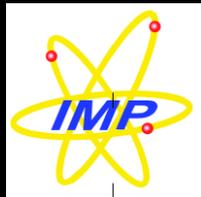
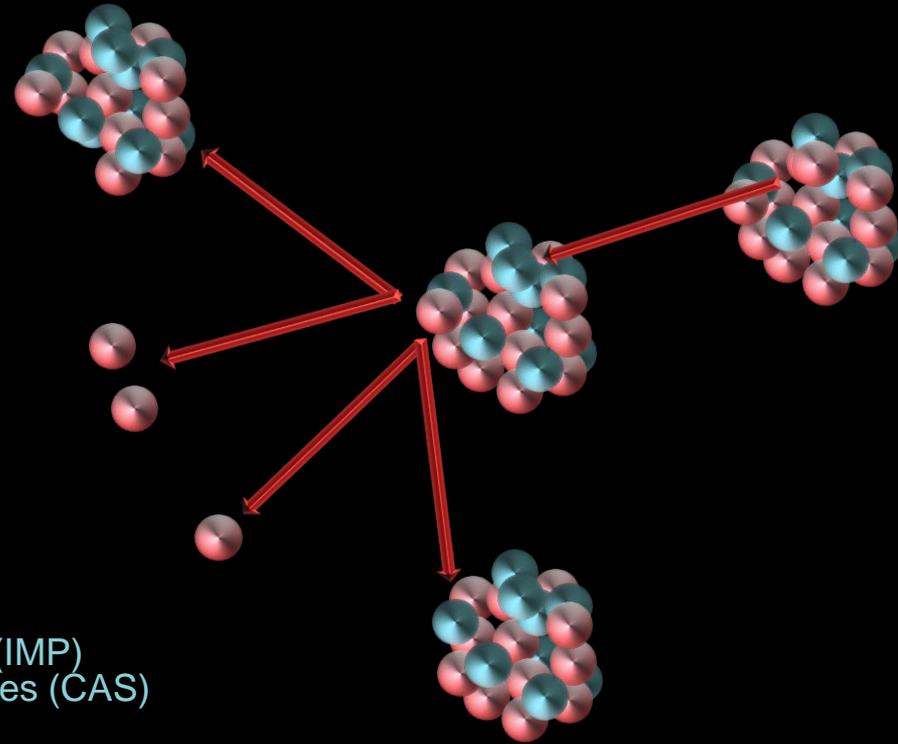


Theoretical Description of β -delayed Proton Emission of Proton Rich *sd*- and *pf*- shell nuclei

Yi Hua LAM (藍乙華), Nadezda A. Smirnova



中国科学院近代物理研究所
Institute of Modern Physics (IMP)
Chinese Academy of Sciences (CAS)

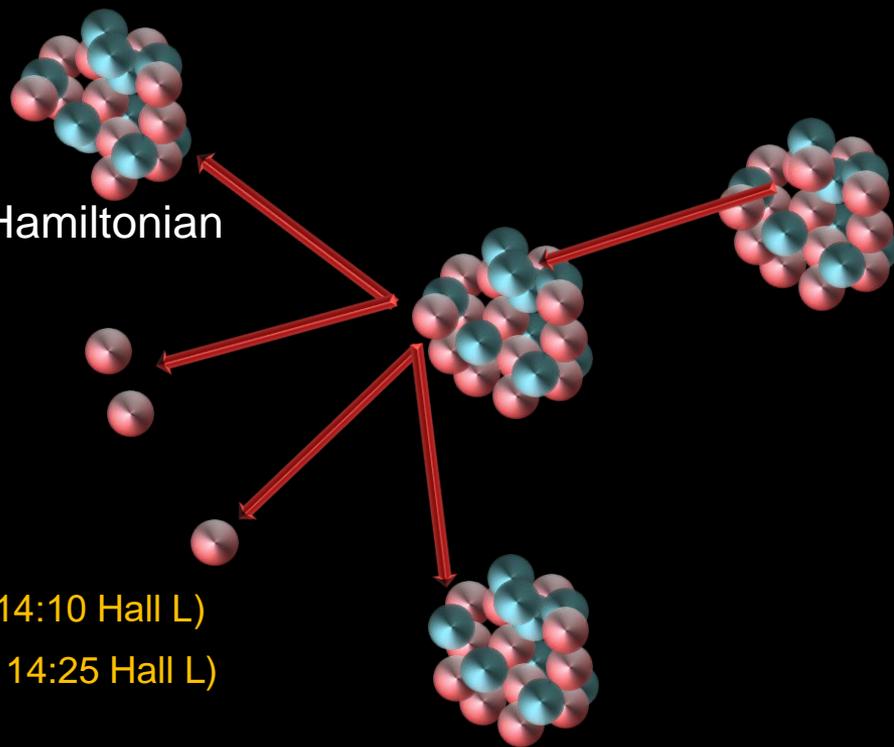


2016 International Nuclear Physics Conference,
11-13 Sept. 2016, Adelaide, Australia.

Theoretical Description of β -delayed Proton Emission of Proton Rich *sd*- and *pf*- shell nuclei

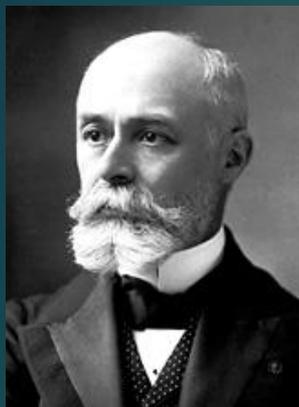
Yi Hua LAM (藍乙華), Nadezda A. Smirnova

- Background
 - Physics motivation
- Shell-model isospin non-conserving (INC) Hamiltonian
 - Construction of INC Hamiltonian
- β -delayed proton emission
 - *sd*-shell: ^{22}Si , ^{23}Si , ^{24}Si , ^{25}Si ,
 - *pf*-shell:
 - ^{53}Ni ,
 - ^{56}Zn (see W. Richter's talk, on Thursday 14:10 Hall L)
 - ^{52}Ni (see Y. H. Zhang's talk, on Thursday 14:25 Hall L)
- Summary and Perspectives



Physics motivation

Discovery of radioactivity



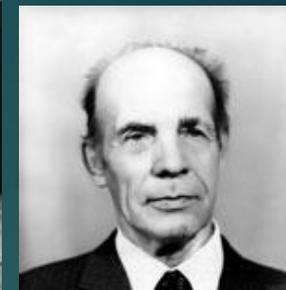
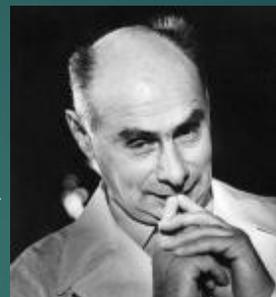
Henri Becquerel

3 groups of radioactivity:
negative, positive,
and electrically
neutral

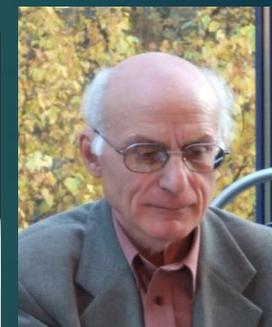


F. Joliot and I. Curie

β^+



G.N. Flerov and K.A. Petzhak
spontaneous fission



V. A. Karnaukhov and G. M. Ter-Akopian
 β -delayed proton emission



S. Hofmann:

p

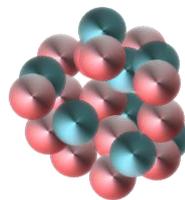
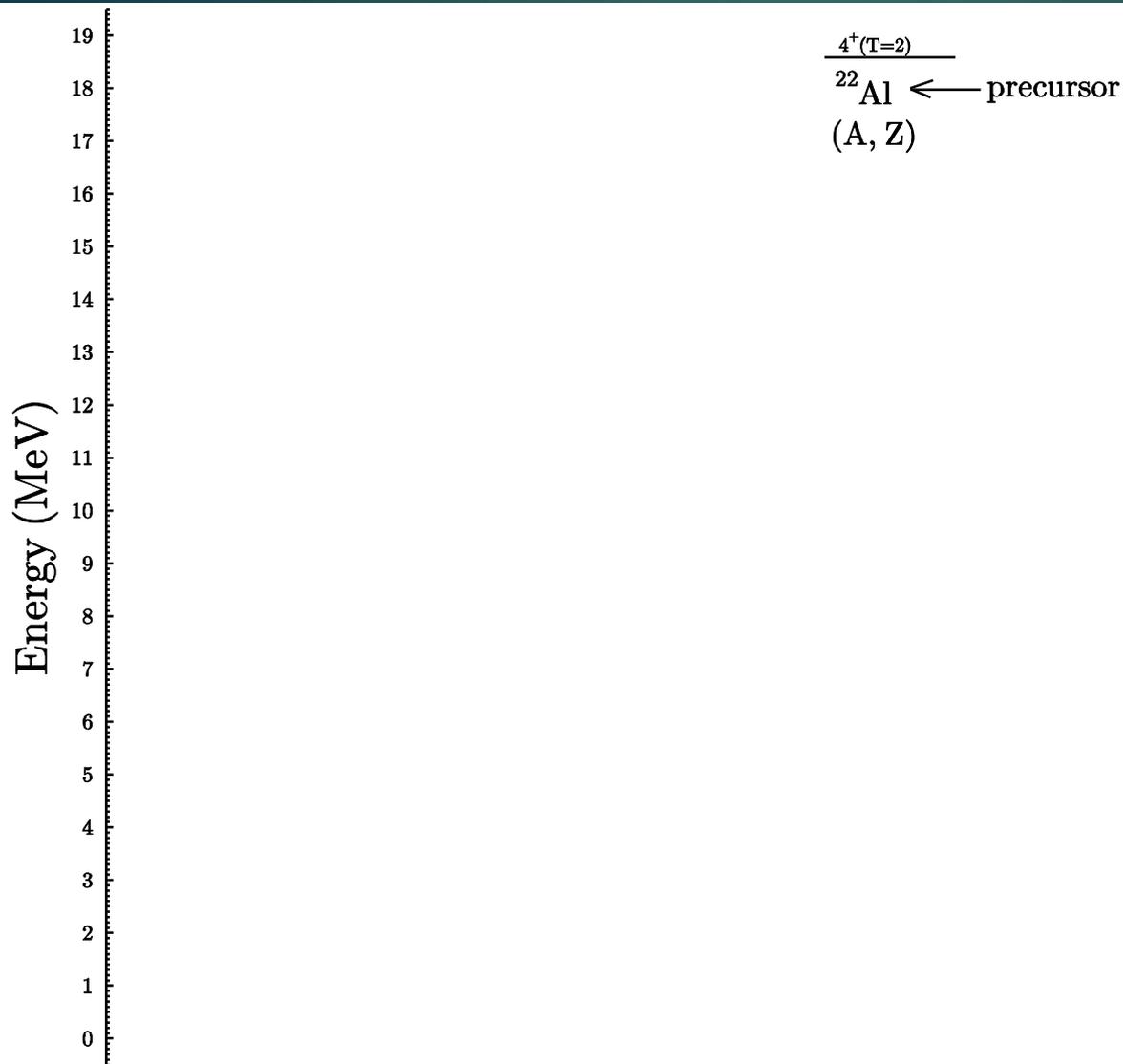


