



Direct (α,p) Reaction Measurements with HELIOS and the study of $^{20}\text{Ne}(\alpha,p)^{23}\text{Na}$

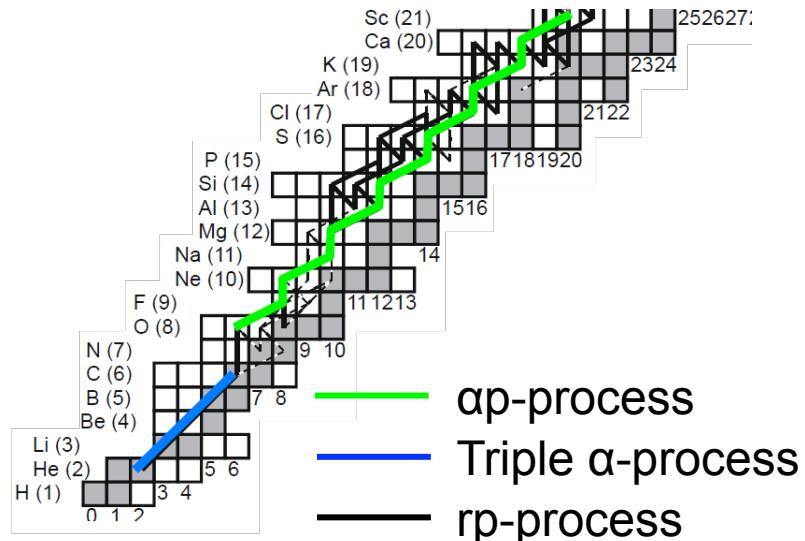
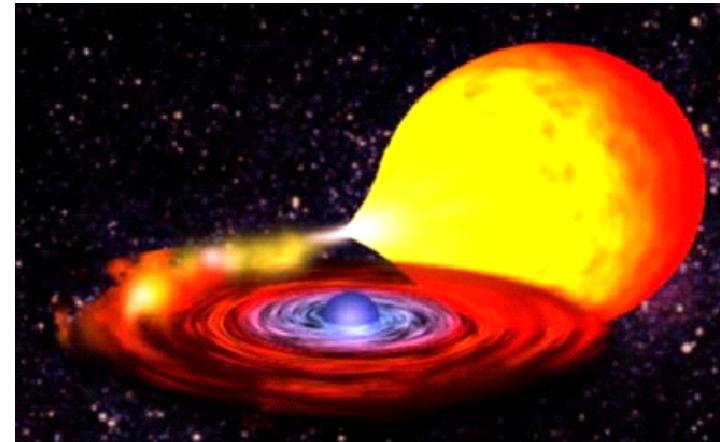
Jianping Lai, Daniel Santiago-Gonzalez

Catherine M. Deibel

Louisiana State University

(α,p) Reactions in Explosive Nucleosynthesis - XRBs

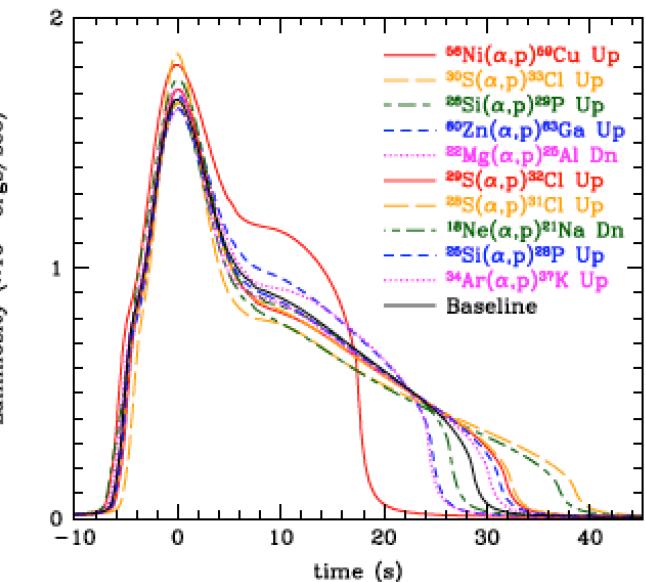
- Type I X-ray bursts (XRBs):
 - thermonuclear explosions on neutron star surface in binary system
 - $T_{\text{peak}} = 1 - 2 \text{ GK}$
 - time scale 10 – 100 s
 - recurrent events (hours to days)
- Nucleosynthesis:
 - triple- α process
 - CNO breakout – $^{18}\text{Ne}(\alpha,p)^{21}\text{Na}$
 - (α,p) process
 - rapid proton capture (rp) process
 - synthesis up to $A \sim 100$



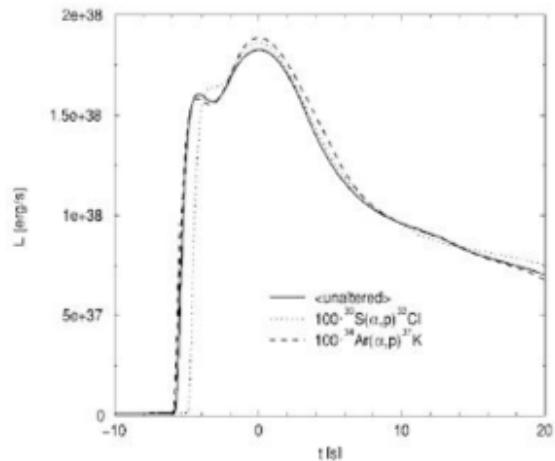
H. Grawe et al., "Rep. Prog. Phys." 70 1525 (2007).

(α,p) Reactions in Explosive Nucleosynthesis - XRBs

- Sensitivity studies of XRB nucleosynthesis:
 - vary reactions individually
 - show only a handful of reactions significantly effect bursts
- Effects of (α,p) reactions in XRBs:
 - shape of light curve
 - energy output
 - elemental abundances
 - double-peaked bursts(?)



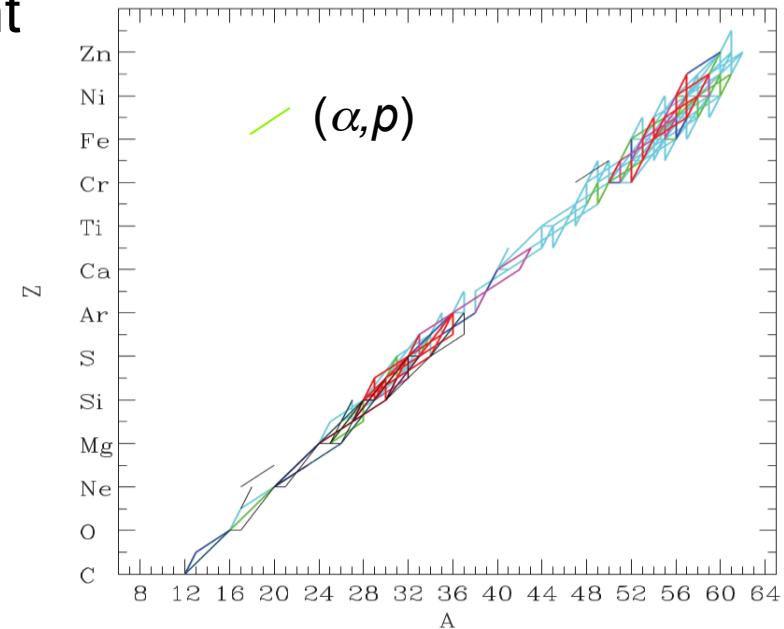
Cyburt *et al.*, 2016 (submitted)



J.L. Fisker, *et al.*, ApJ (2004).

