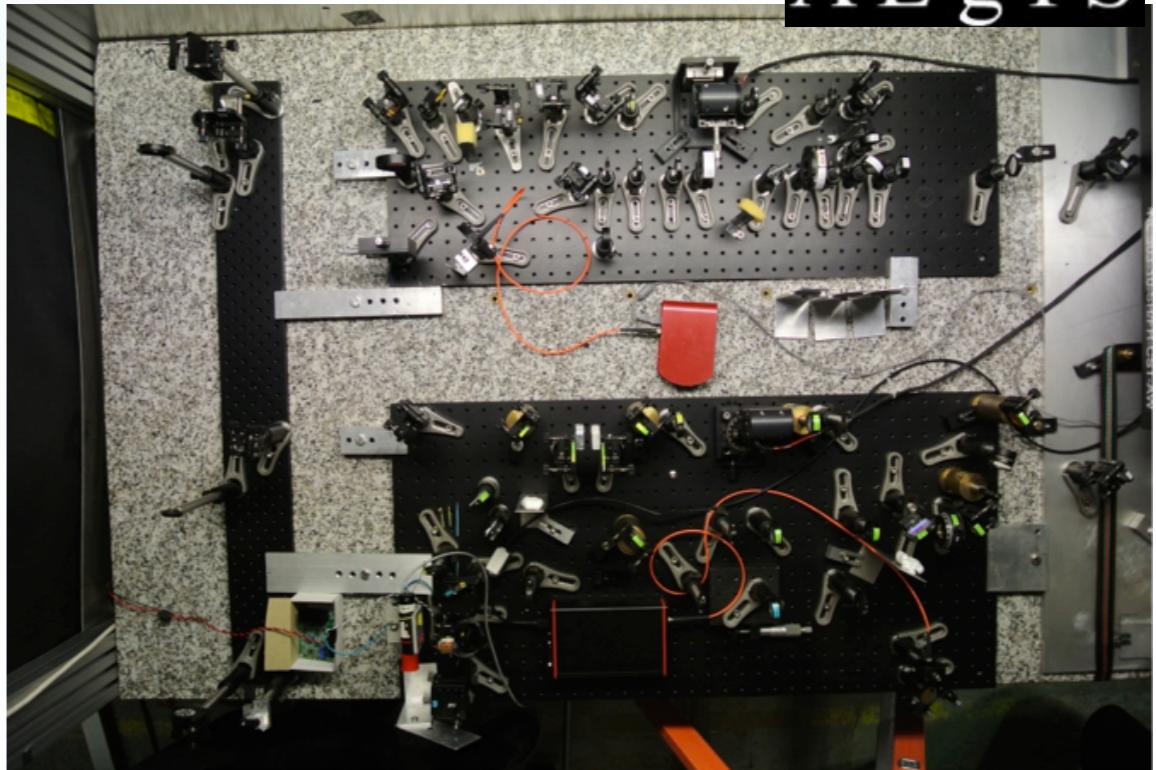
S. Mariazzi et al., Phys. Rev. B 81, 235418 (2010)



