# $\eta N$ interactions in the nuclear medium and $\eta$ -nuclear bound states

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PLB 725 (2013) 334; NPA 925 (2014) 126; PLB 747 (2015) 345



International Nuclear Physics Conference Adelaide Convention Centre, Australia 11-16 September 2016

## $\eta$ nuclei - status

- Haider, Liu (PLB 172 (1986) 257, PRC 34 (1986) 1845) moderate attractive  $\eta N$  interaction with scattering length  $a_{\eta N} \sim 0.27 + i0.22$  fm  $\Rightarrow \exists$  of  $\eta$  nuclear bound states (starting <sup>12</sup>C)
- Numerous studies since then yielding Rea<sub>ηN</sub> from 0.2 fm to 1 fm chiral coupled channel models Rea<sub>ηN</sub> < 0.3 fm;</li>
   K matrix methods fitting πN and γN reaction data in the N\*(1535) resonance region Rea<sub>ηN</sub> ~ 1 fm → bound states already in He isotopes
- Strong final-state interaction have been noted in p- and d-initiated  $\eta$  production (COSY-ANKE, COSY-GEM, LNS-SPES2,3,4)
- ${}^{25}_{\eta}Mg$  ? (COSY-GEM, PRC 79 (2009) 012201(R))  $p + {}^{27}$  Al  $\rightarrow^{25}_{\eta}$  Mg  $+ {}^{3}$  He;  ${}^{25}_{\eta}$ Mg  $\rightarrow (\pi^- + p) + X$  $B_{\eta} = 13.1 \pm 1.6$  MeV and  $\Gamma_{\eta} = 10.2 \pm 3.0$  MeV.
- But NO decisive experimental evidence so far. (negative results for <sup>3</sup><sub>η</sub>He (photoproduction on <sup>3</sup>He - MAMI, PLB 709 (2012) 21.) and for <sup>4</sup>ηHe (dd →<sup>3</sup>Hepπ<sup>-</sup> - WASA@COSY PRC 87 (2013)035204.)

## $\eta N$ interactions

- chiral SU(3)<sub>L</sub> ×SU(3)<sub>R</sub> meson-baryon effective Lagrangian for  $\{\pi, K, \eta\} + \{N, \Lambda, \Sigma, \Xi\}$
- $\exists$  resonances  $\Rightarrow \chi \mathsf{PT}$  not applicable  $\rightarrow$
- nonperturbative coupled-channel resummation techniques

$$T_{ij} = V_{ij} + V_{ik}G_{kl}T_{lj}, V_{ij}$$
 derived from  $\mathcal{L}_{\chi}$ 

Effective potentials are constructed to match the chiral meson-baryon amlitudes (up to NLO order)



- Channels involved: πΝ, ηΝ, ΚΛ, ΚΣ
- Model parameters fixed by fitting low-energy meson-nucleon data:  $\pi N \rightarrow \eta N$  production X-section:  $\pi N$  amplitudes from SAID database:



 $(S_{11} \text{ and } S_{31} \text{ partial waves})$ 



# $\eta N$ scattering amplitudes

- $\eta N$  amplitudes for various models differ considerably
- Strong energy dependence of the scattering amplitudes !



line	$a_{\eta N}$ [fm]	model
dotted	0.46+i0.24	N. Kaiser, P.B. Siegel, W. Weise, PLB 362 (1995) 23
short-dashed	0.26+i0.25	T. Inoue, E. Oset, NPA 710 (2002) 354 (GR)
dot-dashed	0.96+i0.26	A.M. Green, S. Wycech, PRC 71 (2005) 014001 (GW)
long-dashed	0.38+i0.20	M. Mai, P.C. Bruns, UG. Meißner, PRD 86 (2012) 094033 (M2)
full	0.67+i0.20	A. Cieply, J. Smejkal, Nucl. Phys. A 919 (2013) 334 (CS)

## $\eta$ in few-body systems

#### Variational calculation in hypersherical basis: N. Barnea, E. Friedman, A. Gal, PLB 747 (2015) 345

2-body interactions

- NN: Argonne AV4' potential, Minnesota MN (central) potential
- ηN: complex energy-dependent local potential derived from the full chiral coupled-channels model:

$$v_{\eta N}(E, r) = -\frac{4\pi}{2\mu_{\eta N}} b(E) \rho_{\Lambda}(r),$$
  
where  $E = \sqrt{s} - \sqrt{s_{\text{th}}}, \quad \rho_{\Lambda}(r) = (\frac{\Lambda}{2\sqrt{\pi}})^3 exp\left(-\frac{\Lambda^2 r^2}{4}\right)$ 

b(E) fitted to phase shifts  $\delta$  derived from  $F_{\eta N}(E)$  in GW and CS models

# $\eta$ in few-body systems

• Energy dependence of b(E)



 $F_{\eta N}(E)$  generated from  $v_{\eta N}(E)$  (GW), compared with the amplitude generated from  $v_{\eta N}(E = 0)$  (gw).