(α,p) Reactions in Explosive Nucleosynthesis - SNela

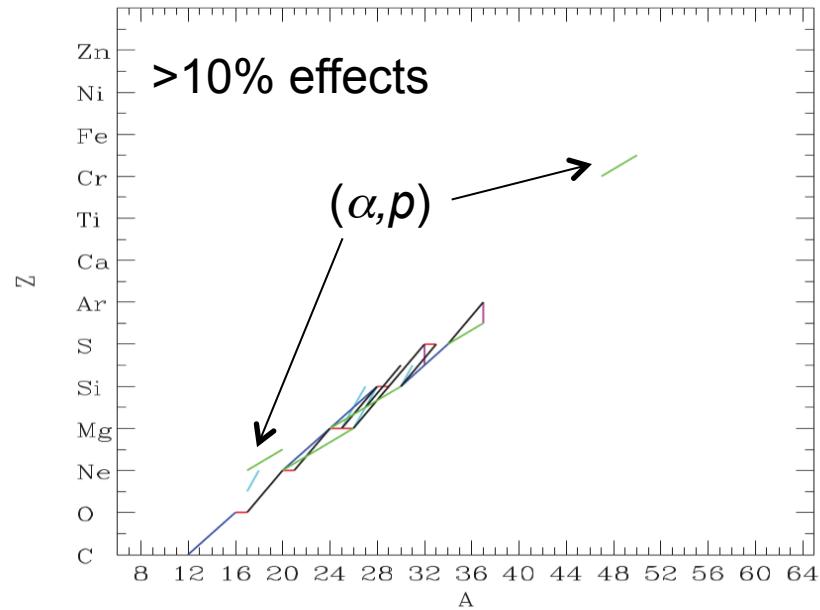
- Type Ia Supernovae:
 - thermonuclear explosion of white dwarf star in binary system
 - progenitor uncertain
 - disruptive explosion; no remnant
 - $T_{\text{peak}} \sim 8 \text{ GK}$
- Nucleosynthesis in SNela:
 - C+C, C+O, O+O fusion
 - α -chain reactions
 - production of >50% Fe content of Galaxy



Bravo *et al.*, PRC **85**, 055805 (2012)

Type Ia Supernovae Sensitivity Studies

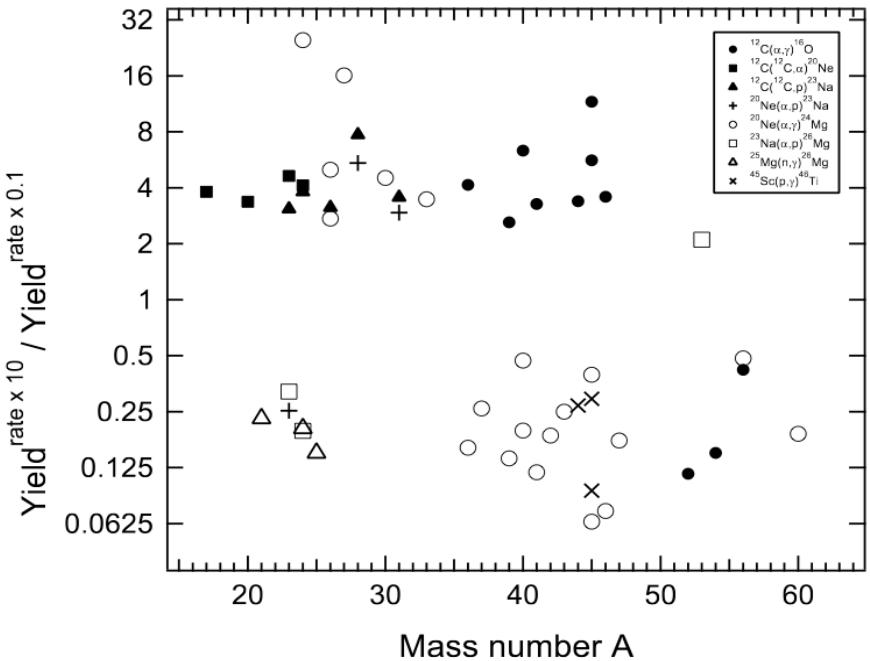
- Variation of reaction rates by factor of 10 up and down
- Two independent studies show high sensitivity to:
 - $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$
 - $^{12}\text{C} + ^{12}\text{C}$
 - $^{20}\text{Ne}(\alpha, p)^{23}\text{Na}$
 - $^{20}\text{Ne}(\alpha, \gamma)^{24}\text{Mg}$
 - $^{30}\text{Si}(p, \gamma)^{31}\text{P}$



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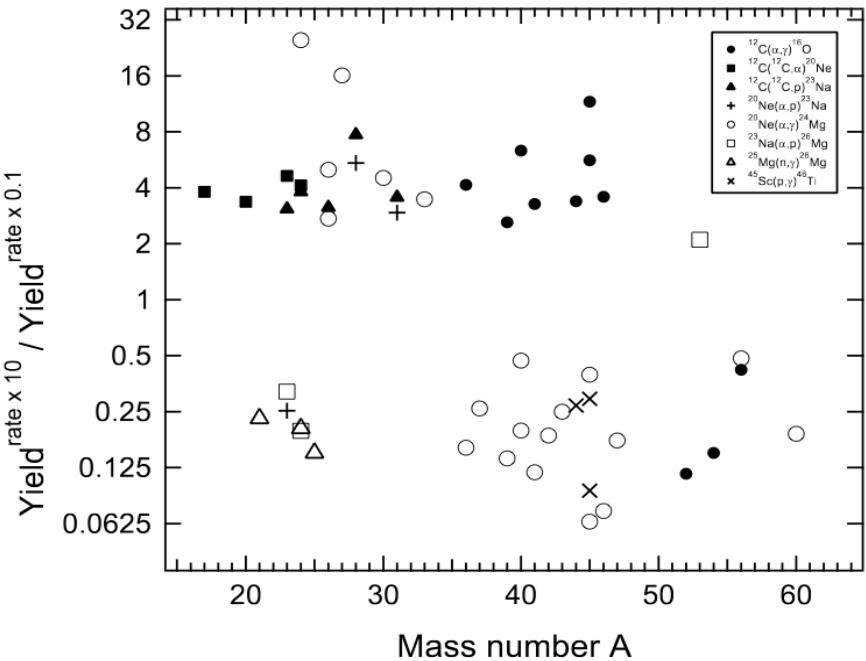