M. Pfutzner:

$2p$

Physics motivation

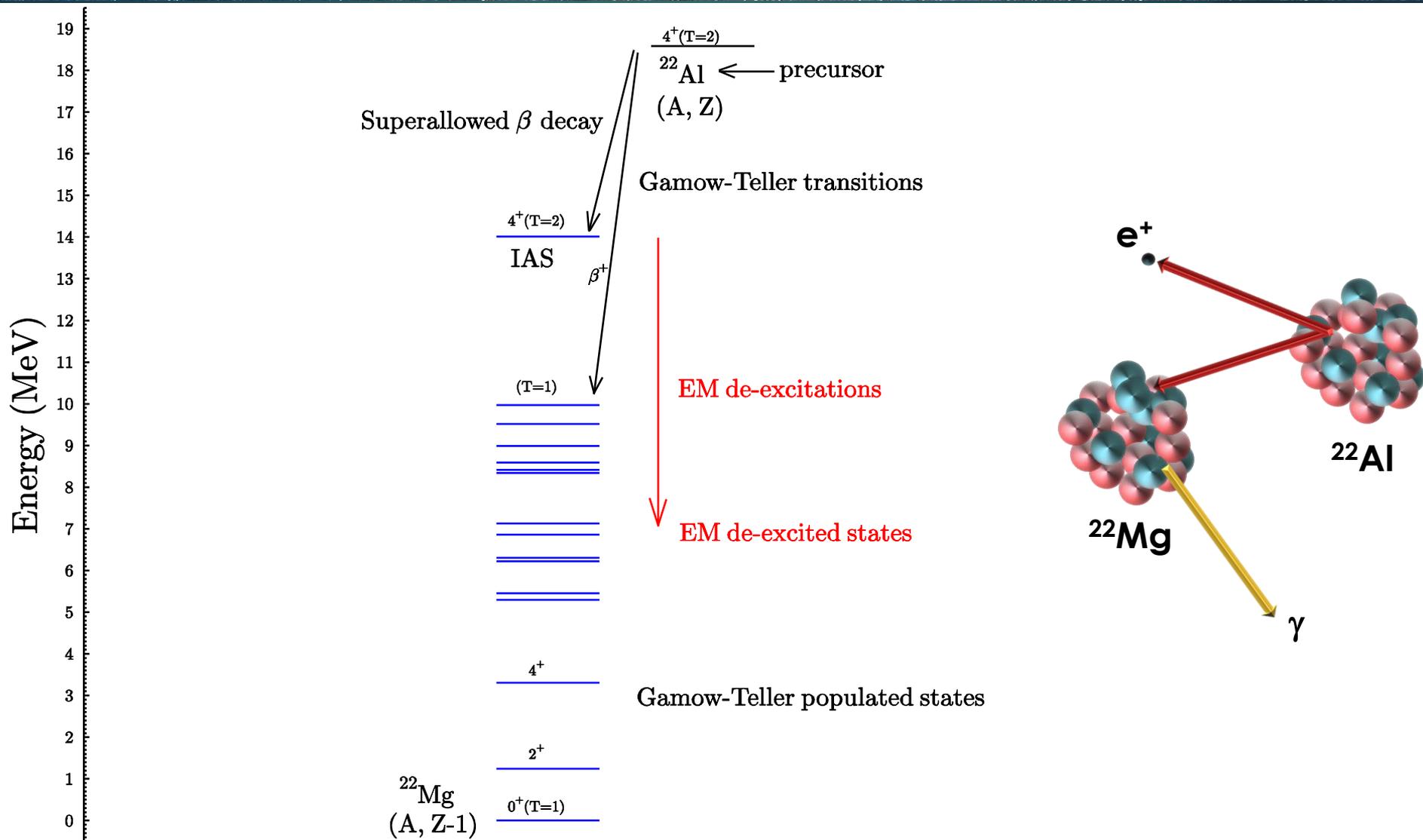
β -delayed proton emissions



^{22}Al

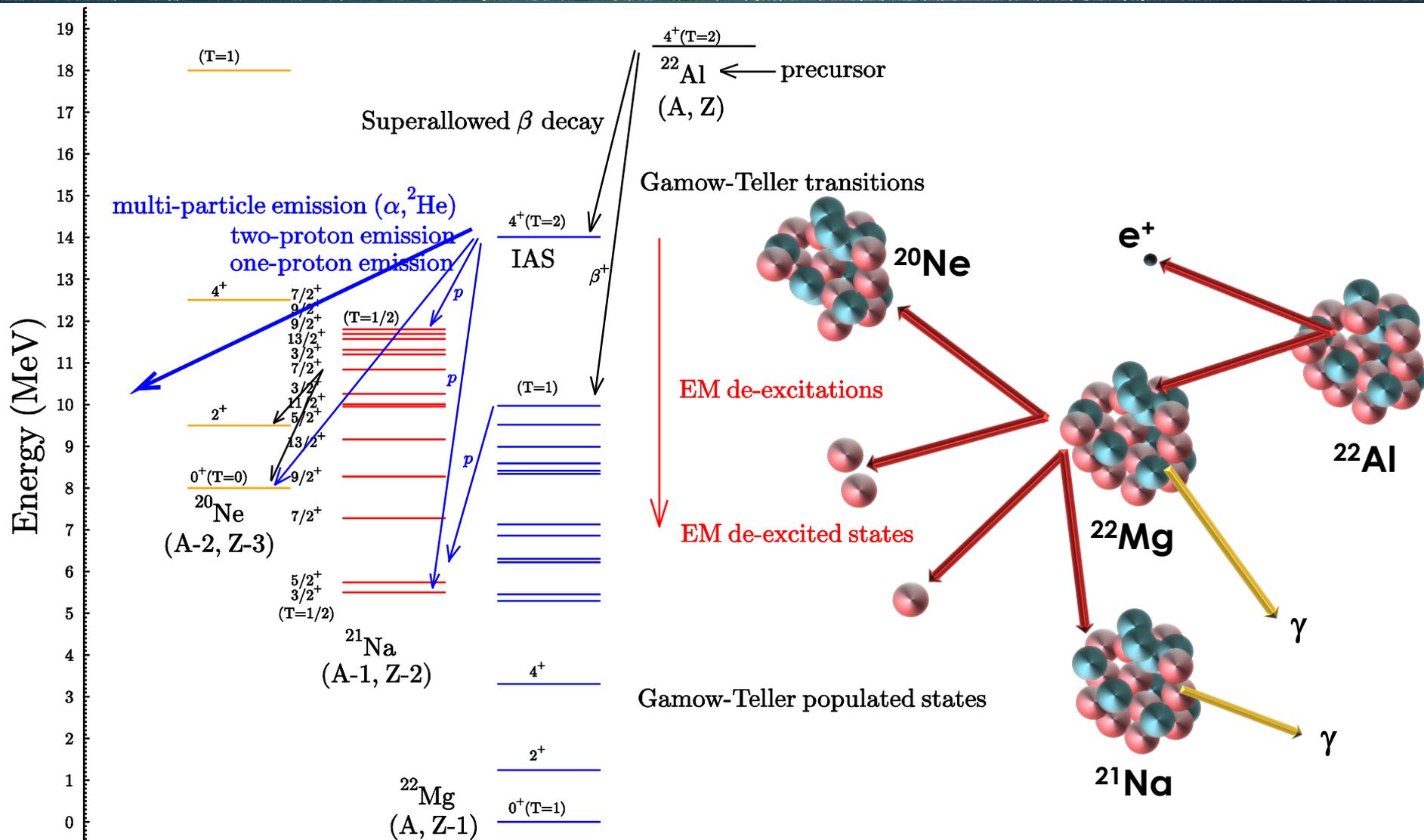
Physics motivation

β -delayed proton emissions



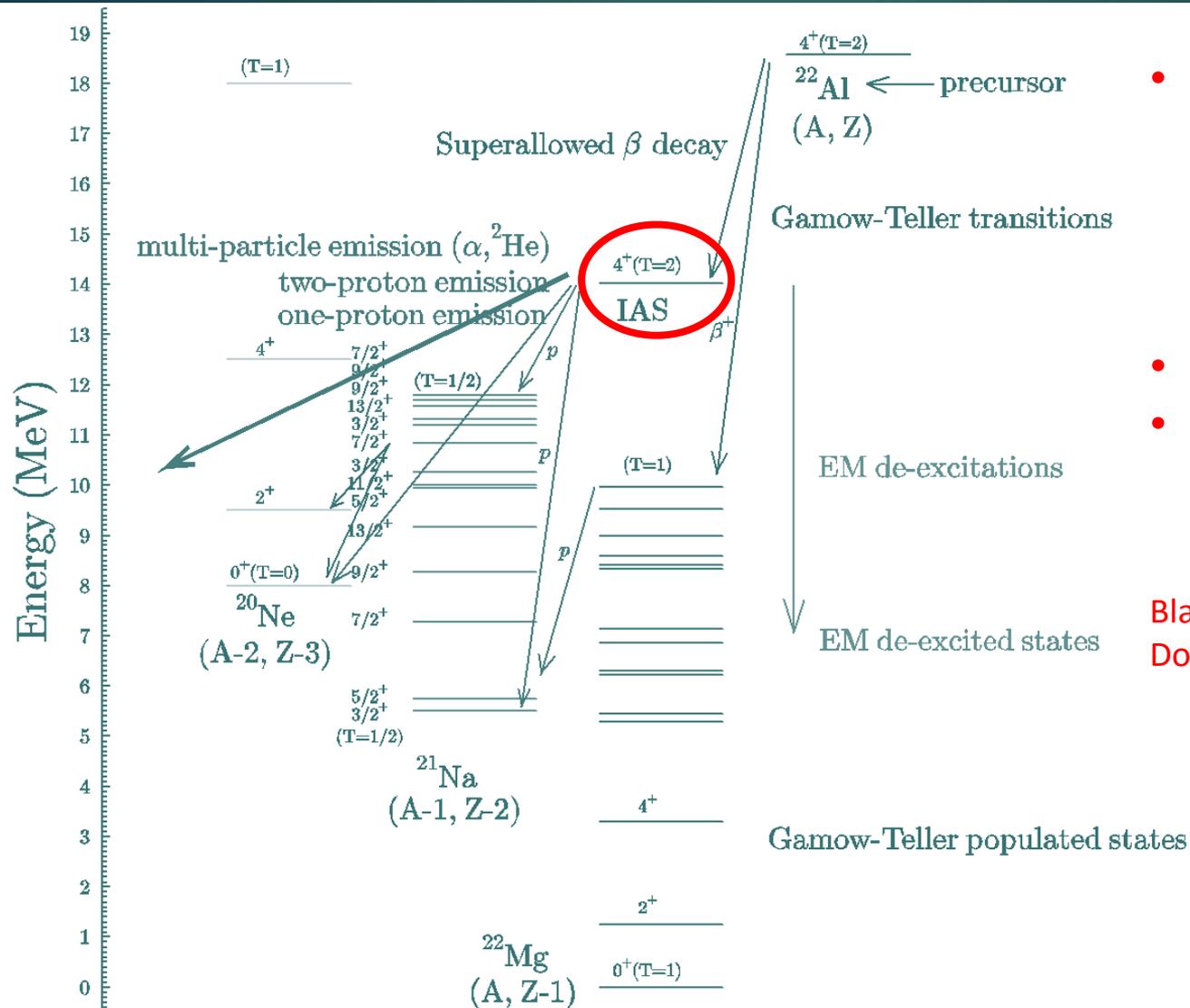
Physics motivation

β -delayed proton emissions



Physics motivation

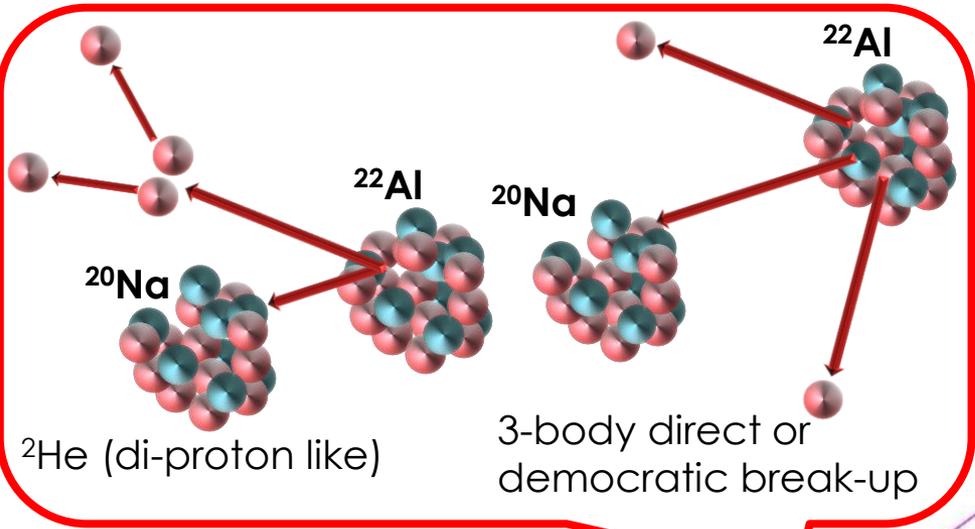
β -delayed proton emissions



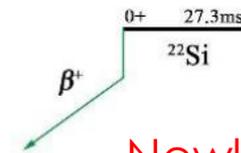
- Determining the **IAS**, and *isovector* and *isotensor* coefficients of **isobaric multiplet mass equation (IMME)**
- Breaking down of IMME?
- Estimating/extrapolating masses of proton-rich nuclei

Blank & Borge (PPNP 60, 403)
Dossat+ (NPA 792, 18)

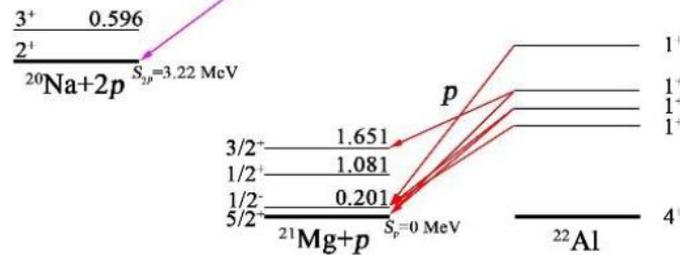
β -delayed proton emissions



$$M({}^{22}\text{Si}) = M({}^{22}\text{Al}, \text{IAS}) + \Delta E_{C(g.s.)} - \Delta_{nH}$$



Newly measured mass of ${}^{22}\text{Si}$ is very close to the prediction of 3 nucleon forces Holt+ (PRL 110 022502)

