Key reaction rates of s-process nucleosynthesis and the impact of nuclear physics uncertainty



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**Facilities** Council



# The s-process

- produces a half amount of heavy isotopes beyond iron (a recent review, Kaeppeler et al., 2011, RvMP, 83, 157)
- (n,  $\gamma$  ) and  $\beta$  -decay from Fe seeds along the stable nuclei



## Two sites of the s-process

### weak s-process

- massive stars (  $> 10 \text{ M}_{\odot}$ )
- core He and shell C-burnings
- neutron source: <sup>22</sup>Ne(a,n)<sup>25</sup>Mg
- weak component (A < 100)

### main s-process

- low mass AGB stars
- thermal pulses
- neutron source: <sup>13</sup>C(a,n)<sup>16</sup>O
- main component (up to Pb, Bi)







## The s-process in massive stars: uncertainties



### Stellar environment

### Stellar (structure) evolution models

 mass, metallicity, dynamics: convection, rotation and magnetic fields, single/binary stars etc.

### Nuclear burnings in the stellar interior

- main fusion reactions: triple- $\alpha$ , <sup>12</sup>C(a,g)<sup>16</sup>O, …
- n-source and n-poison reactions: <sup>22</sup>Ne(a,n)<sup>25</sup>Mg, …

(see e.g., Nishimura et al., AIPC 1594 p 146, 2014)

Network calculation

# Monte-Carlo simulation

- <u>Nucleosynthesis</u>
  - (n,g) reaction
  - beta-decay

uncertainty in the final abundances

feedback (find key reaction/decay)

## Contents

### Introduction

- <u>Methods</u>
  - the PizBuin MC nucleosynthesis code
  - T-dependent reaction rate uncertainty
  - MC and analysis
- <u>Results</u>
  - s-process in massive stars
  - main s-process in AGB stars
  - other processes
- <u>Summary</u>
- Rauscher, NN+ (2016), MNRAS in press; arXiv1606.05671
- NN+ 2016, MNRAS, in prep.
- Cescutti, NN+ 2017(?), in prep.



Nuclear chart : <u>http://www.nndc.bnl.gov/nudat2/</u>

# Monte-Carlo network code

- Monte-Carlo framework
  - PizBuin MC-driver
    - (developed by Rauscher, NN, Hirschi)
  - a simple "Brute-force" approach
  - parallelized by OpenMP for shared memory architectures
     (paralleled easily, but harder debugging. . .)



Piz Buin (mountain)

### Nuclear Reaction network

- Network solver:
  - WinNet: the latest Basel network, Winteler+, 2012
- Reaction rates:
  - Reaclib: (Rauscher & Thielemann 2000)
  - T-dependent beta-decay (Takahashi & Yokoi 1987, Goriely 1999)

### T-dependent uncertainty:

- Provided by Reaclib format, based on Rauscher 2012

## <u>T-dependent uncertainty: (n,γ) & β-decay</u>

### (n,g) reactions

- <u>Experimental rates</u>
  - base rates: KADoNiS v0.3 (Dillmann+, 2009)
  - consider both g.s. and e.s. (Rauscher, ApJ, 775, 2011  $U(T) = U_{\rm g.s.}X + U_{\rm e.s.}(1-X)$

 $\beta$ -decay: a similar method using partition functions

(See, NN+2016) <sup>83</sup>Kr(n,ღ)<sup>84</sup>Kr



## MC calculation with reaction network



## Propagation of uncertainty: <sup>64</sup>Cu(β-)<sup>64</sup>Zn



### Uncertainty by MC: (n,g) and β-decay

(n,g) &  $\beta^{\pm}$ 



### Uncertainty by MC: (n,g) and β-decay

only (n,g) only  $\beta^{\pm}$ 



### **Correlation factors**



### **Correlation factors**



We can derive key reaction rates from the MC results

## Screening of the reaction rates



# Key rate Level 1 $^{64}Cu(+\beta)^{64}Zn$ $^{67}Zn + n \leftrightarrow \gamma + {}^{68}Zn$ $^{72}Ge + n \leftrightarrow \gamma + {}^{73}Ge$ $^{73}Ge + n \leftrightarrow \gamma + {}^{74}Ge$ $^{77}Se + n \leftrightarrow \gamma + {}^{78}Se$ $^{78}Se + n \leftrightarrow \gamma + {}^{79}Se$ $^{80}Br(+\beta)^{80}Kr$ $^{81}Kr + n \leftrightarrow \gamma + {}^{82}Kr$ $^{83}Kr + n \leftrightarrow \gamma + {}^{84}Kr$ $^{85}Kr + n \leftrightarrow \gamma + {}^{86}Kr$