Reaction	Nuclide	W7 model		DDT model	
		×10	×0.1	×10	×0.1
$^{20}\text{Ne}(\alpha, p)^{23}\text{Na}$	^{18}O	0.44	2.4		
	^{23}Na	0.47	2.2	0.48	
	^{26}Al			2.1	
	^{28}Mg			5.4	
	^{31}Si			3.0	

A. Parikh *et al.*, A&A **557**, 11 (2013)

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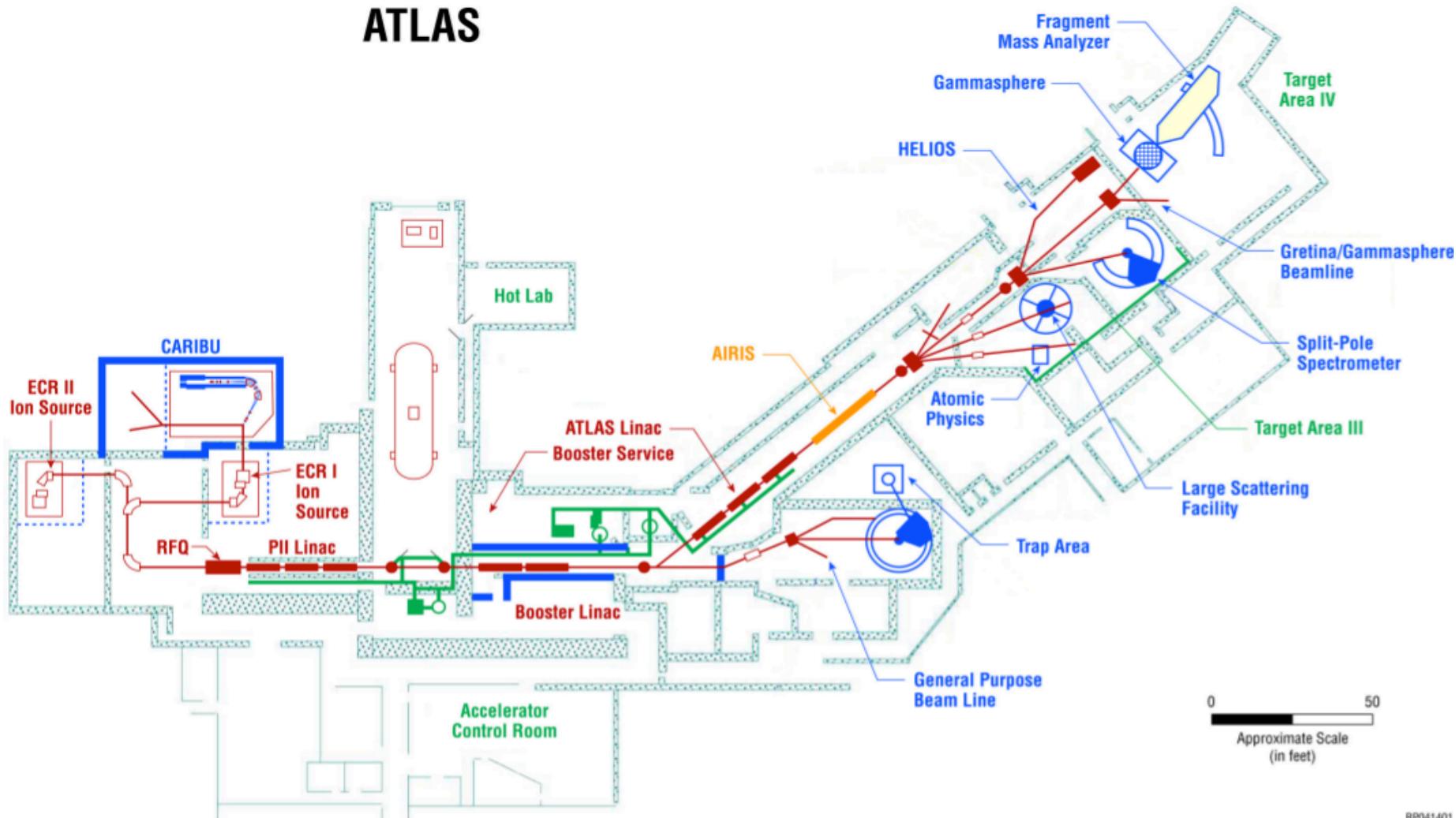
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Direct (α, p) measurements with HELIOS at Argonne National Laboratory

ATLAS

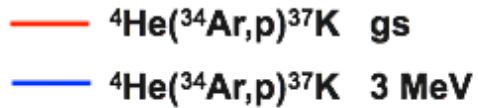
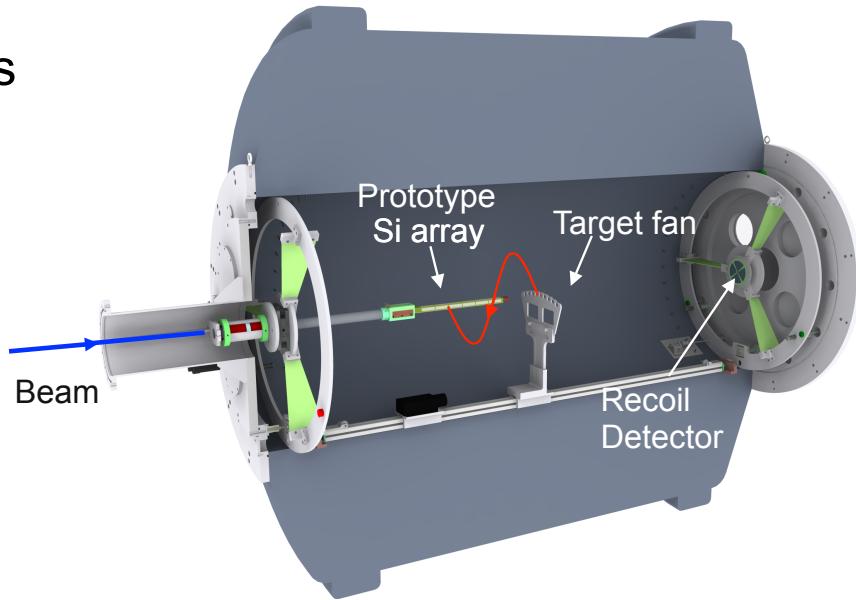