Ps excitation

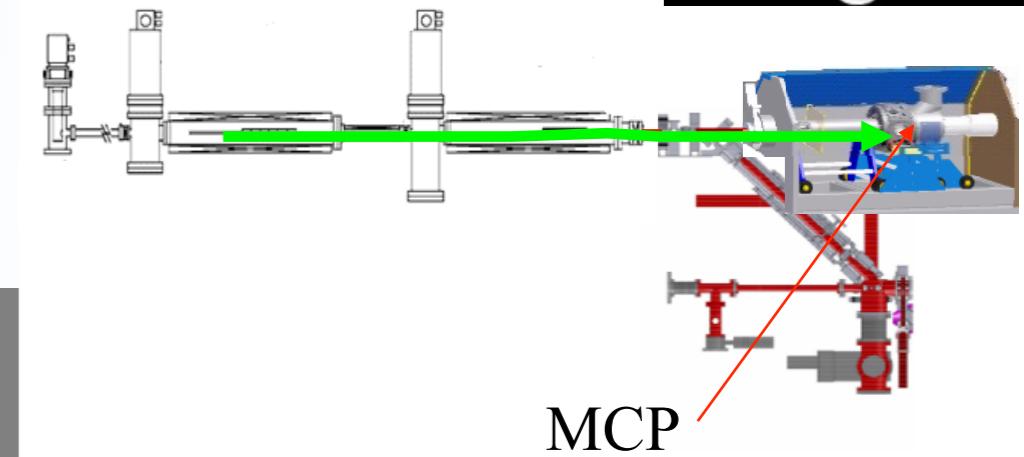
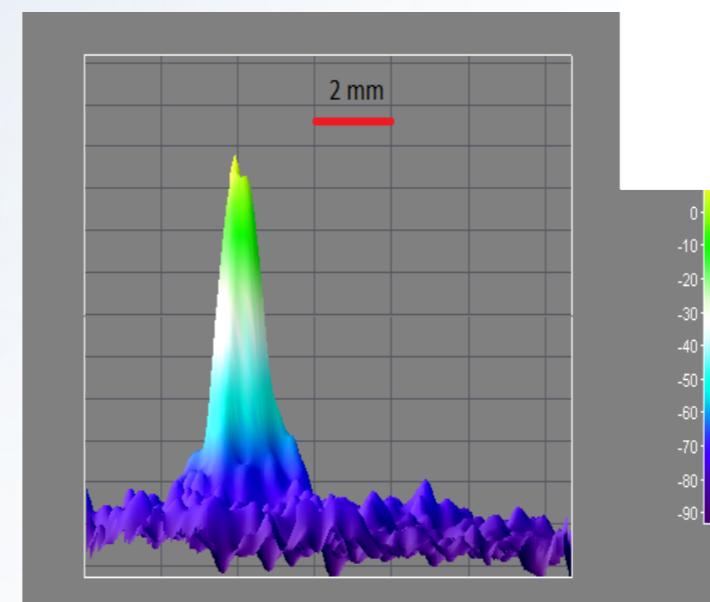
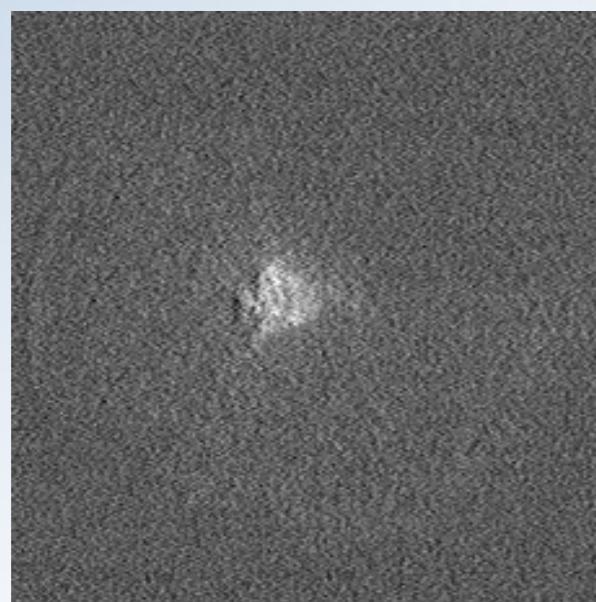
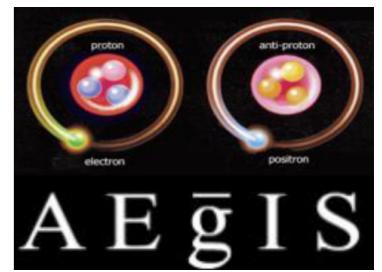


- 2-stage excitation
- UV: $n=1 \rightarrow 3$
- IR: $n=3 \rightarrow$ Rydberg ($n=26$)
- Alternative: $n=1 \rightarrow 2 \rightarrow$ Ry