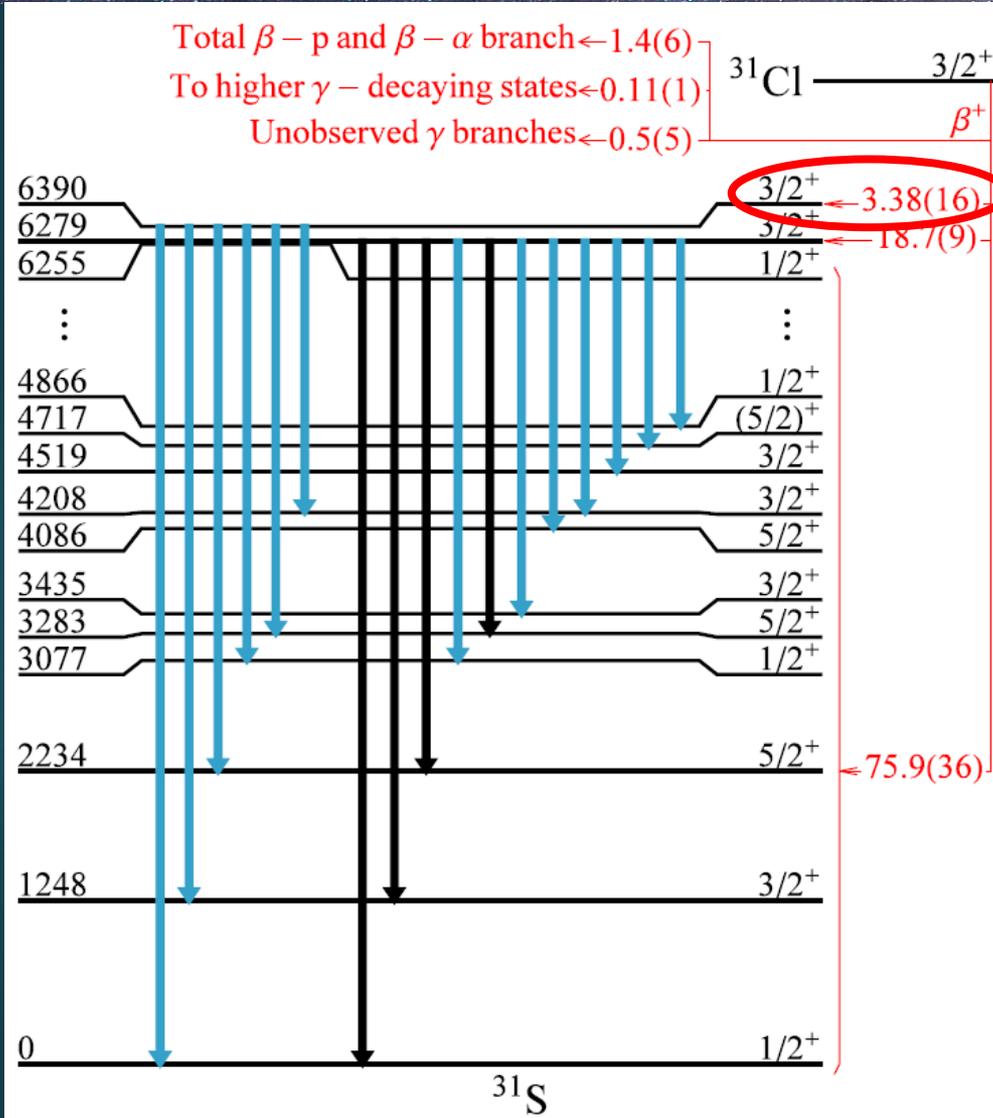


Xin Xing XU, Chen Jian LIN, Li Jie SUN+ (Preliminary)

YHL & N. A. Smirnova calculated for Xin Xing XU+ (in preparation)

Physics motivation

Determining structure of proton-rich nuclei

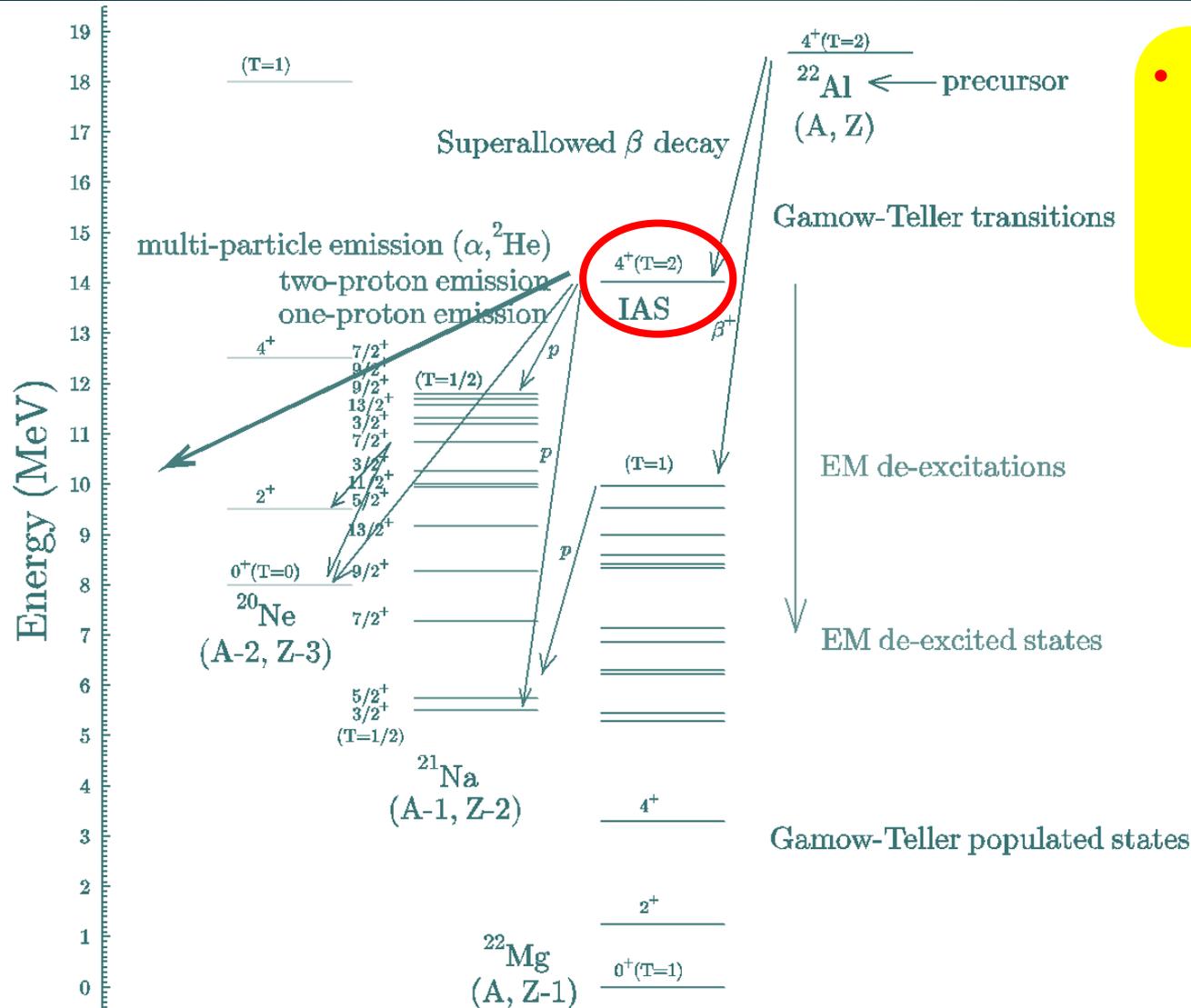


Obtaining the pivotal level of ^{31}S which determines $^{30}\text{P}(p,\gamma)^{31}\text{S}$ reaction rates that influence nova nucleosynthesis

Bennett+ (PRL 116, 102502)

Physics motivation

β -delayed proton emissions



• Only uses **one set of Hamiltonian** to describe all physical phenomena of a given partial decay scheme.

Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian



Isospin Conserving Nucl. Hamiltonian

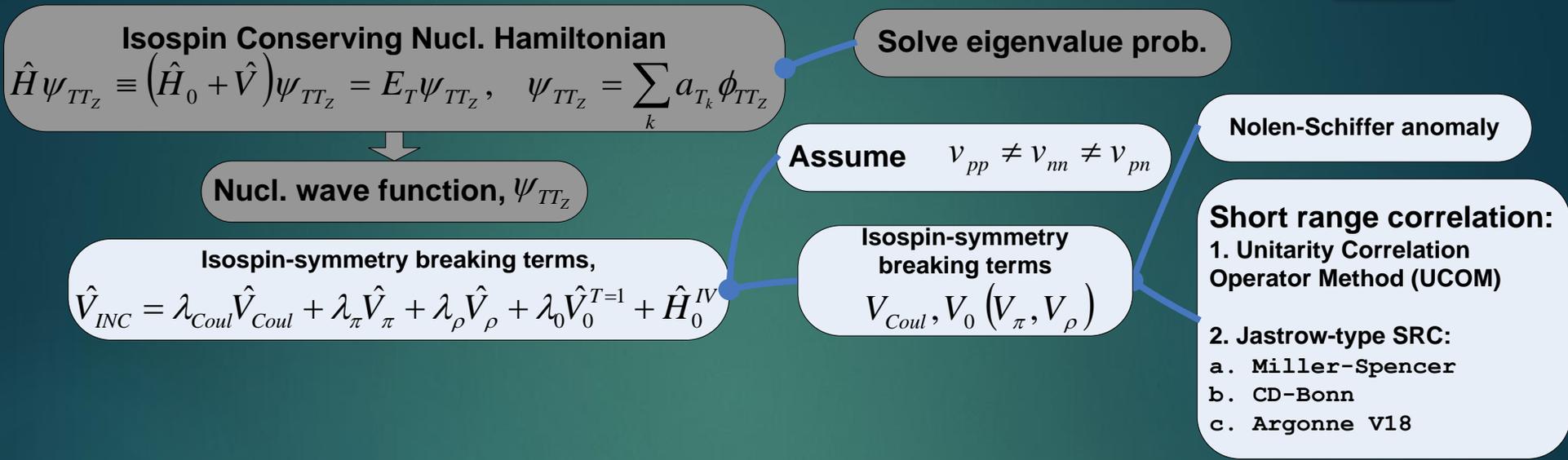
$$\hat{H}\psi_{TT_Z} \equiv (\hat{H}_0 + \hat{V})\psi_{TT_Z} = E_T\psi_{TT_Z}, \quad \psi_{TT_Z} = \sum_k a_{T_k}\phi_{TT_Z}$$

Solve eigenvalue prob.