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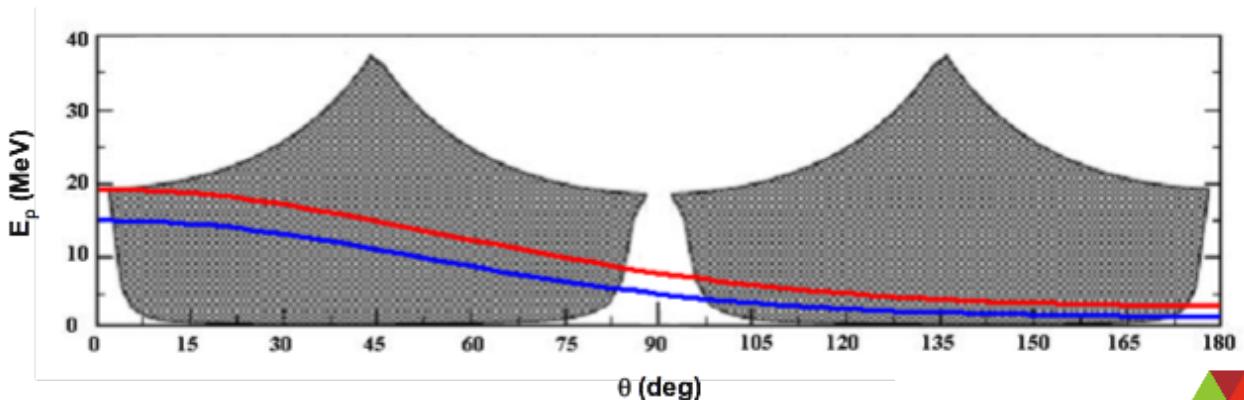
Direct (α, p) measurements with HELIOS



- HELical Orbit Spectrometer (HELIOS)
 - repurposed MRI magnet
 - magnetic field aligned with beam axis
 - commissioned 2008
- Ideal for studying reactions with radioactive ion beams
 - unique particle identification from time-of-flight
 - high geometrical efficiency
 - improve resolution (avoid kinematic compression)



Particle	p	${}^3\text{He}$	d, ${}^4\text{He}$	t
TOF(ns)	21.9	32.8	43.7	65.6



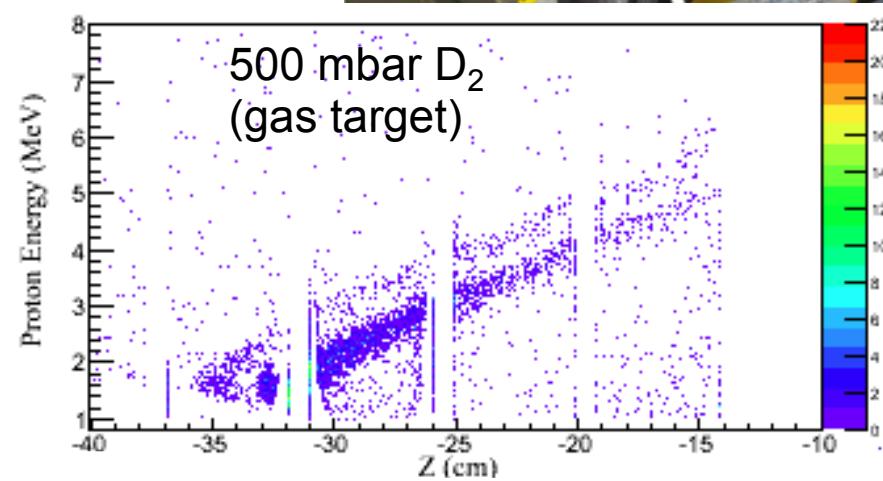
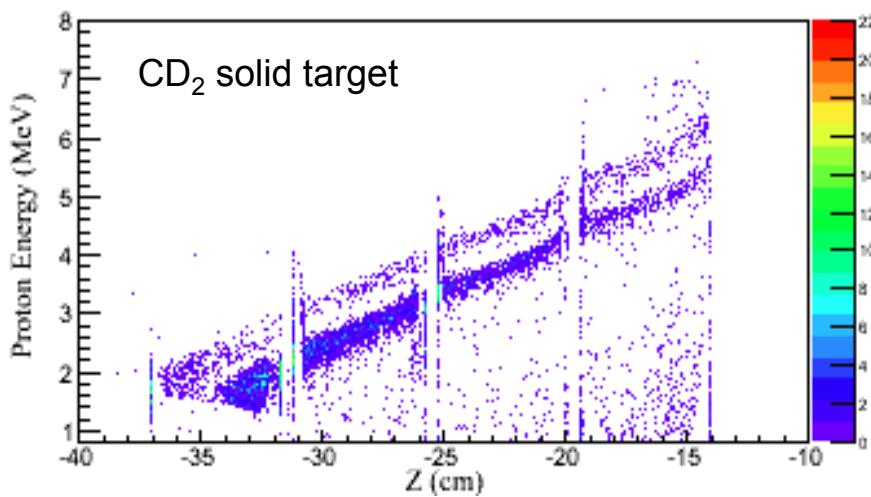
Direct (α,p) measurements with HELIOS