## Lv2





Lv3



I v/4

## Key reaction list

all Lv1 reaction all Lv1+Lv2 reaction

are fixed

are fixed

							1	
Nuclide	r <sub>cor,0</sub>	r <sub>cor,1</sub>	r <sub>cor,2</sub>	Key rate Level 1	Key rate Level 2	Key rate level 3	$X_0$ (30 keV)	$X_0$ (80 keV)
<sup>64</sup> Zn	0.76			$^{64}$ Cu(+ $\beta$ ) <sup>64</sup> Zn				
	-0.47	-0.73			$^{64}Cu + e^- \leftrightarrow \nu_e + {}^{64}Ni$			
<sup>67</sup> Zn	-0.67			$^{67}$ Zn + n $\leftrightarrow \gamma$ + $^{68}$ Zn			1.00	1.00
<sup>72</sup> Ge	-0.85			$^{72}\text{Ge} + n \leftrightarrow \gamma + ^{73}\text{Ge}$			1.00	1.00
<sup>73</sup> Ge	-0.84			$^{73}\text{Ge} + n \leftrightarrow \gamma + {}^{74}\text{Ge}$			0.88	0.81
<sup>74</sup> Ge	-0.44	-0.53	-0.67	·		$^{74}\text{Ge} + n \leftrightarrow \gamma + ^{75}\text{Ge}$	1.00	1.00
<sup>75</sup> As	-0.50	-0.58	-0.70			$^{75}As + n \leftrightarrow \gamma + ^{76}As$	1.00	1.00
<sup>77</sup> Se	-0.86			$^{77}\text{Se} + n \leftrightarrow \gamma + {}^{78}\text{Se}$			1.00	1.00
<sup>78</sup> Se	-0.71			<sup>78</sup> Se + n $\leftrightarrow \gamma$ + <sup>79</sup> Se			1.00	1.00
	0.37	0.68			$^{68}$ Zn + n $\leftrightarrow \gamma$ + $^{69}$ Zn		1.00	1.00
<sup>80</sup> Se	-0.76			$^{80}\mathrm{Br}(+m{eta})^{80}\mathrm{Kr}$	-			
	0.27	0.73			$^{80}{ m Br}(-m eta)^{80}{ m Se}$			
	0.16	0.44	0.88			$^{80}Br + e^- \leftrightarrow \nu_e + {}^{80}Se$		
<sup>79</sup> Br	-0.63	-0.73			$^{79}Br + n \leftrightarrow \gamma + {}^{80}Br$		1.00	1.00
<sup>81</sup> Br	-0.80			$^{81}$ Kr + n $\leftrightarrow \gamma$ + $^{82}$ Kr	-		1.00	1.00
<sup>83</sup> Kr	-0.76			$^{83}$ Kr + n $\leftrightarrow \gamma$ + $^{84}$ Kr			0.81	0.74
<sup>84</sup> Kr	-0.49	-0.64	-0.76			$^{84}$ Kr + n $\leftrightarrow \gamma$ + $^{85}$ Kr	1.00	1.00
<sup>86</sup> Kr	0.84			$^{85}$ Kr + n $\leftrightarrow \gamma$ + $^{86}$ Kr			1.00	1.00
	-0.31	-0.71			$^{86}$ Kr + n $\leftrightarrow \gamma$ + $^{87}$ Kr		1.00	1.00
	-0.33	-0.62	-0.90			$^{85}\mathrm{Kr}(+m{eta})^{85}\mathrm{Rb}$		
<sup>87</sup> Rb	-0.57	-0.64	-0.95			$^{87}$ Rb + n $\leftrightarrow \gamma$ + $^{88}$ Rb	1.00	1.00
						ji		