Nucl. wave function, ψ_{TT_Z}

Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian

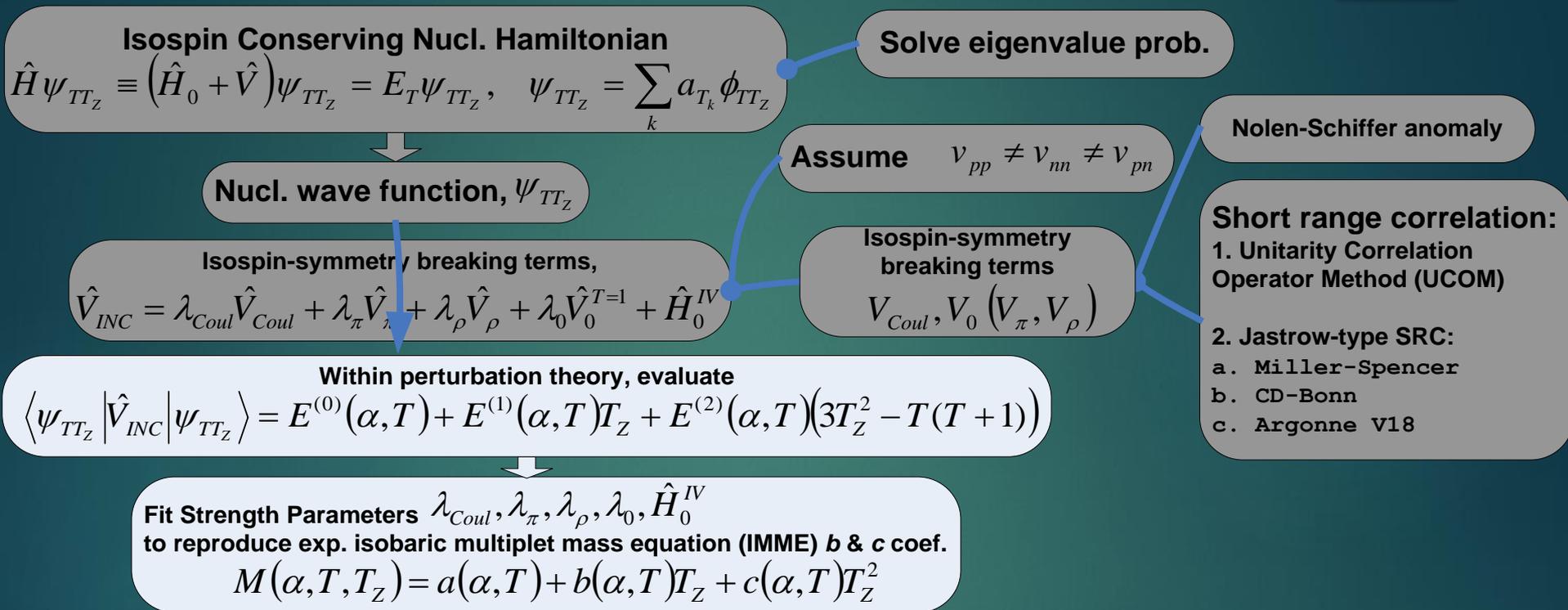


YHL & N. A. Smirnova & E. Caurier (PRC 87, 054304)

The β -delayed proton emission of proton-rich *sd*- and *pf*-shell nuclei

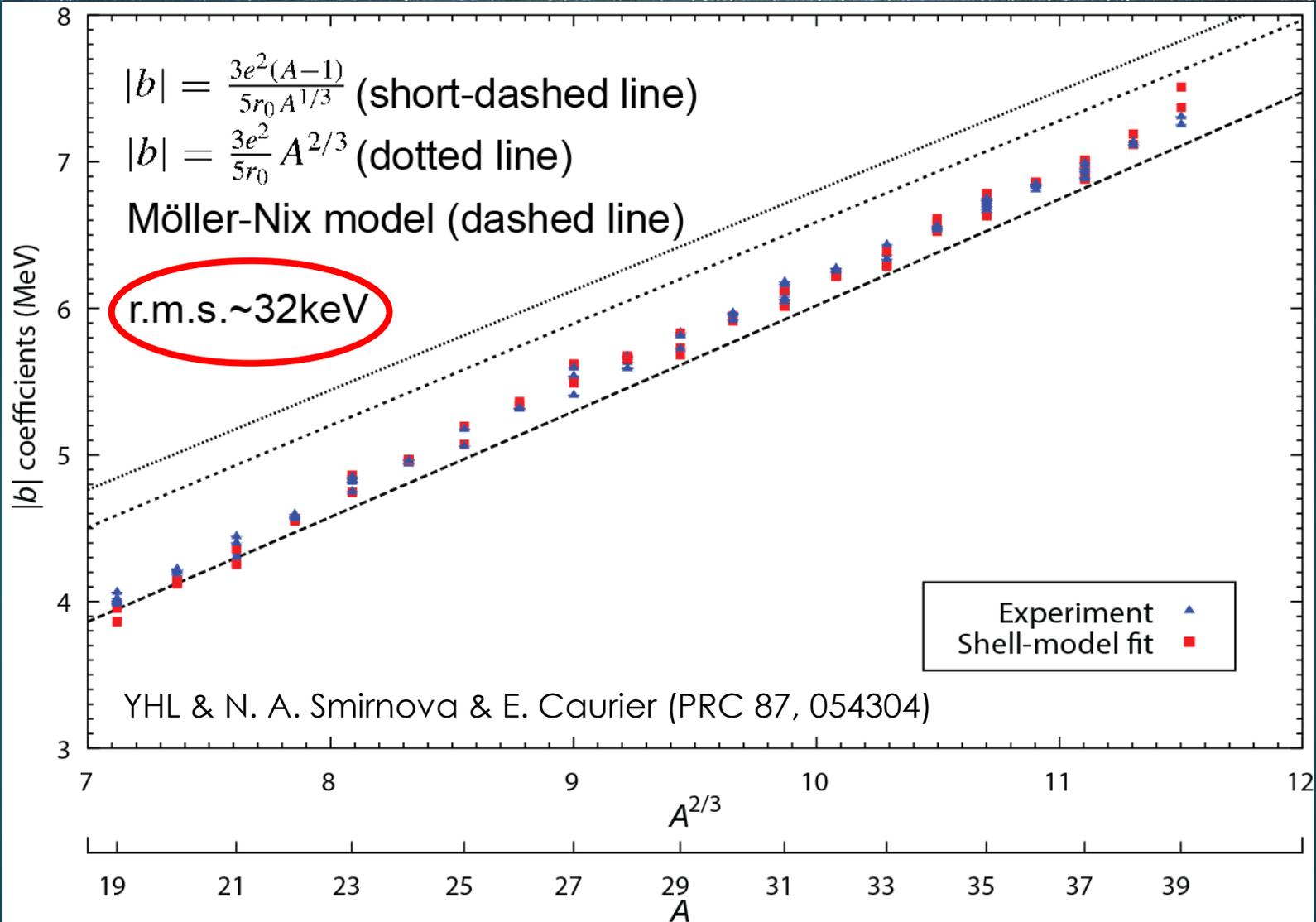
Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian



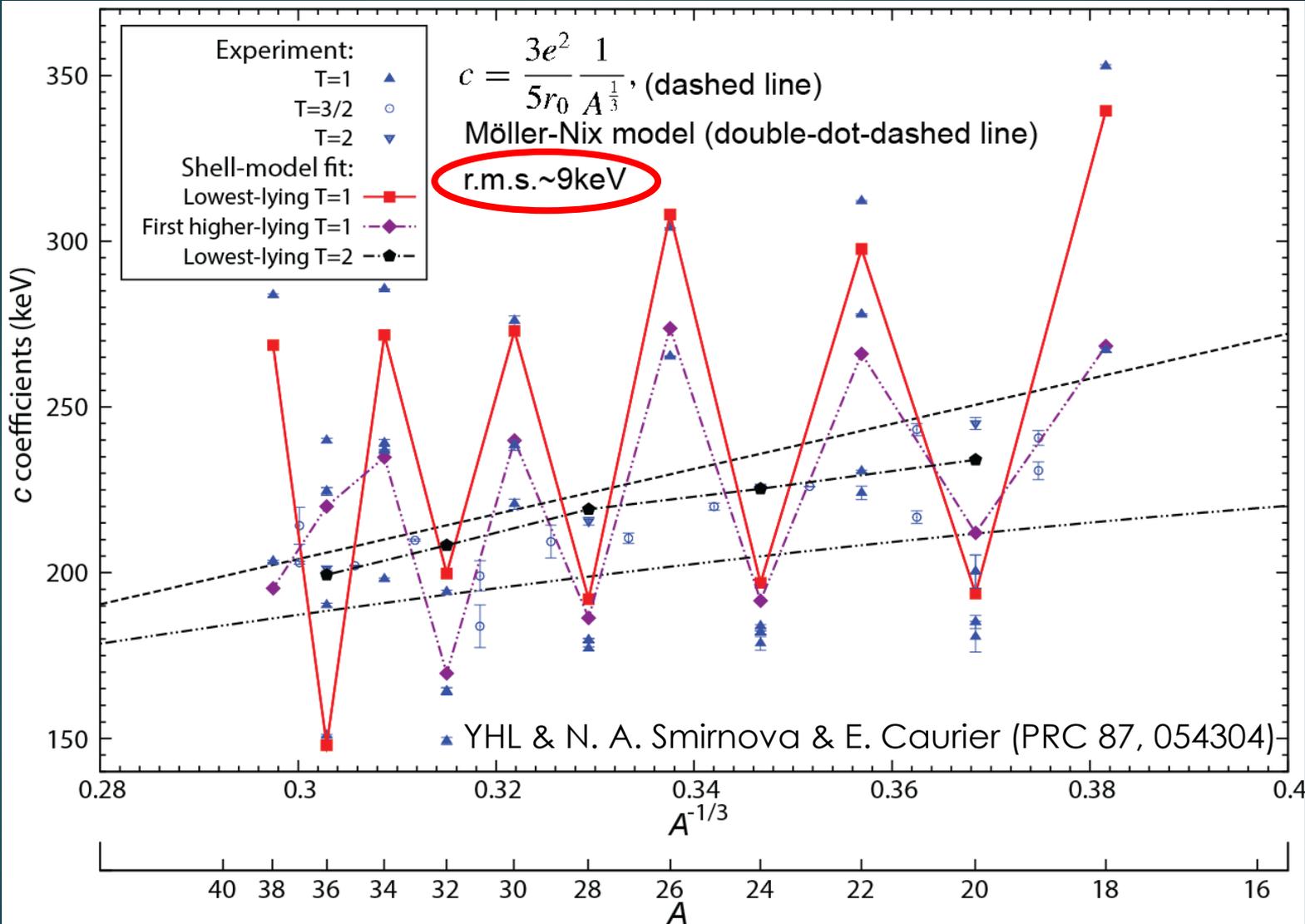
Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian



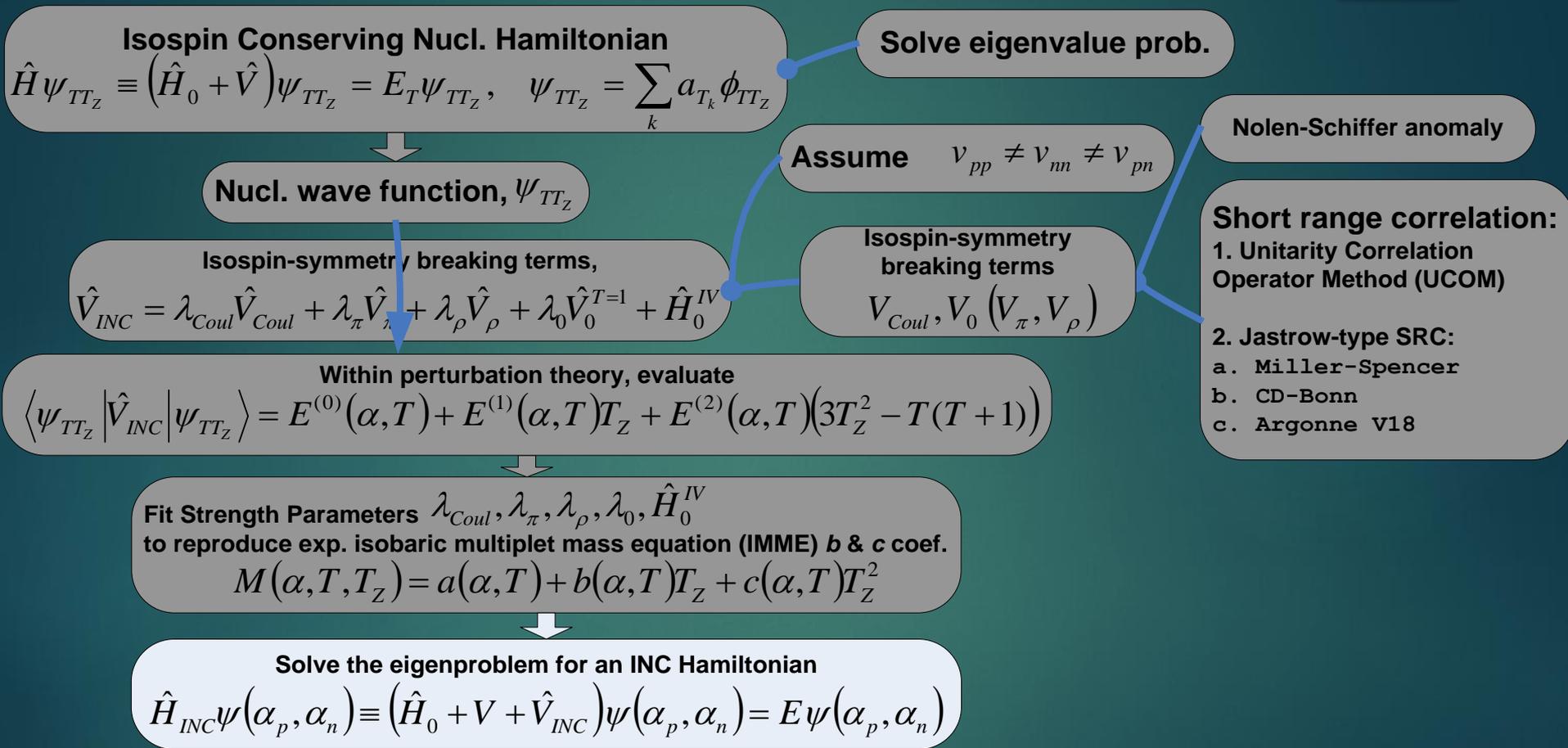
Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian



Shell-model isospin non-conserving (INC) Hamiltonian

Construction of INC Hamiltonian



YHL & N. A. Smirnova & E. Caurier (PRC 87, 054304)

The β -delayed proton emission of proton-rich *sd*- and *pf*-shell nuclei

Isospin symmetry & Isospin admixed states



Isospin symmetry states,

$$|\psi_m^A(JT = 3/2)\rangle = \sum_i \alpha_i |\phi_{n,l,j,T=3/2}\rangle_i$$

Isospin admixed states,

$$|\psi_m^A(JT = 3/2)\rangle = \sum_i \alpha_i |\phi_{n,l,j,T=3/2}\rangle_i + \sum_j \beta_j |\phi_{n,l,j,T=1/2}\rangle_j + \sum_k \chi_k |\phi_{n,l,j,T=5/2}\rangle_k + \dots$$

Isobaric analogue state of ^{53}Co with isospin symmetry,

$$|7/2^-, \text{IAS}_{T=3/2}\rangle = \sum_i \alpha_i |7/2^-, T = 3/2\rangle_i$$

Isobaric analogue state of ^{53}Co with admixed isospin,

$$|7/2^-, \text{IAS}_{T=3/2}\rangle = \sum_i \alpha_i |7/2^-, T = 3/2\rangle_i + \sum_j \beta_j |7/2^-, T = 1/2\rangle_j + \sum_k \chi_k |7/2^-, T = 5/2\rangle_k + \dots$$

Isospin non-conserving Hamiltonian



***sd*-shell nuclei**

$$\text{cd - USDB} = \text{USDB} + V_{Coul}(\text{UCOM}) + V_0(\text{USDB}, T = 1) + \text{isovector SPE}$$

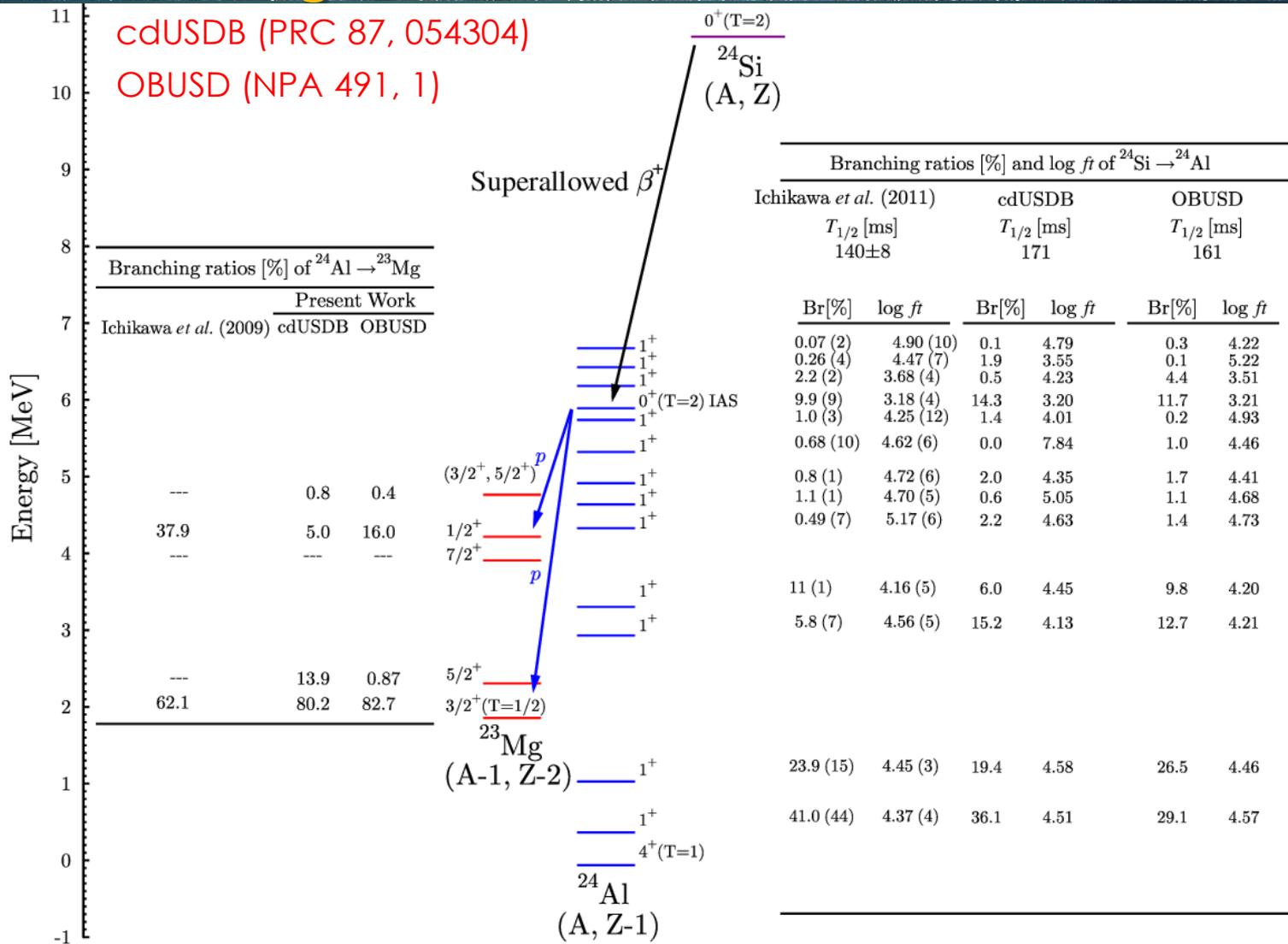
***pf*-shell nuclei (unoptimized INC)**

$$\text{cd - KB3G} = \text{KB3G} + V_{Coul} + \text{isovector SPE}$$

$$\text{cd - GXPF1a} = \text{GXPF1a} + V_{Coul} + \text{isovector SPE}$$

YHL & N. A. Smirnova & E. Caurier (PRC 87, 054304); Ormand & Brown (NPA 491, 1)