- Direct (α,p) measurement needs:
 - ${}^4\text{He}$ gas target
 - cryogenically cooled
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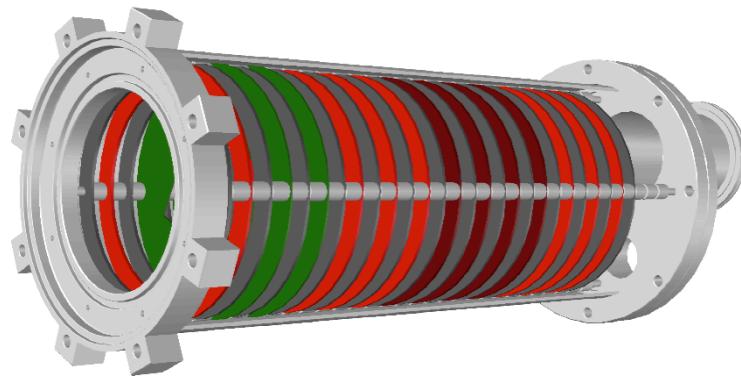
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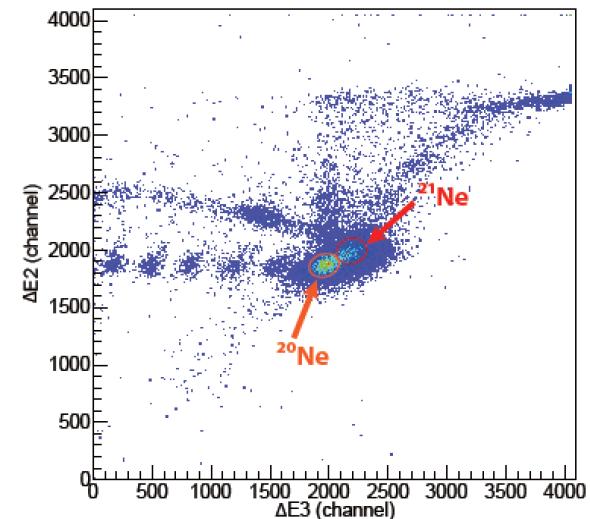
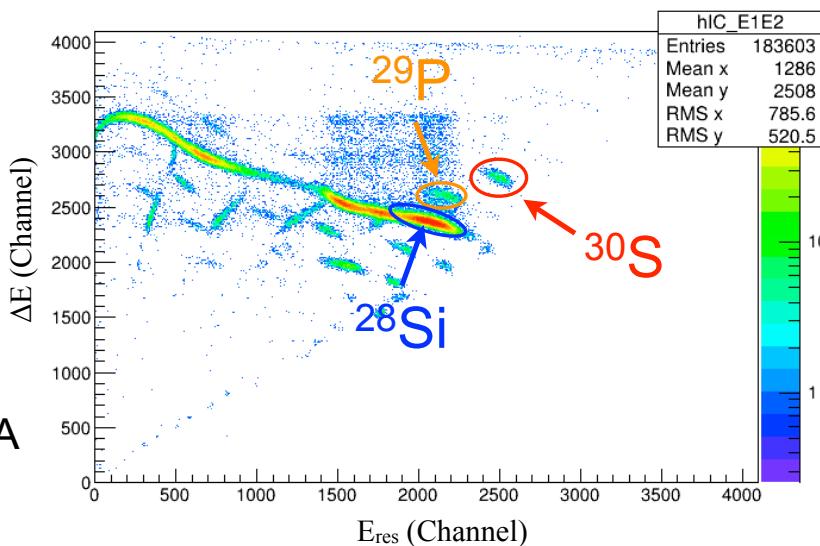
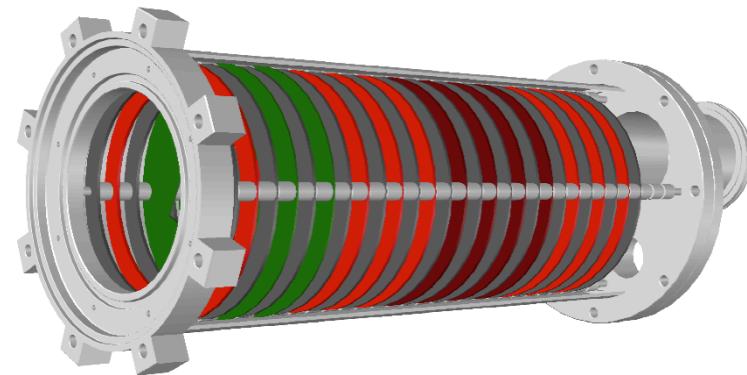
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 - recoil identification
 - beam monitoring/normalization
 - position information



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high rate: 500 kHz

energy resolution: <5%

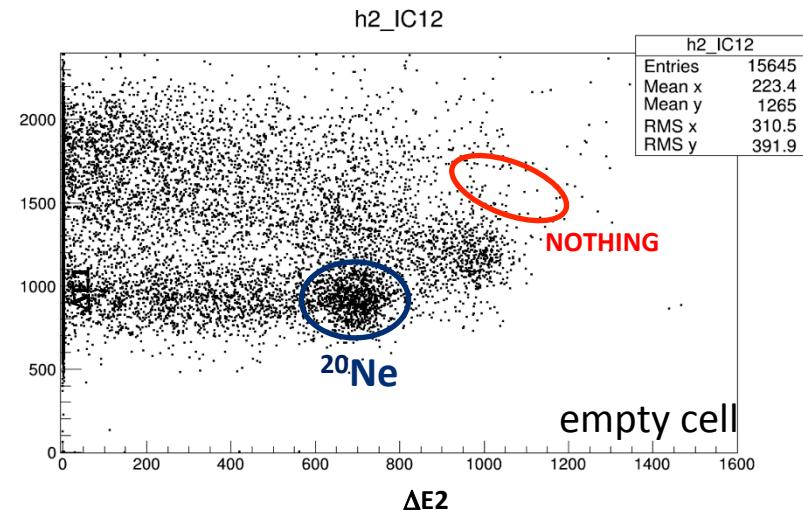
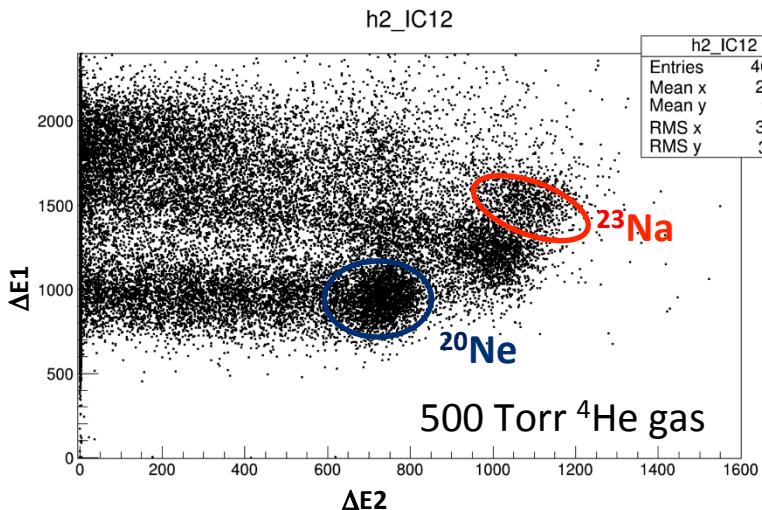
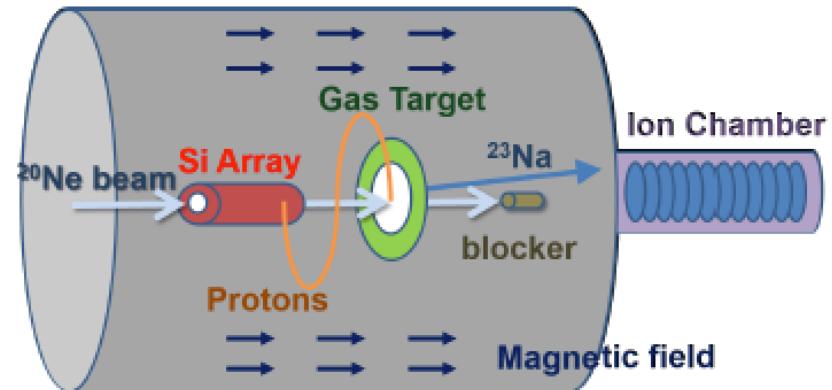
position sensitivity

