

Exploring Nuclear Shell Evolution with Nucleon Transfer Reactions

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and the TIARA and SHARC/TIGRESS collaborations

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Experiments performed at TRIUMF and at Texas A&M University, USA



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Adelaide, 1977

EXPLORING SHELL EVOLUTION with nucleon transfer reactions specifically, via (d,pγ) neutron transfer

(a) motivation: changing shell structure
(b) experiments: SHARC + TIGRESS + trifoil
(c) update on recent PLB ²⁵Na,pγ)²⁶Na
(d) new results from d(²⁴Na,pγ)²⁵Na
(e) new results from d(²⁸Mg,pγ)²⁹Mg
(f) perspectives: T-Rex at Texas A&M

Adelaide Oval close to 1977

²⁷ P	²⁸ P	²⁹ P	³⁰ P	³¹ P	³² P	³³ P	³⁴ P	³⁵ P	³⁶ P	³⁷ P	³⁸ P
²⁶ Si	²⁷ Si	²⁸ Si	²⁹ Si	³⁰ Si	³¹ Si	³² Si	³³ Si	³⁴ Si	³⁵ Si	³⁶ Si	³⁷ Si
²⁵ AI	²⁶ AI	²⁷ AI	²⁸ AI	²⁹ AI	³⁰ AI	³¹ AI	³² AI	³³ AI	³⁴ AI	³⁵ AI	³⁶ AI
²⁴ Mg	²⁵ Mg	^{₂6} Mg	²⁷ Mg	²⁸ Mg	²⁹ Mg	³⁰ Mg	³¹ Mg	³² Mg	³³ Mg	³⁴ Mg	³⁵ Mg
²³ Na	²⁴ Na	²⁵ Na	²⁶ Na	²⁷ Na	²⁸ Na	²⁹ Na	³⁰ Na	³¹ Na	³² Na	³³ Na	³⁴ Na
²² Ne	²³ Ne	²⁴ Ne	²⁵ Ne	²⁶ Ne	²⁷ Ne	²⁸ Ne	²⁹ Ne	³⁰ Ne	³¹ Ne	³² Ne	³³ Ne
²¹ F	²² F	²³ F	²⁴ F	²⁵ F	²⁶ F	²⁷ F	²⁸ F	²⁹ F	³⁰ F	³¹ F	

orange – nuclei studied by us, using (d,p) green – (a) N=20, (b) island of inversion (intruder structure dominates ground state structure)



G.L. Wilson et al.

Theory: effective spe's

Experiment: energies of just the lowest levels



- Our aim is to identify single-particle-like levels and determine their spin/parity
- We use the selective nature of (d,p) neutron transfer (with radioactive beams)
- We aim to track the evolution of these levels and compare to the shell model



 $S = |J\downarrow SP\uparrow \pi J\downarrow i\uparrow \pi |\uparrow 2$

spectroscopic factor
= overlap with pure SP state



 $| J\downarrow i\uparrow \pi > = \sqrt{S} | J\downarrow SP\uparrow \pi > + \sum k\uparrow m \alpha \downarrow k | J\downarrow k\uparrow \pi >$

- we measure transferred $\ell \downarrow n$
- we measure gamma-decays
- we aim to identify J and π
- we deduce S



 $S=|J\downarrow SP\uparrow\pi J\downarrow i\uparrow\pi |\uparrow 2$

spectroscopic factor
= overlap with pure SP state

Theory: effective spe's

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- We aim to track the evolution of these levels and compare to the shell model
- We use the large SF for theory/experimental states to associate them with each other
- Details of the precise numerical value of the SF don't affect this process
- Results will be shown here for Z=11,12 for N=14,15,17, probing higher orbitals



can make beams of VERY exotic nuclei and learn properties by removing neutrons **OR**

can learn the important interactions that explain the structure by isolating p-n interactions, using a **single nucleon** to probe the additional orbitals one at a time

thus **transfer** is an excellent way to isolate the separate interactions





G.L. Wilson et al., Physics Letters B 759 (2016) 417

Gemma Wilson, Surrey Proton Ex (keV) 0002 0009 0009 cascade decays eround state decan

Data from d(²⁵Na,p)²⁶Na at 5 MeV/A using SHARC at ISAC2 at TRIUMF

Ε_γ (keV)

Doppler corrected (β =0.10) gamma ray energy measured in TIGRESS

Excitation energy deduced from proton energy and angle

Experimental Results from studying d(²⁵Na,p)²⁶Na at TRIUMF



Differential cross sections and spectroscopic factors

First analysis of this type:

Each of these distributions is:

- (a) gated on a gamma-ray peak
- (b) background-subtracted
- (c) corrected for gamma ray efficiency
- (d) corrected for gamma ray branching ratio