β -delayed proton emissions



β -delayed proton emissions



TBME $\langle ij|V|kl\rangle_{JT}^{pp}$

$$\langle (1s_{1/2})^2 | V | (1s_{1/2})^2 \rangle_{01}^{pp}$$

+20%

$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{31}^{pp}$$

-10%

$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{21}^{pp}$$

-15%

$$\langle (0d_{5/2})^2 | V | (1s_{1/2})^2 \rangle_{01}^{pp}$$

-18%

$$\langle 1s_{1/2}0d_{5/2} | V | (0d_{5/2})^2 \rangle_{21}^{pp}$$

-8%

TBME $\langle ij|V|kl\rangle_{JT}^{pn}$

$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{31}^{pn}$$

-10%

$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{21}^{pn}$$

-15%

$$\langle 1s_{1/2}0d_{5/2} | V | (0d_{5/2})^2 \rangle_{21}^{pn}$$

-8%

$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{30}^{pn}$$

-8%

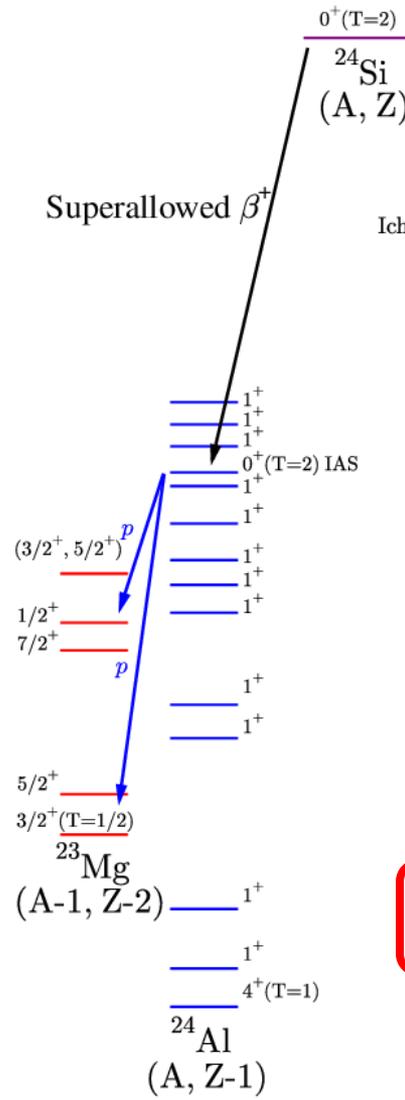
$$\langle 1s_{1/2}0d_{5/2} | V | 1s_{1/2}0d_{5/2} \rangle_{20}^{pn}$$

+78%

$$\langle 1s_{1/2}0d_{5/2} | V | (0d_{5/2})^2 \rangle_{30}^{pn}$$

+2%

USDB:USD



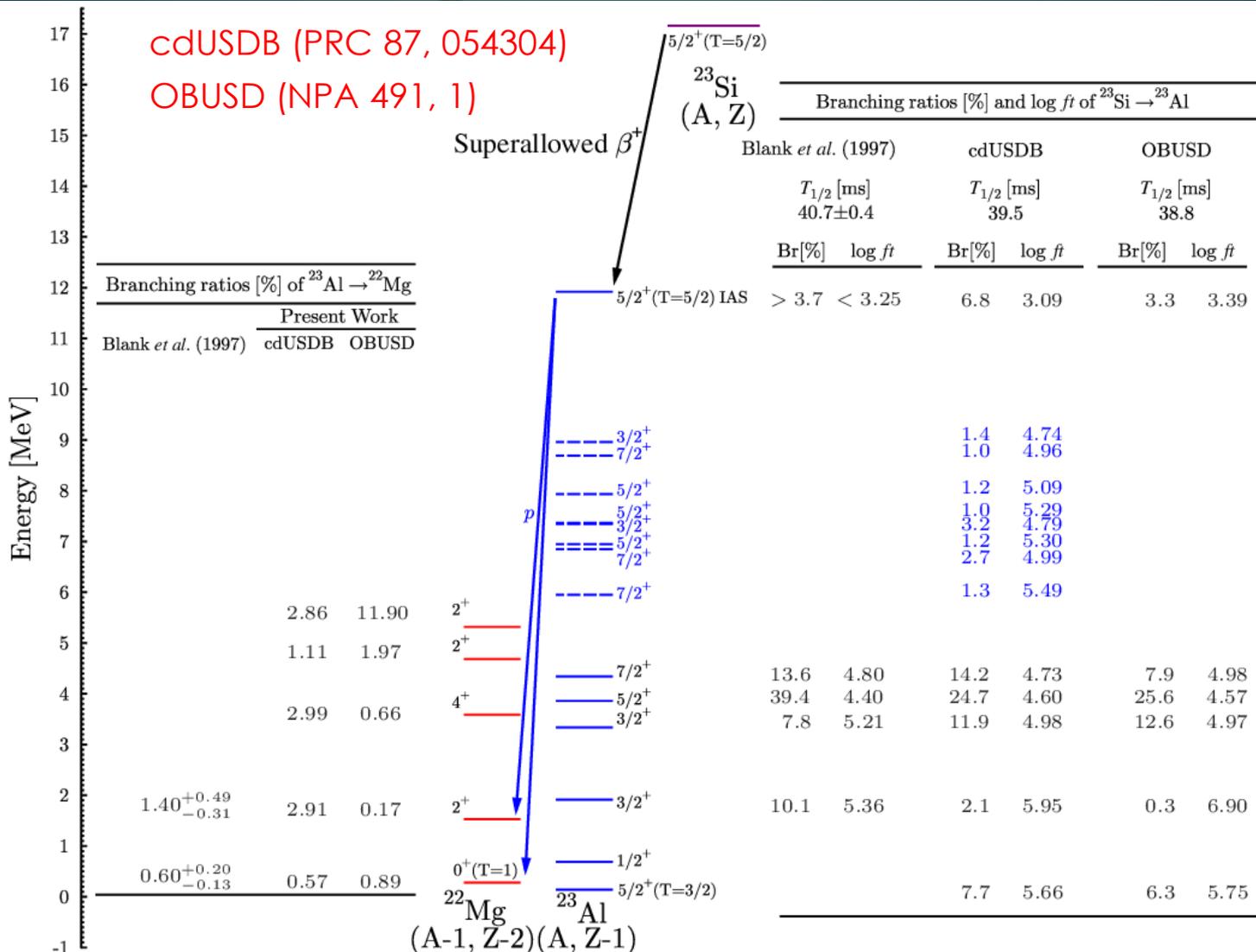
Branching ratios [%] and log ft of ${}^{24}\text{Si} \rightarrow {}^{24}\text{Al}$					
Ichikawa <i>et al.</i> (2011)		cdUSDB		OBUSD	
$T_{1/2}$ [ms]		$T_{1/2}$ [ms]		$T_{1/2}$ [ms]	
140±8		171		161	
Br[%]	log ft	Br[%]	log ft	Br[%]	log ft
0.07 (2)	4.90 (10)	0.1	4.79	0.3	4.22
0.26 (4)	4.47 (7)	1.9	3.55	0.1	5.22
2.2 (2)	3.68 (4)	0.5	4.23	4.4	3.51
9.9 (9)	3.18 (4)	14.3	3.20	11.7	3.21
1.0 (3)	4.25 (12)	1.4	4.01	0.2	4.93
0.68 (10)	4.62 (6)	0.0	7.84	1.0	4.46
0.8 (1)	4.72 (6)	2.0	4.35	1.7	4.41
1.1 (1)	4.70 (5)	0.6	5.05	1.1	4.68
0.49 (7)	5.17 (6)	2.2	4.63	1.4	4.73
11 (1)	4.16 (5)	6.0	4.45	9.8	4.20
5.8 (7)	4.56 (5)	15.2	4.13	12.7	4.21
23.9 (15)	4.45 (3)	19.4	4.58	26.5	4.46
41.0 (44)	4.37 (4)	36.1	4.51	29.1	4.57