mass separation at low A

First Direct (α,p) Study with HELIOS: $^{20}\text{Ne}(\alpha,p)^{23}\text{Na}$



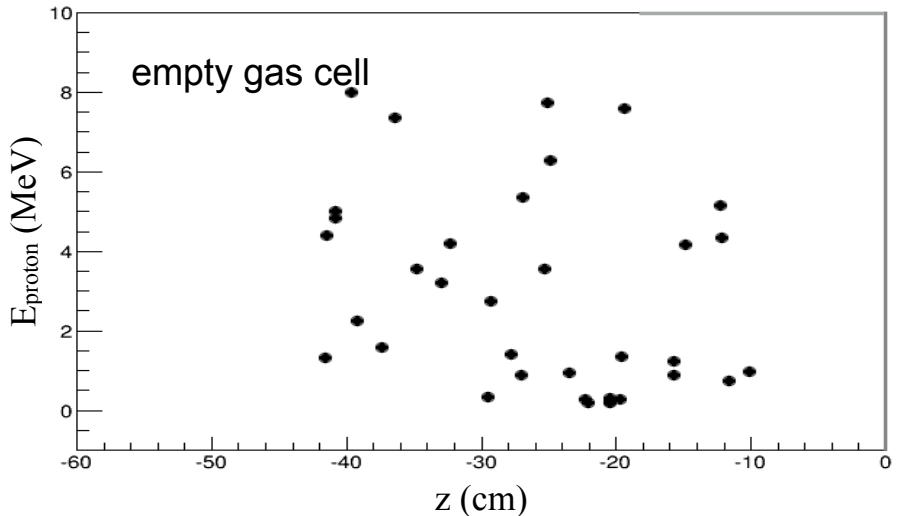
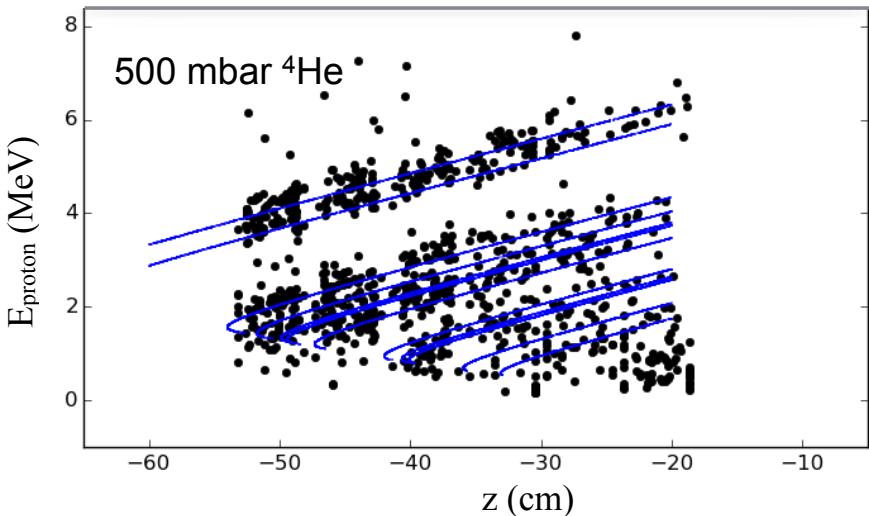
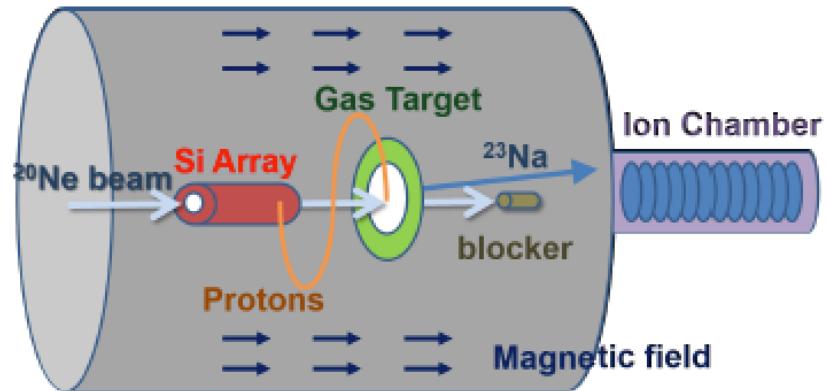
- Direct study of $^{20}\text{Ne}(\alpha,p)^{23}\text{Na}$
 - protons detected in HELIOS Si array
 - ^{23}Na detected in ionization chamber
 - normalized via known (d,p) reaction
- Part I completed December 2014
 - $E_{\text{beam}} = 107, 100, 88 \text{ MeV}$