## $\eta$ in few-body systems

#### Energy dependence of $v_{\eta N}(\sqrt{s})$

• A nucleons  $+ \eta$  meson:

$$m{s} = (\sqrt{s_{ ext{th}}} - B_\eta - B_N)^2 - (ec{p_\eta} + ec{p_N})^2 \le s_{ ext{th}}$$
  
where  $\sqrt{s_{ ext{th}}} = m_N + m_\eta$ 

near threshold approximated by:

$$\begin{split} \sqrt{s} &= \sqrt{s_{\text{th}}} + \delta\sqrt{s}, \quad \delta\sqrt{s} < 0! \\ \langle \delta\sqrt{s} \rangle &= -\frac{B}{A} - \frac{A-1}{A}B_{\eta} - \xi_N \frac{A-1}{A} \langle T_{N:N} \rangle - \xi_\eta \left(\frac{A-1}{A}\right)^2 \langle T_{\eta} \rangle, \\ \text{where } B &= \text{total binding energy, } \xi_{N(\eta)} = m_{N(\eta)} / (m_N + m_\eta), \\ T_{\eta} &= \eta \text{ kin. energy, } T_{N:N} = \text{pairwise } NN \text{ kin. energy} \end{split}$$

•  $\langle \delta \sqrt{s} \rangle \Rightarrow$  selfconsistency

• Energy dependence  $\Rightarrow$  selfconsistency



The  $\eta NNN$  g.s. energy  $E_{g.s.}$  (solid curves) +  $\delta \sqrt{s}$  (dashed curves)

- Conversion widths  $\Gamma$  of  $\eta$  nuclear few-body systems perturbative estimate:  $\Gamma = -2\langle \Psi_{gs} | \text{Im} V_{\eta N} | \Psi_{gs} \rangle$
- $\eta NN NO$  bound state
- $\eta NNN$  bound state ?

NN int.	E(NNN)	$E_{ m gs}^{ m no~sc}$	$E_{\eta}^{\text{no sc}}$	$\delta\sqrt{s_{ m sc}}$	$E_{ m gs}^{ m sc}$	$E_{\eta}^{ m sc}$	$\Gamma_{ m gs}^{ m sc}$
MN	-8.38	-11.26	2.88	-13.52	-9.33	-0.95	6.76
AV4'	-8.99	-11.33	2.34	-15.83	-9.03	-0.04	7.88

A. Cieply, E. Friedman, A. Gal, J. Mares, PLB 725 (2013) 334, NPA 925 (2014) 126

• K.-G. equation:

$$\left[\omega_\eta^2+ec
abla^2-m_\eta^2-\Pi_\eta(\omega_\eta,
ho)
ight]\phi_\eta=0$$

complex energy  $\omega_\eta = m_\eta - B_\eta - \mathrm{i} \Gamma_\eta/2$ 

• 
$$\Pi_{\eta}(\omega_{\eta},\rho) = 2\omega_{\eta}V_{\eta} = -4\pi \frac{\sqrt{s}}{E_{N}}F_{\eta N}(\sqrt{s},\rho)\rho$$

•  $\eta$  in a nucleus  $\Rightarrow$  polarized (compressed)  $\rho \longrightarrow \Pi_{\eta}(\rho)$  $\Rightarrow$  selfconsistent solution Selfenergy operator

$$\Pi_{\eta}(\omega_{\eta}) = 2\,\omega_{\eta}V_{\eta} = -4\pi \frac{\sqrt{s}}{E_{N}}F_{\eta N}(\sqrt{s},\rho)\,\rho$$

- $F_{\eta N} = \eta N$  scattering amplitude with two-body argument:  $\sqrt{s} (s = (\omega_{\eta} + E_N)^2 - (\vec{p}_{\eta} + \vec{p}_N)^2)$
- $\eta N$  c.m. frame  $\rightarrow \eta$ -nucleus c.m. frame  $\Rightarrow \vec{p}_{\eta} + \vec{p}_{N} \neq 0$   $\Rightarrow \sqrt{s} \approx m_{\eta} + m_{N} - B_{\eta} - B_{N} - \xi_{N} \frac{p_{N}^{2}}{2m_{N}} - \xi_{\eta} \frac{p_{\eta}^{2}}{2m_{\eta}} = E_{\text{th}} + \delta \sqrt{s},$  $\delta \sqrt{s} = B_{N} \frac{\rho}{\bar{\rho}} - \xi_{N} B_{\eta} \frac{\rho}{\rho_{0}} - \xi_{N} T_{N} (\frac{\rho}{\rho_{0}})^{2/3} + \xi_{\eta} \text{Re} V_{\eta} (\sqrt{s}, \rho)$
- $\rho$  = nucl. medium density (RMF calculations)
- $V_\eta, B_\eta \Rightarrow$  self-consistent solution