β -delayed proton emissions

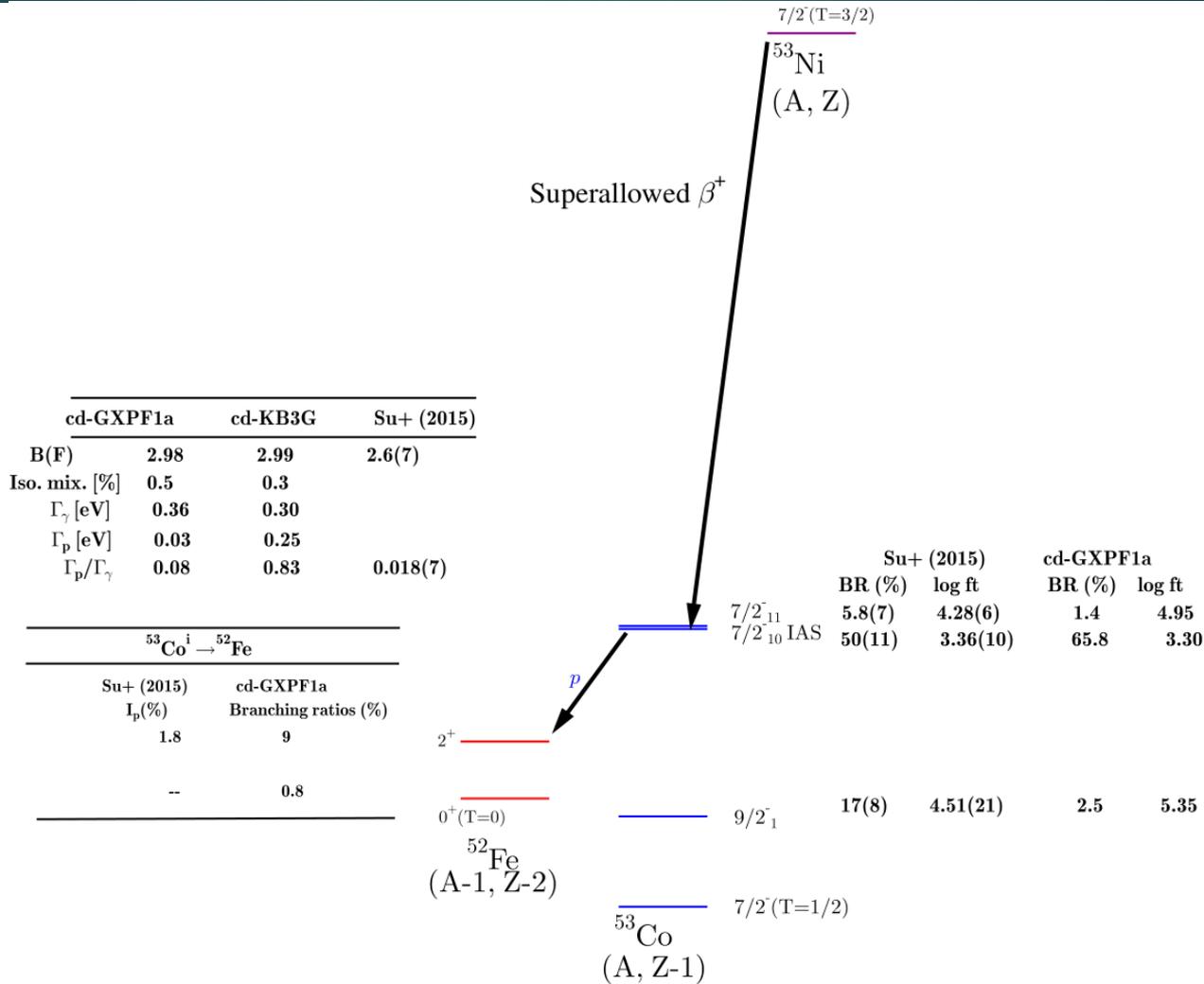


cdUSDB (PRC 87, 054304)

OBUSD (NPA 491, 1)



β -delayed proton emissions



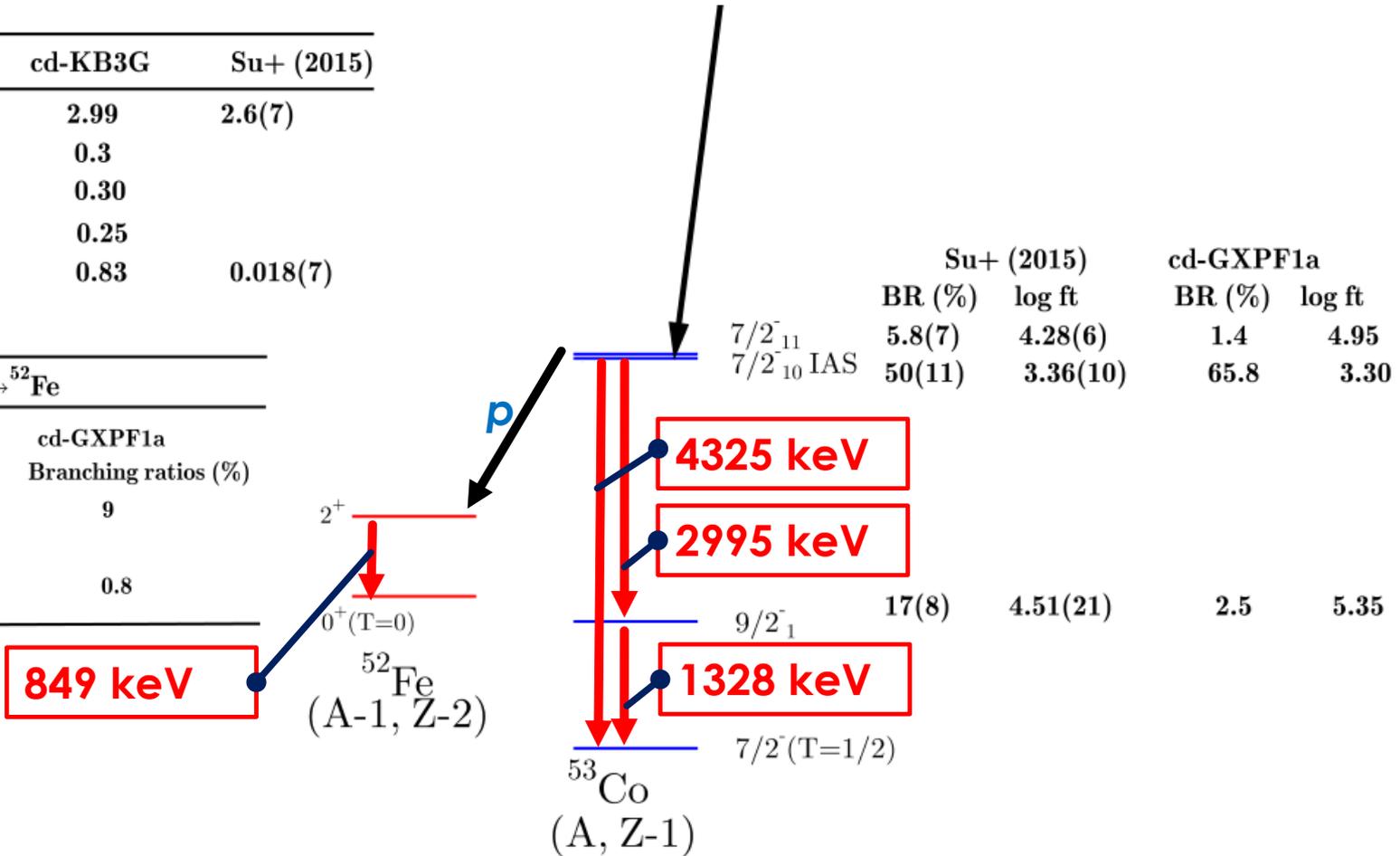
J. Su+ PLB 756, 323

β -delayed proton emissions



	cd-GXPF1a	cd-KB3G	Su+ (2015)
B(F)	2.98	2.99	2.6(7)
Iso. mix. [%]	0.5	0.3	
Γ_γ [eV]	0.36	0.30	
Γ_p [eV]	0.03	0.25	
Γ_p/Γ_γ	0.08	0.83	0.018(7)

${}^{53}\text{Co}^i \rightarrow {}^{52}\text{Fe}$	
Su+ (2015)	cd-GXPF1a
I_p (%)	Branching ratios (%)
1.8	9
--	0.8



YHL & N. A. Smirnova calculated for J. Su+ (PLB 756, 323)

Summary & Perspective



- ▶ Most of the partial decay scheme of *sd*-shell precursors can be described with the newly constructed isospin non-conserving Hamiltonian.
- ▶ Without nuclear-origin isospin symmetry breaking force, we can still describe some partial decay scheme of *pf*-shell precursors. Will inclusion of nuclear-origin isospin symmetry breaking force improve the calculation?
- ▶ USD/USDB is widely used in the study of neutron-rich nuclei. Thomas-Ehrman shift may cause the change of configuration in wave functions, but is not considered in USD/USDB, (and maybe also GXPF1).
- ▶ Evolution of the SPE of $1s_{1/2}$, and of TBME involving $1s_{1/2}$, and $1s_{1/2}$ - $0d_{5/2}$ orbits are not considered for the proton-rich side.

Acknowledgments



COLLABORATORS:

CENBG: Betram Blank;

MSU : B. Alex Brown;

CIAE: Jun SU, WeiPing LIU, Xin Xing XU, Cheng Jian LIN, Li Jie SUN, et al.

IMP: Xing XU, Peng ZHANG, Peng SHUAI, Yu Hu ZHANG, Meng WANG, et al.

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Thank you...