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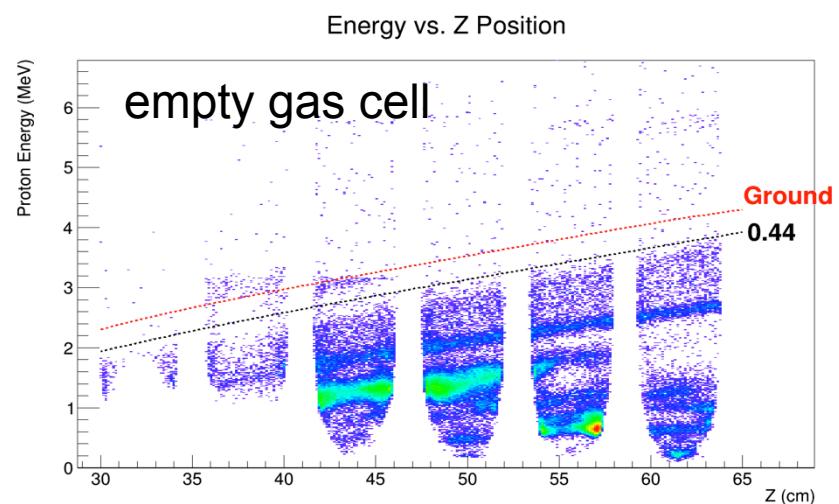
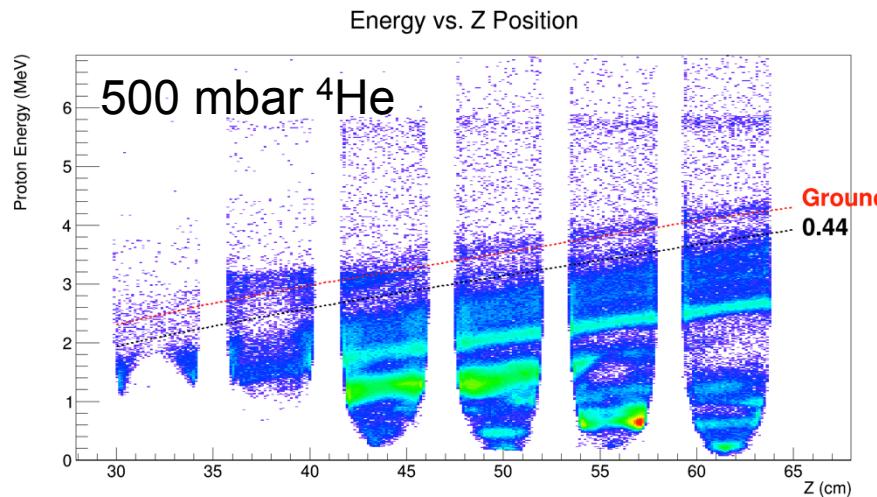
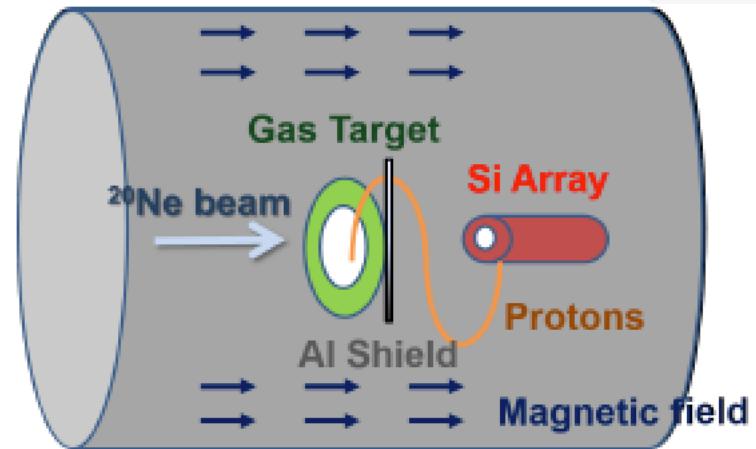
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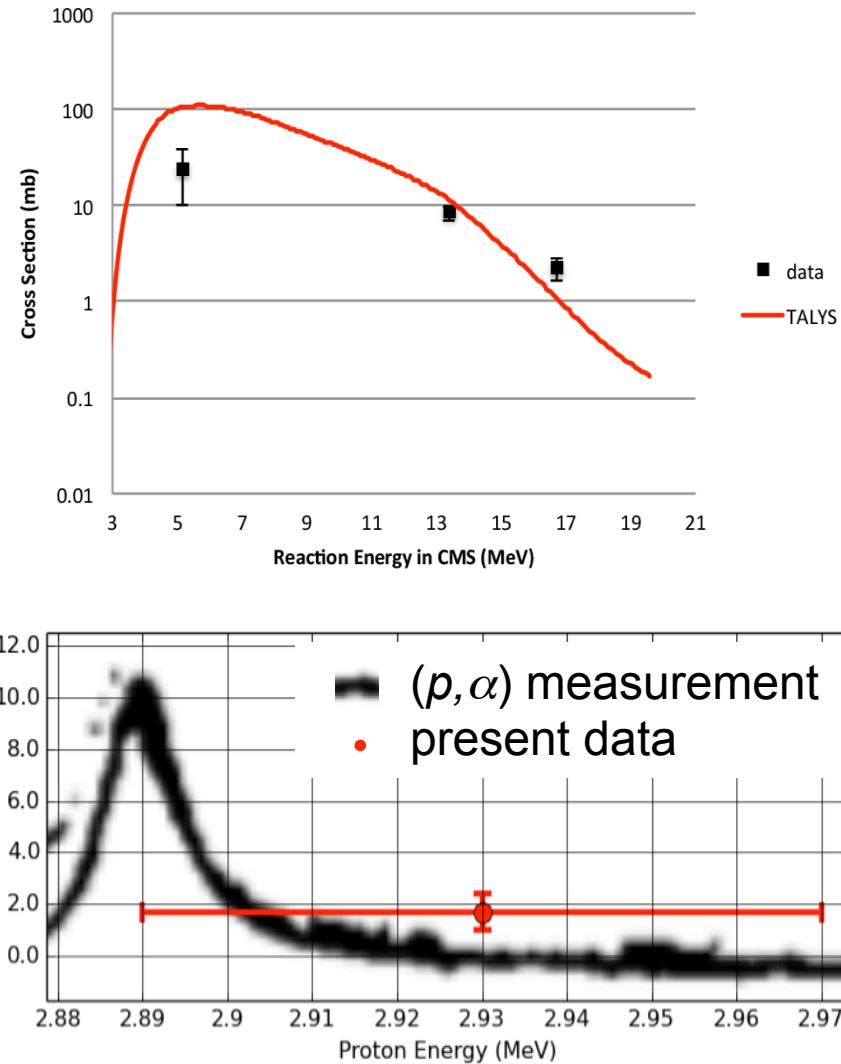


- Part II completed July 2015
- Normalization via (α,α) scattering and (d,p) on solid target
- Dissertation of Jianping Lai (LSU)
– August 2016



$^{20}\text{Ne}(\alpha, p)^{23}\text{Na}$ Cross Sections

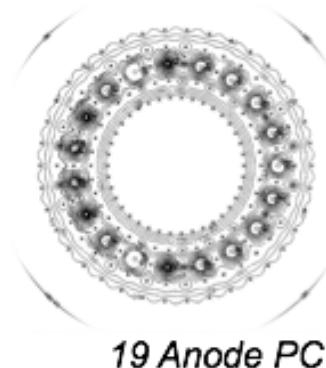
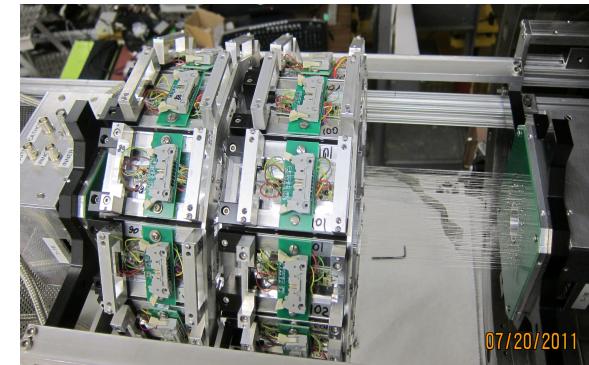
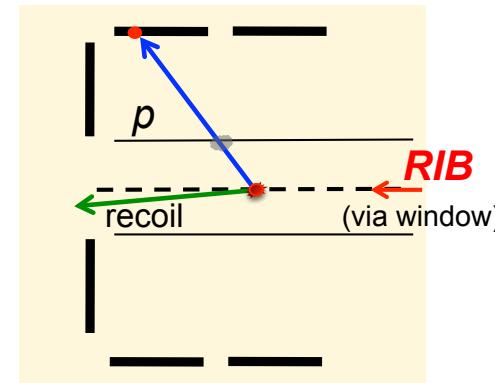
- Normalized cross section for $^{20}\text{Ne}(\alpha, p)^{23}\text{Na}$:
 - in agreement with Hauser-Feshbach calculations
 - in agreement with inverse reaction cross section
- Future plans:
 - determining contribution from excited states (present data)
 - measurements at more (and lower) energies
 - study affects on SNela models