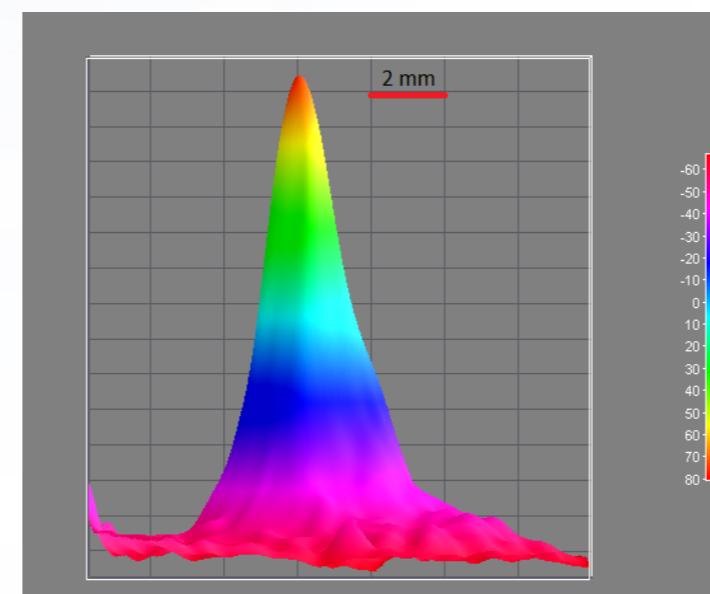
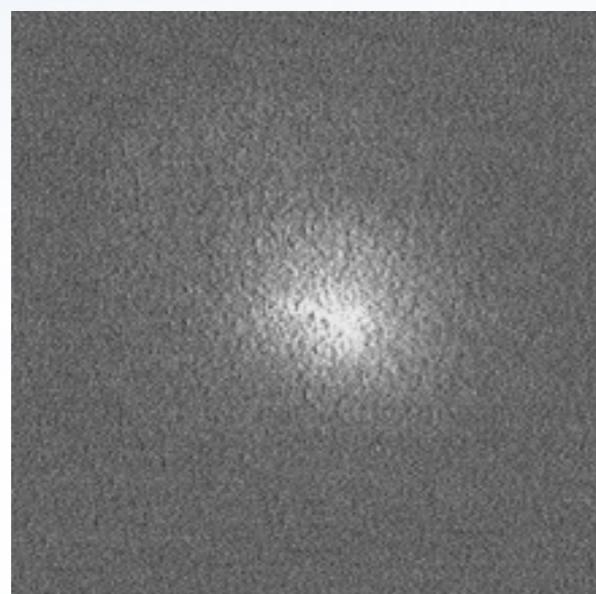