## Key reaction list: <sup>86</sup>Kr



### vs <sup>85</sup>Kr(n,g)<sup>86</sup>Kr

	<sup>86</sup> Kr(n,g)	<sup>85</sup> Kr(b+)
upper	-0.42	-0.68
standard	-0.71	-0.62
lower	-0.84	-0.42

# main s-process

- 'one zone' a low mass AGB star
- 2Msun model by MESA code
- the initial <sup>13</sup>C is adjusted
- a typical s-process pattern



### lighter



medium

### Cescutti+NN+, in prep.

heavier

# main s-process

(n,g)-contribution  $\beta$ -decay contribution



# One more example: gamma-process

### Rauscher, NN+ (2016), MNRAS (in press) arXiv:1606.05671



15M<sub>sun</sub>

Nuclide	r <sub>corr,0</sub>	r <sub>corr,1</sub>	r <sub>corr,2</sub>	Key rate Level 1	Key rate Level 2	Key rate Level 3	$X_0 (2 \text{ GK})$ capture	X <sub>0</sub> (3 GK capture
<sup>78</sup> Kr	-0.84			$^{77}Br + p \leftrightarrow \gamma + {}^{78}Kr$			$9.63 \times 10^{-2}$	$4.44 \times 10^{-1}$
	0.34	0.87			$^{79}$ Kr + n $\leftrightarrow \gamma$ + $^{80}$ Kr		$1.28 \times 10^{-1}$	$7.94 \times 10^{-3}$
<sup>92</sup> Mo	-0.74			$^{91}Nb + p \leftrightarrow \gamma + ^{92}Mo$			$8.88 \times 10^{-1}$	$8.24 \times 10^{-1}$
<sup>96</sup> Ru	-0.73			$^{92}Mo + \alpha \leftrightarrow \gamma + {}^{96}Ru$			1.00	$9.86 \times 10^{-1}$
	-0.43	-0.69			$^{95}$ Tc + p $\leftrightarrow \gamma + {}^{96}$ Ru		$7.64 \times 10^{-1}$	$6.60 \times 10^{-1}$
$^{102}Pd$	-0.87			$^{101}Pd + n \leftrightarrow \gamma + ^{102}Pd$			$5.62 \times 10^{-1}$	$3.97 \times 10^{-1}$
<sup>112</sup> Sn	-0.88			$^{111}$ Sn + n $\leftrightarrow \gamma$ + $^{112}$ Sn			$7.79 \times 10^{-1}$	6.73 × 10 <sup>-</sup>
<sup>114</sup> Sn	-0.77			$^{113}$ Sn + n $\leftrightarrow \gamma$ + $^{114}$ Sn			$1.82 \times 10^{-1}$	$1.28 \times 10^{-1}$
$^{120}$ Te	-0.64	-0.66			$^{119}\text{Te} + n \leftrightarrow \gamma + ^{120}\text{Te}$		$2.43 \times 10^{-1}$	$1.77 \times 10^{-1}$
<sup>124</sup> Xe	-0.74			$^{123}$ Xe + n $\leftrightarrow \gamma$ + $^{124}$ Xe			$8.25 \times 10^{-2}$	$4.38 \times 10^{-3}$
<sup>126</sup> Xe	-0.75			$^{125}Cs + p \leftrightarrow \gamma + ^{126}Ba$			$1.17 \times 10^{-1}$	$7.41 \times 10^{-3}$
	0.30	0.64	0.65			$^{127}Ba + n \leftrightarrow \gamma + ^{128}Ba$	$5.78 \times 10^{-2}$	$3.59 \times 10^{-3}$