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Experimental Results from studying d(²⁵Na,p)²⁶Na at TRIUMF





comparison between revised shell model energies and SFs

the results are somewhat subtle

evidence for stronger influence of the 1p3/2 orbital in the low-lying negative parity states, compared to the less exotic isotone ²⁸Al

this is evidence for the 1p3/2 orbital becoming lower, relative to the 0f7/2 orbital which is clear, in ²⁷Ne and ²⁹Mg

the shell model works surprisingly well wbc spsdpf 0+1ħω

Experimental Results from studying d(²⁵Na,p)²⁶Na at TRIUMF

						single L analysis				two L analysis (where applicable)							
	No.	$\mathbf{E}_x^{(a)}$	$\mathbf{E}_x^{SM \ b)}$	J ^{π c)}	\mathbf{J}_{SM}^{π}	L	nlj	S	S^{SM}	${\rm L}_1$	$n_1l_1j_1$	S_1	S_1^{SM}	L_2	$n_2 l_2 j_2$	S_2	S_2^{SM}
UPDATE		0	0	3^{+}	3^{+}_{1}	*	$1s_{1/2}$		0.61	*	$1s_{1/2}$		0.61	*	$0d_{3/2}$		0.01
<u></u>		a anad													$0d_{5/2}$		0.01
		0.082^{a})	0.077	1+	1_{1}^{+}	*	$0d_{3/2}$		0.29								
8 new states		0.000	0.140	o+	0 +	0	$0d_{5/2}$	0.19	0.11	0	1	0.10	0.15	0	0.1	0.101	0.10
nlus		0.232	0.149	21	z_1	0	$1s_{1/2}$	0.13	0.15	0	$1s_{1/2}$	0.10	0.15	2	$0d_{3/2}$	0.19Ţ	0.10
		0.405	0.416	9+	9+	0	1	0.33	0.97	0	1	0.30	0.97	9	$0d_{5/2}$	0.134	0.09
4 new ℓ value	es	0.400	0.410	2	42	0	151/2	0.00	0.21	0	151/2	0.00	0.21	4	$0d_{5/2}$	0.10	0.03
		1.507	1.409	1+	1^{+}_{0}	2	0d2/9	0.39	0.09						043/2		0.00
					2		$0d_{5/2}$		0.10								
		1.805	1.676	(3^{+})	3^{+}_{2}	2	$0d_{3/2}$	0.37	0.33	2	$0d_{3/2}$	0.33†	0.33	0	$1s_{1/2}$	0.01‡	0.00
background							$0d_{5/2}$		0.02	2	$0d_{5/2}$		0.02				
subtraction		1.992	1.758	4+	4_{1}^{+}	2	$0d_{3/2}$	0.07	0.07								
		2.116	2.241	5^{+}	5^{+}_{1}	2	$0d_{5/2}$	0.16	0.08								
		2.195	2.142	2+	2^+_3	2	$0d_{3/2}$	0.49	0.06								
NEW		2.225	2.048	(4^{+})	4^+_2	2	$0d_{3/2}$	0.43	0.51								
gamma-rav		0.400	9.459	0+	o+		$0d_{5/2}$		0.01	0	1-	0.00	0.19	0	0.1	0.14	0.99
Barrina ray		2.423	2.402	(2^{-})	$\frac{24}{2^{-}}$	2	1.000		0.20	2	$18_{1/2}$	1 10	0.15	2	$1_{D_{3/2}}$	0.14	0.25
angular		2.040	2.550	(2)	21	0	$1p_{3/2}$		0.20	0	017/2 Ofr. /0	1.10	0.20	1	1P3/2	0.10	0.03
correlations		3.135	3.228	3-	3^{-}_{1}	1	1D2/9	0.07t	0.15	1	1D2/0	0.06 1	0.15	3	$0f_{7/2}$	0.10t	0.13
		0.100	0.220		-1	-	$1p_{1/2}$	0.0.1	0.02	-	1p _{1/2}	0.001	0.02		$0f_{5/2}$	01107	0.00
		3.511	3.513	4^{-}	4^{-}_{1}	1	$1p_{3/2}$	0.30	0.44	1	$1p_{3/2}$	0.25	0.44	3	$0f_{7/2}$	0.51†	0.00
I.C. Celik							- /				- /				$0f_{5/2}$		0.00
PhD thesis		4.087	3.690	2^{-}	2^{-}_{2}	3				1	$1p_{3/2}$	0.34	0.31	3	$0f_{7/2}$	0.78	0.03
Surroy 2015		4.239	3.975	4+	4_{5}^{+}	2	$0d_{3/2}$	0.12	0.12								
Surrey 2015		4.305	4.401	(5^{-})	5^{-}_{1}	3				1	$1p_{3/2}$	0.01	0.00	3	$0f_{7/2}$	0.25	0.46
		4.597	4.460	3-	3_{2}^{-}	3	$0f_{7/2}$			1	$1p_{3/2}$	0.02	0.10	3	$0f_{7/2}$	0.76	0.10
		4.800	4.730	4^{-}	$\frac{4}{c^{-}}$	3	$0f_{7/2}$	0.51	0.61	1	$1p_{3/2}$	0.00	0.05	3	$0f_{7/2}$	0.62	0.37
		4.917	4.001	(0)	0 ₁	0 0	$01_{7/2}$	0.51	0.01	1	1.5	0.00	0.99	9	Of	0.62	0.05
		$\frac{4.952}{5.000}$	4.770	$(3^{-} 4^{-})$	\mathcal{O}_4	د *	017/2			1	$1p_{3/2}$	0.00	0.20	0	017/2	0.03	0.00
		0.009		(0,4)													