Current Developments: ANASEN

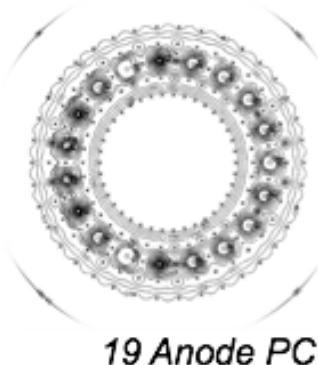
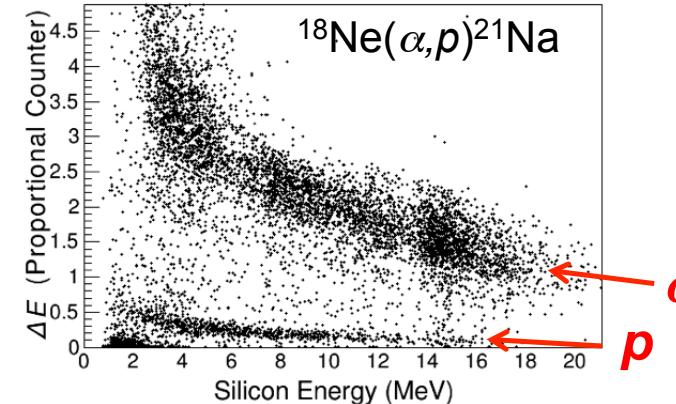
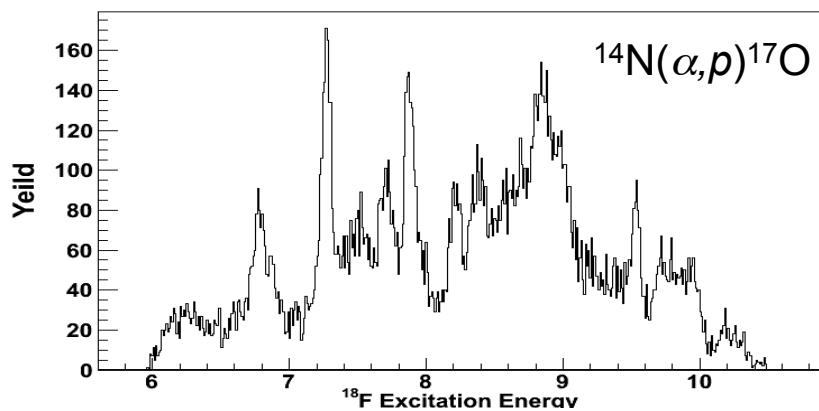
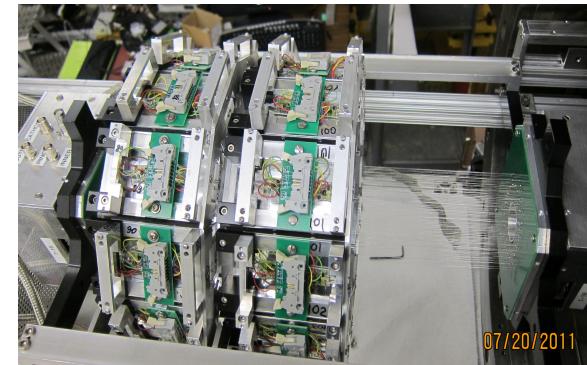
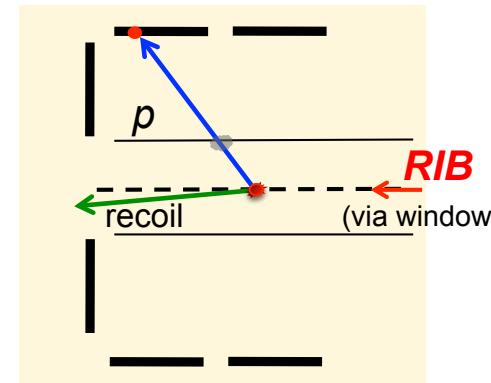
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- Nuclear Astrophysics measurements:
 - $^{14}\text{N}(\alpha, p)^{17}\text{O}$ (stable beam FSU)
 - $^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$ (RIB from RESOLUT @ FSU)
 - Kevin Macon, PhD Thesis (LSU)
 - $^{37}\text{K}(p, p)^{37}\text{K}$ (first RIB measurement @ ReA3)



Summary

- (α,p) reactions play a significant role in stellar explosive nucleosynthesis
- Direct measurements are challenging:
 - radioactive ion beams
 - gas targets
 - low cross sections
- Multiple methods underdevelopment for direct (α,p) measurements:
 - HELical Orbit Spectrometer (HELIOS) at ATLAS facility
 - ANASEN
 - JENSA gas-jet target with Si detector array
- First direct measurement of $^{20}\text{Ne}(\alpha,p)^{23}\text{Na}$:
 - proof-of-principle of HELIOS method
 - important for Type Ia Supernova nucleosynthesis

THANKS!

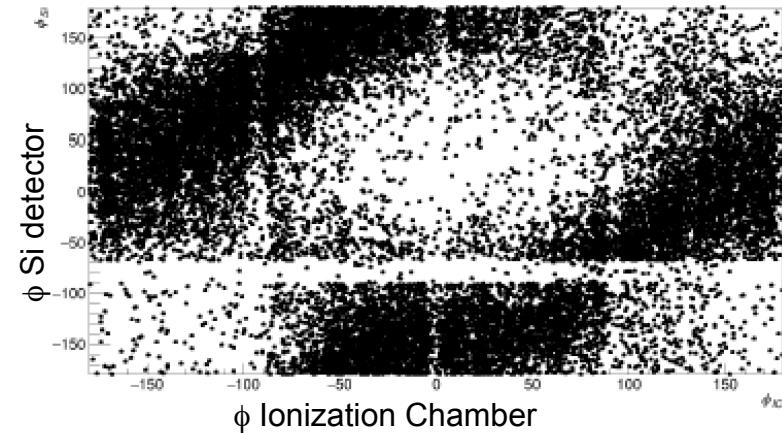
- ANL
 - Jason Clark
 - John Greene
 - Calem Hofman
 - Ben Kay
 - Richard Pardo
 - Ernst Rehm
 - Birger Back
 - Akaa Ayangeakaa
 - Melina Avila
- Louisiana State University
 - Jianping Lai
 - Jeff Blackmon
 - Daniel Santiago-Gonzalez
 - Kevin Macon (Postdoc – Notre Dame)
 - Amber Lauer (graduate student)
 - Liudmyla Afanasieva (Postdoc - WashU)
- Florida State University
 - Sergio Almaraz



****ATLAS staff and operators****

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