<sup>130</sup> Ba	-0.66			$^{129}Ba + n \leftrightarrow \gamma + ^{130}Ba$			$5.77 \times 10^{-2}$	$3.55 \times 10^{-3}$
<sup>132</sup> Ba	-0.77			$^{131}Ba + n \leftrightarrow \gamma + ^{132}Ba$			$1.07 \times 10^{-1}$	$5.85 \times 10^{-3}$
<sup>136</sup> Ce	-0.69			$^{135}Ce + n \leftrightarrow \gamma + ^{136}Ce$			$1.86 \times 10^{-1}$	$8.94 \times 10^{-3}$
	0.31	0.72			$^{139}Ce + n \leftrightarrow \gamma + {}^{140}Ce$		$8.56 \times 10^{-1}$	$6.09 \times 10^{-1}$
$^{138}Ce$	-0.66			$^{137}Ce + n \leftrightarrow \gamma + ^{138}Ce$			$4.16 \times 10^{-1}$	$2.54 \times 10^{-1}$
	-0.16	-0.19	-0.66			$^{136}Ce + n \leftrightarrow \gamma + ^{137}Ce$	$7.57 \times 10^{-1}$	$4.70 \times 10^{-1}$
<sup>144</sup> Sm	0.70			$^{145}Eu + p \leftrightarrow \gamma + {}^{146}Gd$			$8.06 \times 10^{-1}$	$6.02 \times 10^{-1}$
<sup>152</sup> Gd	-0.74			$^{151}\text{Gd} + n \leftrightarrow \gamma + ^{152}\text{Gd}$			$6.18 \times 10^{-1}$	$3.87 \times 10^{-1}$
	0.43	0.76			$^{153}\text{Gd} + n \leftrightarrow \gamma + ^{154}\text{Gd}$		$5.38 \times 10^{-2}$	$2.78 \times 10^{-3}$
	-0.14	-0.26	-0.73			$^{148}\text{Sm} + \alpha \leftrightarrow \gamma + ^{152}\text{Gd}$	$8.14 \times 10^{-1}$	$5.22 \times 10^{-1}$
<sup>164</sup> Er	-0.78			$^{160}\text{Er} + \alpha \leftrightarrow \gamma + {}^{164}\text{Yb}$			$2.13 \times 10^{-1}$	$1.24 \times 10^{-1}$
$^{180}W$	-0.83			$^{176}W + \alpha \leftrightarrow \gamma + {}^{180}Os$			$1.83 \times 10^{-1}$	$1.04 \times 10^{-1}$
	-0.19	-0.60	-0.68			$^{179}Os + n \leftrightarrow \gamma + ^{180}Os$	$4.89 \times 10^{-2}$	$2.49 \times 10^{-1}$
<sup>196</sup> Hg	-0.83			$^{195}Pb + n \leftrightarrow \gamma + ^{196}Pb$			$2.97 \times 10^{-1}$	$1.89 \times 10^{-1}$
	0.31	0.70			$^{197}\text{Pb} + n \leftrightarrow \gamma + ^{198}\text{Pb}$		$3.28 \times 10^{-1}$	$2.39 \times 10^{-1}$
	0.17	0.35	0.67			$^{199}Pb + n \leftrightarrow \gamma + ^{200}Pb$	$6.37 \times 10^{-1}$	$3.47 \times 10^{-1}$
<sup>92</sup> Nb	0.76			$^{90}$ Zr + p $\leftrightarrow \gamma + ^{91}$ Nb			1.00	9.95 × 10-
<sup>146</sup> Sm	-0.57	-0.75			$^{144}Sm + \alpha \leftrightarrow \gamma + ^{148}Gd$		$9.99 \times 10^{-1}$	$9.65 \times 10^{-1}$
	0.34	0.44	0.79			$^{147}\text{Gd} + n \leftrightarrow \gamma + ^{148}\text{Gd}$	$9.92 \times 10^{-1}$	9.28 × 10 <sup>-</sup>