Experimental Setup to Measure d(²⁴Na,p)²⁵Na at TRIUMF



d(²⁴Na,p)²⁵Na at 8.0 MeV/u with 10,000 pps



Excitation energy from (E, θ) of proton, MeV

Andy Knapton, Surrey PhD

Doppler corrected E(gamma), MeV

d(²⁴Na,p)²⁵Na at 8.0 MeV/u with 10,000 pps



Excitation energy from (E, θ) of proton, MeV

Andy Knapton, Surrey PhD

d(²⁴Na,p)²⁵Na – fits to excitation energy spectrum at each angle



Excitation Energy in ²⁵Na (MeV)

Andy Knapton, Surrey PhD

d(²⁴Na,p)²⁵Na – spectroscopic factors in ²⁵Na compared to theory



BIG IMPROVEMENTS IN LEVEL IDENTIFICATIONS

Andy Knapton, Surrey PhD

Using the ²⁵Na SFs to calculate ²⁴Al(p, γ)²⁵Si widths and $\omega\gamma$'s for novae



d(²⁸Mg,p)²⁵Na at 8.0 MeV/u with 3,000 pps

Secondary Beam

- ²⁸Mg beam 3000 pps at 8 AMeV → With strong contamination
 - → With strong contamination ²⁸Si cont. (3. 10⁵ pps) ²⁸Al cont. (300 pps)





Experiment – SHARC and TIGRESS at TRIUMF





Microscopic Shell Model 3/2* 0d3/2 2.13 3/2* 1p3/2 1.92 -1f7/2 1.43 0p3/2 1.09 3/2* 1p3/2 0.87 1d3/2 0.05 0s1/2 0.00 Exp Tsunoda

We have preliminary results from this experiment using a heavily contaminated beam.

New shell model calculations with realistic interactions and expanded sdpf model space... Tsunoda, Otsuka EEdf1 (EKK)

Too early to judge agreement.

Future plans – d(⁶⁰Cr,p)⁶¹Cr at 10.0 MeV/u

We have plans to move towards studying the second island of inversion e.g. via ⁶⁰Cr(d,p) at Texas A&M...



B.A. Brown, http://link.aps.org/doi/10.1103/Physics.3.104

Texas A&M– radioactive beams using gas catcher and cyclotron reacceleration

installed, first run

Aug 2016

zero-degree detection using Oxford MDM

THE OXFORD MDM-2 MAGNETIC SPECTROMETER

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Nuclear Instruments and Methods A245 (1986) 230

TIARA for TEXAS



<u>Summary</u>



- We found that just outside the borders of the island of inversion, the shell model that was adapted for the island (i.e. USD-A, wbc) seems to work reasonably well we have very useful discussions with those developing the new EEdf1 interaction
- Even in some less exotic nuclei, the selectivity of (d,p) has been shown to be hugely powerful in identifying the most interesting states (for the first time) e.g. ²⁵Na, and WE STUDY THE SAME orbitals and physics as in much more exotic nuclei.
- The new technique of gating on the coincident gamma rays to separate states that are not otherwise resolved has worked well
- We are edging closer towards the island of inversion to test the shell model further and improve it, and have plans to move attention to the second island of inversion
- We are preparing for new availability of beams at Texas A&M (also HIE-ISOLDE and MUGAST at GANIL)



Summary



- We found that just outside the borders of the island of inversion, the shell model that was adapted for the island (i.e. USD-A, wbc) seems to work reasonably well we have very useful discussions with those developing the new EEdf1 interaction
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In-built normalisation from d(²⁴Na,d)²⁴Na near 70° (lab)



centre of mass angle, degrees

<u>AIM</u>:

Perform fits to spectra of E_x for each of a number of angle bins

Constraints in fitting:

- Excitation energy scale calibration checked using states of known E_x e.g. 3.455 MeV
- Allow for slight shift from angle to angle, in case corrections are imperfect
- Constrain energies of peaks using evidence from E_x spectra and supporting gamma rays
- Constrain the widths guided by simulations (observe little E_x change, but angular dependence)



proton laboratory angle (degrees)