F. Castelli et al., Phys. Rev. B 78, 052512 (2008)
U. Warring et al., Phys. Rev. Lett. 102, 043001 (2009)

Chamber for Ps experiments: focus

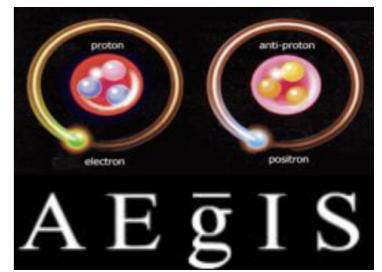


No field in the target region.
FWTM <4 mm



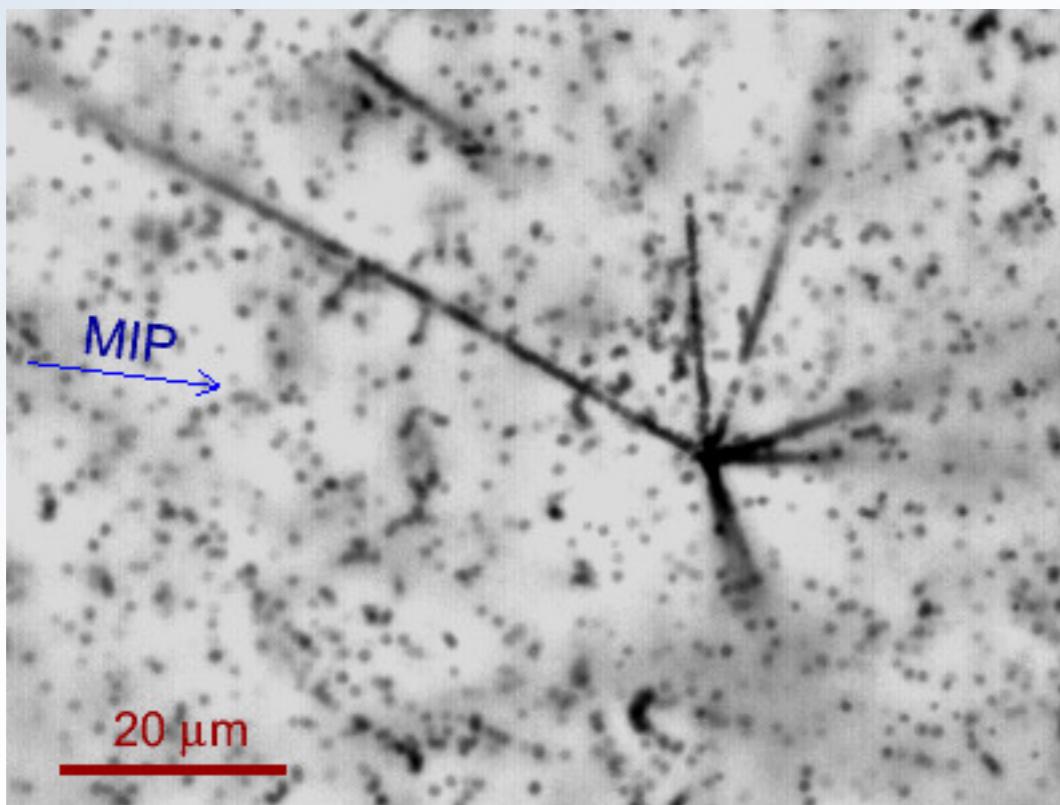
250 Gauss in the target region.
FWTM 5 mm