## Free space amplitudes $\rightarrow$ in-medium amplitudes

• WRW method - T. Wass, M. Rho, W. Weise, NPA 617 (1997) 449.

$$F_{\eta N}(\sqrt{s},
ho) = rac{F_{\eta N}(\sqrt{s})}{1+\xi(
ho)(\sqrt{s}/E_N)F_{\eta N}(\sqrt{s})
ho} \; ,$$

$$\xi(\rho) = \frac{9\pi}{4\rho_f^2} I(\kappa), \qquad I(\kappa) = 4 \int_0^\infty \frac{dt}{t} \exp(iqt) j_1^2(t), \qquad \kappa = \frac{1}{\rho_f} \sqrt{2m_\eta (B_\eta + i\Gamma/2)}.$$

Chiral coupled-channels model - A. Cieply, J. Smejkal, NPA 919 (2013) 334.
 multi-channel L.-Sch. equation:

$$F = V + VGF, \quad F, V \text{ in separable form,}$$

$$G_n(\sqrt{s}; \rho) = -4\pi \int_{\Omega_n(\rho)} \frac{d^3\vec{p}}{(2\pi)^3} \frac{g_n^2(p)}{k_n^2 - p^2 - \prod_n(\sqrt{s}, \vec{p}; \rho) + i0}$$

 $\Omega_n(\rho) \rightarrow \text{intermediate } N \text{ energy is above Fermi level (Pauli blocking)}$  $\Pi \rightarrow \text{hadron self-energies in } G \ (+SE \text{ option})$ 

 $\Rightarrow$  self-consistency

#### • Energy dependence of $f_{\eta N}(\sqrt{s})$



chiral CS model (A. Cieply, J. Smejkal, Nucl. Phys. A 919 (2013) 334)

dotted curve: free-space, dot-dashed: Pauli blocked, full: Pauli blocked + hadron selfenergies

 Nuclear medium reduces the ηN attraction at threshold, the amplitude becomes smaller when going subthreshold

#### $\eta$ nuclear states

#### Energy dependence of $V_{\eta}(\sqrt{s}) \leftarrow$ due to $N^*(1535)$

• In-medium (subthreshold) energy shift:

 $\delta\sqrt{s} = -B_N \frac{\rho}{\bar{\rho}} - \xi_N B_\eta \frac{\rho}{\rho_0} - \xi_N T_N (\frac{\rho}{\rho_0})^{2/3} + \xi_\eta \operatorname{Re} V_\eta(\sqrt{s},\rho)$ 



#### • $B_{\eta}, V_{\eta}, \rho \Rightarrow$ selfconsistent solution $\rightarrow$

40 - 60 MeV energy shift at  $\rho_0$  – larger than shift by  $B_\eta$  (GR) or by 30 MeV (Haider, Liu)

## $\eta$ nuclear states

 Sensitivity to the energy shift: selfconsistent δ√s reduces both 1s B<sub>n</sub> and Γ<sub>n</sub>



• GR widths too large to resolve  $\eta$  bound states !

#### • Model dependence:



• Larger Re  $a_{\eta N}$  gives larger  $B_{\eta}$  vs. no relation between Im  $a_{\eta N}$  and  $\Gamma_{\eta}$ 

#### • Predictions of GW and CS models:

all states in selected nuclei are shown; both models give small widths ( $\Gamma_\eta < 5~{\rm MeV})$ 



- Large energy shift and rapid decrease of the  $\eta N$  amplitudes lead to relatively small binding energies and widths of the calculated  $\eta$  nuclear bound states
- CS and GW models predict  $\eta$  nuclear states with small widths (< 5 MeV)  $\rightarrow$  this might encourage further attempts to produce and identify  $\eta$  nuclear bound states.
- additional width contribution not considered in this work due to  $\eta N \rightarrow \pi \pi N$  and  $\eta NN \rightarrow NN \Rightarrow$  estimated to add few MeV
- Subthreshold behavior of  $F_{\eta N}$  is crucial to decide whether  $\eta$  nuclear states exist, in which nuclei, and if their widths are small enough to be resolved in experiment.