Nuclide	r <sub>corr,0</sub>	r <sub>corr,1</sub>	r <sub>corr,2</sub>	Key rate Level 1	Key rate Level 2	Key rate Level 3	$X_0 (2 \text{ GK})$ capture	$X_0 (3 \text{ GK})$ capture
<sup>78</sup> Kr	-0.77			$^{77}$ Br + p $\leftrightarrow \gamma + ^{78}$ Kr			$9.63 \times 10^{-2}$	$4.44 \times 10^{-2}$
	0.38	0.66			$^{79}$ Kr + n $\leftrightarrow \gamma + {}^{80}$ Kr		$1.28 \times 10^{-1}$	$7.94 \times 10^{-2}$
<sup>92</sup> Mo	-0.87			$^{91}\text{Nb} + p \leftrightarrow \gamma + ^{92}\text{Mo}$			$8.88 \times 10^{-1}$	$8.24 \times 10^{-1}$
<sup>94</sup> Mo	0.78			$^{95}Mo + n \leftrightarrow \gamma + {}^{96}Mo$			$9.14 \times 10^{-1}$	$7.69 \times 10^{-1}$
<sup>96</sup> Ru	-0.67			$^{92}Mo + \alpha \leftrightarrow \gamma + {}^{96}Ru$			1.00	$9.86 \times 10^{-1}$
<sup>102</sup> Pd	-0.71			$^{101}Pd + n \leftrightarrow \gamma + {}^{102}Pd$			$5.62 \times 10^{-1}$	$3.97 \times 10^{-1}$
$^{112}Sn$	-0.74			$^{111}$ Sn + n $\leftrightarrow \gamma$ + $^{112}$ Sn			$7.79 \times 10^{-1}$	$6.73 \times 10^{-1}$
$^{136}Ce$	0.53	0.66			$^{137}Ce + n \leftrightarrow \gamma + ^{138}Ce$		$4.16 \times 10^{-1}$	$2.54 \times 10^{-1}$
$^{138}Ce$	0.71			$^{139}\text{Ce} + n \leftrightarrow \gamma + {}^{140}\text{Ce}$			$8.71 \times 10^{-1}$	$6.43 \times 10^{-1}$
$^{138}La$	0.94			$^{138}$ La + n $\leftrightarrow \gamma$ + $^{139}$ La			$6.18 \times 10^{-1}$	$4.92 \times 10^{-1}$
$^{144}Sm$	0.79			$^{145}\text{Eu} + p \leftrightarrow \gamma + {}^{146}\text{Gd}$			$8.06 \times 10^{-1}$	$6.02 \times 10^{-1}$
<sup>164</sup> Er	-0.76			$^{160}\mathrm{Er} + \alpha \leftrightarrow \gamma + {}^{164}\mathrm{Yb}$			$2.13 \times 10^{-1}$	$1.24 \times 10^{-1}$
<sup>168</sup> Yb	-0.80			$^{164}$ Yb + $\alpha \leftrightarrow \gamma$ + $^{168}$ Hf			$2.12 \times 10^{-1}$	$1.26 \times 10^{-1}$
	-0.14	-0.67			$^{166}$ Yb + $\alpha \leftrightarrow \gamma$ + $^{170}$ Hf		$1.80 \times 10^{-1}$	$1.10 \times 10^{-1}$
$^{180}$ Ta	-0.88			$^{180}\mathrm{Ta} + \mathrm{n} \leftrightarrow \gamma + ^{181}\mathrm{Ta}$			$7.09 \times 10^{-2}$	$3.96 \times 10^{-2}$
	0.09	0.90			$^{179}\mathrm{Ta} + \mathrm{n} \leftrightarrow \gamma + ^{180}\mathrm{Ta}$		$2.37 \times 10^{-1}$	$1.46 \times 10^{-1}$
$^{180}W$	-0.82			$^{176}W + \alpha \leftrightarrow \gamma + {}^{180}Os$			$1.83 \times 10^{-1}$	$1.04 \times 10^{-1}$
<sup>190</sup> Pt	-0.79			$^{190}$ Pt + n $\leftrightarrow \gamma$ + $^{191}$ Pt			$3.58 \times 10^{-1}$	$1.58 \times 10^{-1}$
<sup>196</sup> Hg	-0.86			$^{195}Pb + n \leftrightarrow \gamma + {}^{196}Pb$			$2.97 \times 10^{-1}$	$1.89 \times 10^{-1}$
	0.17	0.64	0.65			$^{197}\text{Pb} + n \leftrightarrow \gamma + ^{198}\text{Pb}$	$3.28 \times 10^{-1}$	$2.39 \times 10^{-1}$
<sup>92</sup> Nb	0.75			$^{92}$ Zr + p $\leftrightarrow \gamma + ^{93}$ Nb			$9.91 \times 10^{-1}$	$9.76 \times 10^{-1}$
$^{98}Tc$	0.89			$^{96}Mo + p \leftrightarrow \gamma + ^{97}Tc$			$9.50 \times 10^{-1}$	$8.56 \times 10^{-1}$
<sup>146</sup> Sm	-0.65			$^{144}\text{Sm} + \alpha \leftrightarrow \gamma + {}^{148}\text{Gd}$			$9.99 \times 10^{-1}$	$9.65 \times 10^{-1}$
	0.33	0.79		-	$^{147}\text{Gd} + n \leftrightarrow \gamma + ^{148}\text{Gd}$		$9.92 \times 10^{-1}$	$9.28 \times 10^{-1}$