Gravity measurement - H^- detectors

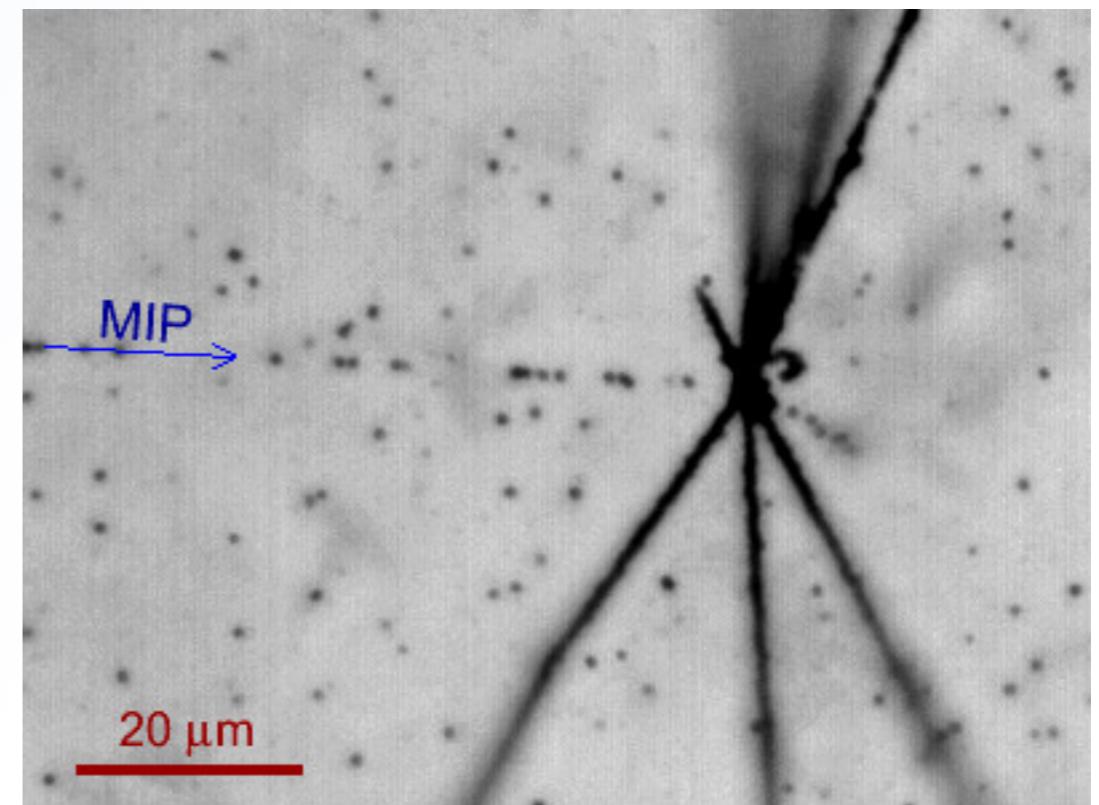


- Emulsion detectors at low temperatures and in vacuum

current emulsion

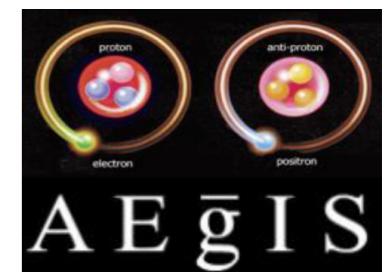


test emulsion

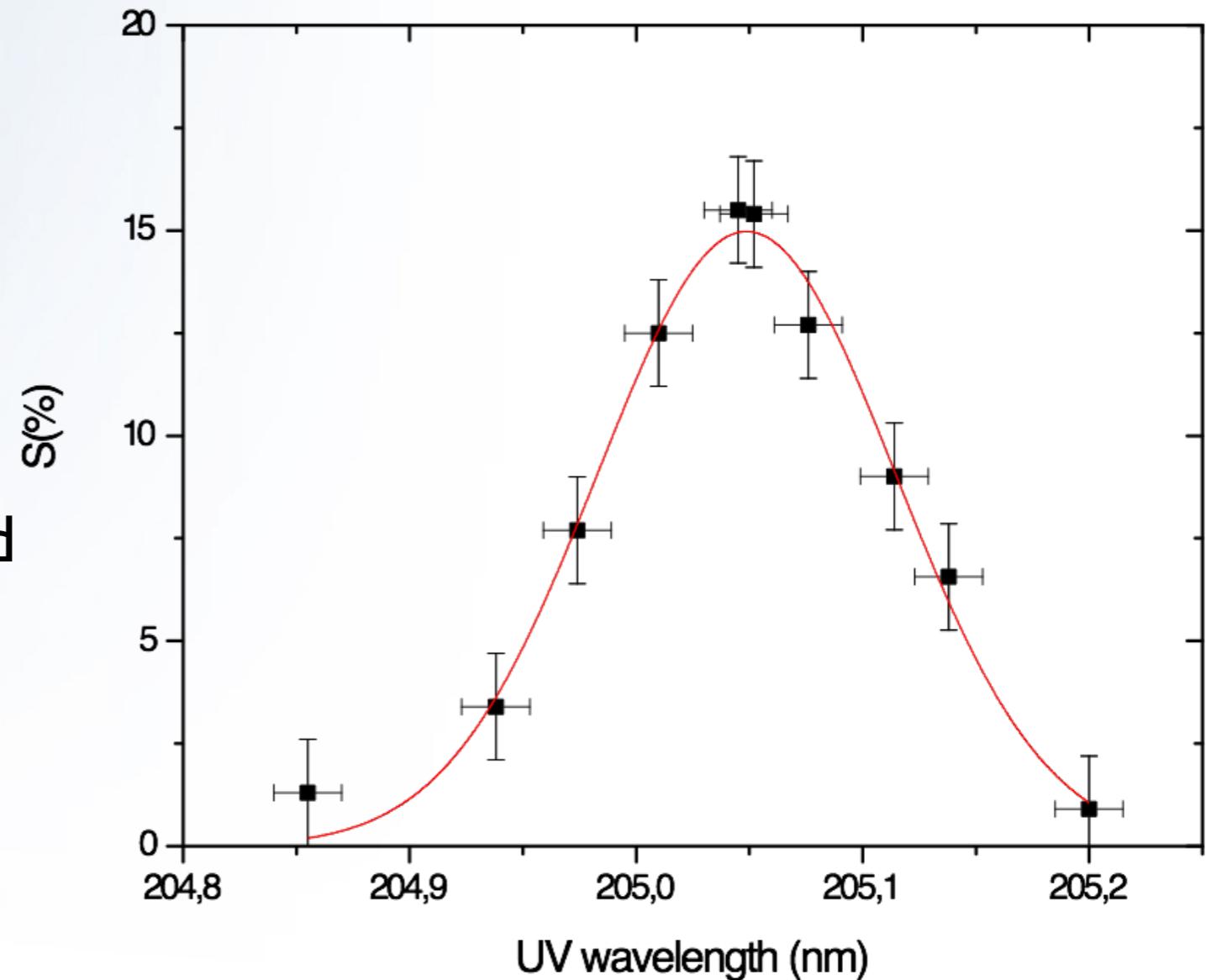


C. Amsler et al. JINST 8, P02015 (2013), S. Aghion et al., JINST 8, P08013 (2013)

Ps n=3 laser excitation



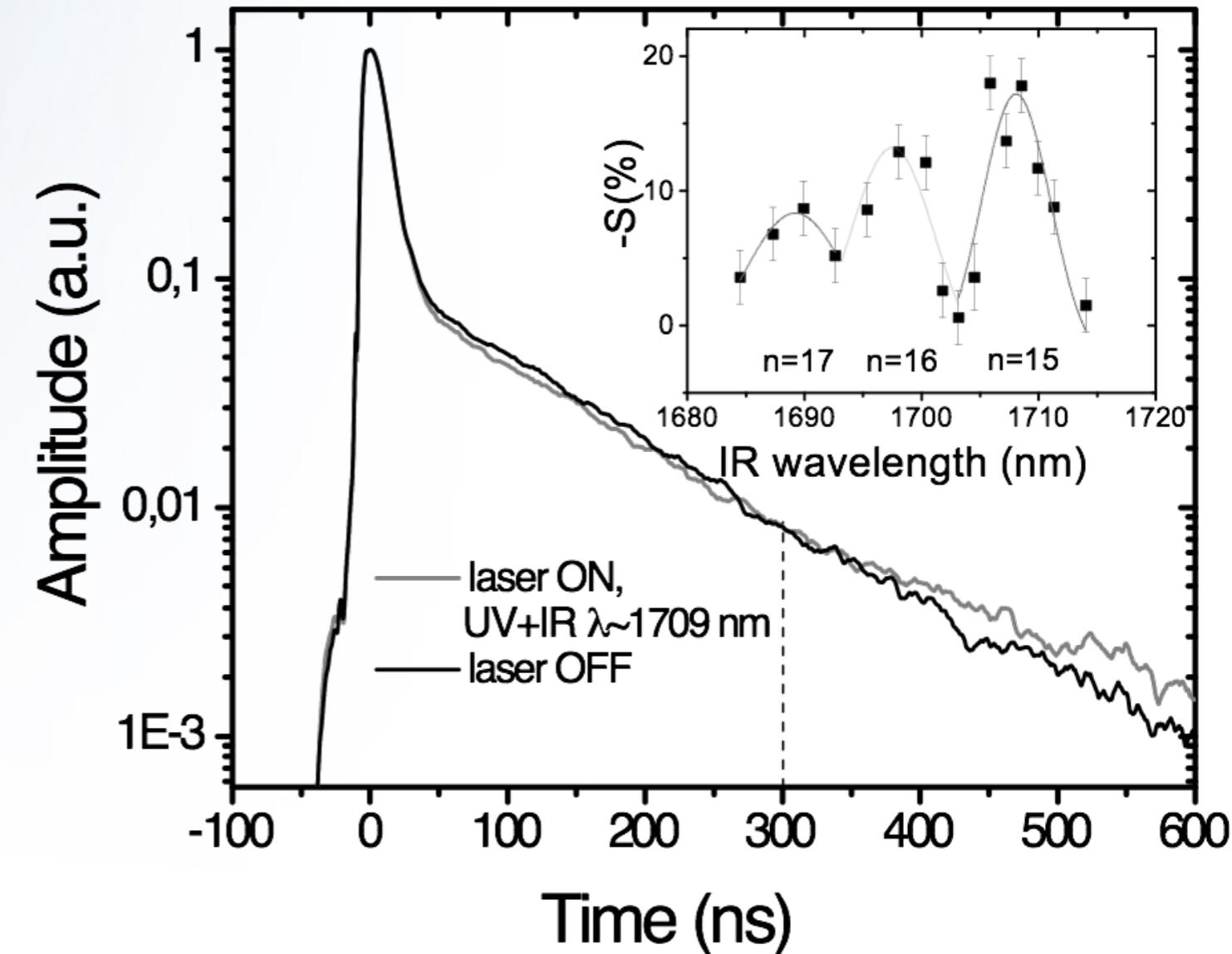
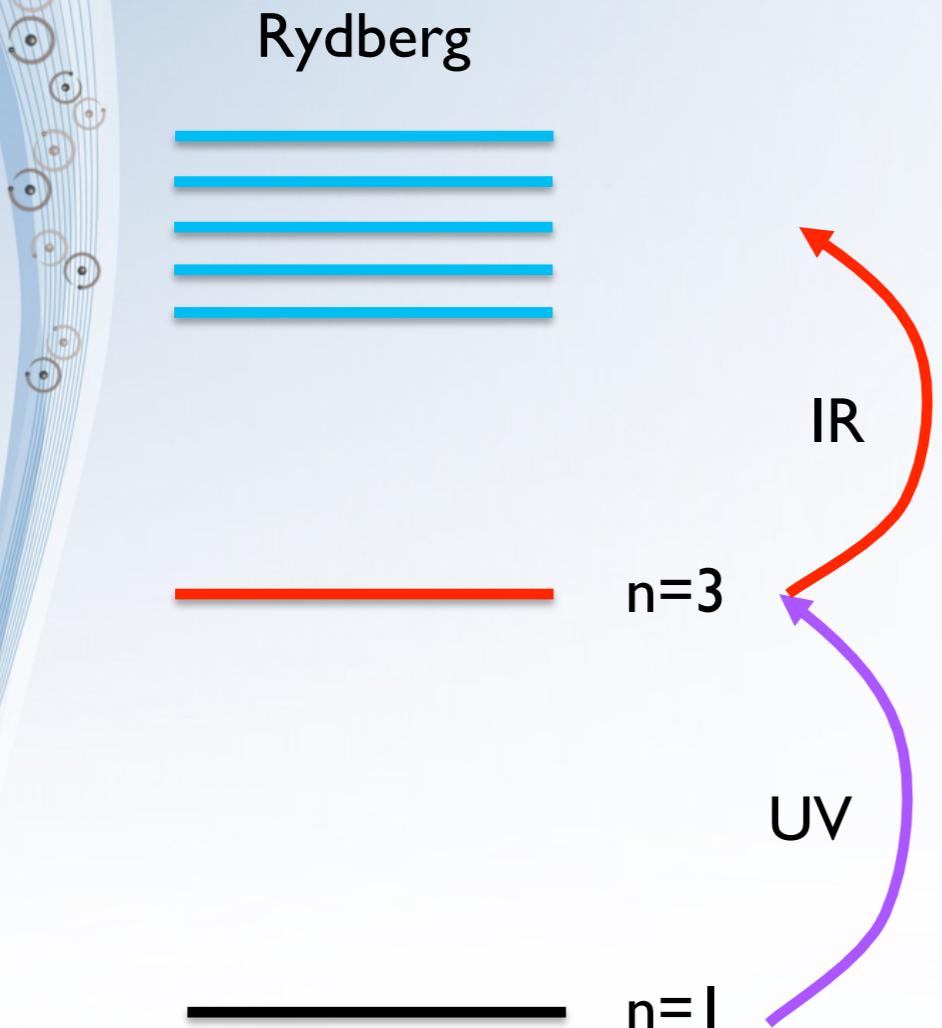
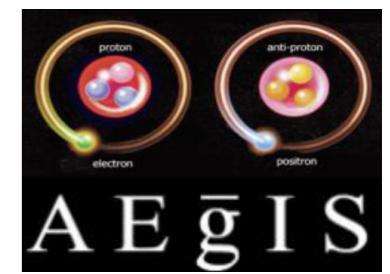
- 3P excitation line centered at 205.05 ± 0.02 nm
- excitation-ionization efficiency $\sim 15\%$



$$S(\%) = (\text{Area laser OFF} - \text{Area laser ON}) / \text{Area laser OFF}$$

Aghion S et al., submitted for publication

Ps Rydberg excitation



$$S(\%) = (\text{Area laser OFF} - \text{Area laser ON}) / \text{Area laser OFF}$$

AEGIS principle