# <u>Future</u>

- Other nucleosynthesis

p-process, vp-process, rp-process, rp-process, r-process etc.



- UK Supercomputer facility
DiRAC



- Improved nuclear uncertainty model

w/ T. Rauscher @UK BRIDGCE meeting (Keele U)

# Summary:

- MC nuclear reaction network code
  - applicable to general nucleosynthesis
  - statistical analysis to find important reactions
     parallelized by OpenMP for shared memory systems

•<u>s-process</u>

- T-dependent uncertainty enhanced by exited state contribution
- $\cdot$  (n,  $\gamma$  ) contributes global uncertainty, while few beta-decay affects uncertainty around branchings
- $\cdot$  key reactions (n,g) and  $\beta$ -decay are identified
  - •weak s: 10+ reactions/decay
  - ·main s:  $\sim$ 50 mostly (n,  $\gamma$ ) reactions

# Backup slides

- T-dependent uncertainty
  - (n,g)
  - beta
- the main s-process key rate full list
- performance tests

## Importance of T-dependent uncertainty:

## ground state contribution: X, by Rauscher, ApJS, 201, 2012

 $X(T) = \frac{r^{\text{lab}}}{r^* G_0(T)} = \frac{\int_0^\infty \sigma^{\text{lab}}(E) \Phi_{\text{MB}}(E, T) dE}{\int_0^\infty \sigma^{\text{eff}}(E) \Phi_{\text{MB}}(E, T) dE}$ 

$$\sigma^{\text{eff}} = \sum_{\mu} \sum_{\nu} \frac{2J_{\mu} + 1}{2J_0 + 1} \frac{E - E_{\mu}}{E} \sigma^{\mu \to \nu} (E - E_{\mu})$$



## T-dependent uncertainty: (n,g)

For details, see T. Rauscher, ApJL, 775, 2011

- Theoretical rates (incl. some experimental ones)
  - basic rates: Reaclib (Rauscher & Thielemann 2000)
  - a constant factor 2
- Experimental rates
  - base rates: KADoNiS v0.3 (Dillemann el al., 2009)
  - the formula: Rauscher, ApJ, 775, 2011

 $U(T) = U_{g.s.}X + U_{e.s.}(1 - X)$ 

- ground state (experimental based):  $u_{g.s.} \sim 1.0 - 1.3$
- excited states (theory based):  $u_{\text{e.s.}} = 5$  (given constant)
- X(T): the fraction of particles in the ground state



## T-dependent uncertainty: beta-decay

- beta-decay: only the ground state 1.3 (30%)
- beta-decay: T-dependent
   (Bruslib: Takahashi & Yokoi 1987, Goriely 1999)

$$U(T) = \frac{u_{\text{g.s.}}}{g_0(T)} + u_{\text{e.s.}} \left(1 - \frac{1}{g_0(T)}\right)$$

- ground state:  $u_{\text{g.s.}} = 1.3 (30 \%)$
- excited states:  $u_{\text{e.s.}} = 10$
- $g_0$ : partition function of the ground state



# main s-process: full list

Nuclide	$r_{ m cor,0}$	$r_{\rm cor,1}$	$r_{\rm cor,2}$	Key reaction	_					
				Level 1						
Sr88	-0.65			${}^{88}\mathrm{Sr}(\mathrm{n},\gamma){}^{89}\mathrm{Sr}$		Sn122	-0.73			$^{122}Sb(+\beta)^{122}Te$
Y89	-0.83			${}^{89}{ m Y}({ m n},\gamma){}^{90}{ m Y}$			0.57	0.86		(
Zr90	-0.89			$^{90}\mathrm{Zr}(\mathrm{n},\gamma)^{91}\mathrm{Zr}$			-0.33	-0.49	-0.96	
Zr91	-0.91			$^{91}$ Zr $(n,\gamma)^{92}$ Zr		Sb121	-0.92			$^{121}\mathrm{Sb}(\mathrm{n},\gamma)^{122}\mathrm{Sb}$
Zr92	-0.92			$^{92}$ Zr(n, $\gamma$ ) $^{93}$ Zr		Te125	-0.92			$^{125}{ m Te}({ m n},\gamma)^{126}{ m Te}$
Zr94	-0.86			$^{94}\mathrm{Zr}(\mathrm{n},\gamma)^{95}\mathrm{Zr}$		Te126	-0.69			$^{126}{ m Te}({ m n},\gamma)^{127}{ m Te}$
Nb93	-0.97			$^{93}$ Zr(n, $\gamma$ ) $^{94}$ Zr		Xe128	0.66			$^{128}I(+\beta)^{128}Xe$
Mo95	-0.85			$^{95}Mo(n,\gamma)^{96}Mo$		Xe132	-0.97			$^{132}$ Xe(n, $\gamma$ ) $^{133}$ Xe
Mo96	-0.94			${}^{96}Mo(n,\gamma){}^{97}Mo$		Cs133	-0.89			$^{133}Cs(n,\gamma)^{134}Cs$
Mo97	-0.87			$^{97}Mo(n,\gamma)^{98}Mo$		Ba134 Do125	-0.85			$^{135}Ba(n,\gamma)^{136}Ba$
Mo98	-0.94			$^{98}Mo(n,\gamma)^{99}Mo$		Da130 Ba136	-0.70			$^{136}Ba(n,\gamma)^{137}Ba$
Ru99	-0.91			$^{99}Tc(n,\gamma)^{100}Tc$		Ba137	-0.84			$^{137}Ba(n,\gamma)^{138}Ba$
Ru100	-0.93			$^{100}$ Ru(n, $\gamma$ ) $^{101}$ Ru		Ba138	-0.65	-0.71		Du(11,7) Du
Ru102	-0.86			$^{102}Ru(n,\gamma)^{103}Ru$		La139	-0.87	0112		$^{139}$ La $(n,\gamma)^{140}$ La
Rh103	-0.95			$^{103}$ Rh(n, $\gamma$ ) $^{104}$ Rh		Gd152	0.59	0.60	0.66	
Pd104	-0.97			$^{104}Pd(n,\gamma)^{105}Pd$		Er166	-0.81			$^{166}{ m Er}({ m n},\gamma)^{167}{ m Er}$
Pd106	-0.97			$^{106}Pd(n,\gamma)^{107}Pd$		Er168	-0.86			$^{168}\mathrm{Er}(\mathrm{n},\gamma)^{169}\mathrm{Er}$
Pd108	-0.96			$^{108}Pd(n,\gamma)^{109}Pd$		Ta181	-0.84			$^{181}$ Ta $(n,\gamma)^{182}$ Ta
Ag107	-0.81			$^{107}Pd(n,\gamma)^{108}Pd$		Os187	-0.86			$^{187}$ Os(n, $\gamma$ ) $^{188}$ Os
Ag109	-0.80			$109 \operatorname{Ag}(n,\gamma)^{110} \operatorname{Ag}$		Pt192	-0.89			$^{192}$ Pt $(n,\gamma)$ $^{193}$ Pt
In115	-0.97			$^{115}In(n,\gamma)^{116}In$		Hg198	-0.63	-0.65	-0.68	20011-(
Sn110	-0.97			$119 Sn(n \sim) 120 Sn$		Hg200	-0.67			$^{200}$ Hg(n, $\gamma$ ) $^{201}$ Hg
Juiio	-0.01			51(1,7) 51	-	Hg201	-0.77			$-\gamma Hg(n,\gamma) -Hg$

### Large-scale MC calculations

### for shared memory systems

- Fortran + OpenMP
- parallelized well
- optimized code/matrix library for large scared memory computers (multi threads)
- <u>Computer resources</u>
   **numascale**



Shyne cluster @Keele (ERC) performance tests of matrix solvers on shared memory system



Cosmos2 @Cambridge (UK DiRAC facility, STFC)

