# Monte Carlo approach to string/M theory

## Masanori Hanada 花田政範

KEK Theory Center

## Goal (I) : understand these figures.

power law predicted by SUGRA



energy density of SUSY quantum mechanics



# Goal (2)

Convince you that lattice theorists can give very important contributions to string/M-theory, which usual string theorists can never give.

quantum nature of the black hole, Hawking evapolation, birth of the universe, inflation, multiverse,...

# Apologies

Because we have only 30 minutes...

- I skip all the technical details, e.g. how to put supersymmetry on computer. I explain only basic ideas about physics.
- I can explain only a few topics.
- My explanation is sometimes too simplified and not very precise.

## Plan

(1) Gauge/Gravity duality (AdS/CFT) and Super Yang-Mills

(2) ABJM Theory and M-theory ('membrane mini revolution') (M.Honda's talk)

(3) Expanding universe out of a matrix model? (S.-W. Kim's talk & K.N.Anagnostopoulos's talk)

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#### What is string theory? (Only PERTURBATIVE definition exists)

- Point particles are promoted to (I+I)-dimensional 'string'.
- Open string →gauge fields and infinitely many massive fields

gauge degrees of freedom live on the edges

• Closed string  $\rightarrow$  graviton, tensors and massive fields



- The scattering amplitudes can be calculated.
- Spacetime dimension is 1+9=10.



# Supergravity as a low-energy effective theory

 From the scattering amplitudes, one can determine the low-energy effective action in terms of the <u>massless fields</u>, which reproduce the amplitudes.
 Graviton G<sub>µν</sub>, dilaton φ, NS B-field B<sub>µν</sub>, R-R fields C<sub>µνρ</sub>...

IIA/IIB superstring  $\Rightarrow$  IIA/IIB supergravity

There are supersymmetric black p-brane solutions coupling to p-form tensor. (Horowitz-Strominger 1991)
 (p+1)-d analogues of the black hole
 (p=even for IIA, p=odd for IIB)



Can we understand black p-brane from the perturbative string picture?

## black brane = D-brane

 Dp-brane (Dirichlet p-brane) is a (p+1)-d object on which open string can be <u>attached</u>.



Dirichlet boundary condition along the transverse direction

• It has the same charge as the black p-brane.

D-brane open string picture



closed string picture black brane



Polchinski

bunch of many D-branes = black brane (large-N  $\rightarrow$  heavy and big  $\rightarrow$  classical gravity)

# SYM from D-brane (I)



# SYM from D-brane (2)

 $A_{\mu}$ ,  $\Phi$  and  $\psi$  become N×N matrices

(i,j)-component= string connectingi-th and j-th D-branes

(p+1)-d U(N) Super Yang-Mills

(more generally, the Dirac-Born-Infeld action)

N Dp-branes

## Maximally supersymmetric Yang-Mills (Brink-Schwarz-Scherk 1977) IOd $\mathcal{N}=I$ U(N) SYM $S_{10d} = \frac{1}{g_{YM}^2} \int d^{10}x Tr\left(\frac{1}{4}F_{MN}^2 + \frac{1}{2}\bar{\psi}\gamma^M D_M\psi\right)$ dimensional reduction $S_{(p+1)d} = \frac{1}{g_{YM}^2} \int d^{p+1}x Tr\left(\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{2}(D_{\mu}X_i)^2 + \frac{1}{4}[X_i, X_j]^2 + \frac{1}{2}\bar{\psi}\gamma^{\mu}D_{\mu}\psi - \frac{i}{2}\bar{\psi}\gamma^i[X_i, \psi]\right)$ **Dp-brane** worldvolume SUSY transf. $\delta A_M \sim \bar{\epsilon} \Gamma_M \psi$ $\delta\psi\sim F^{MN}\Gamma_{MN}\epsilon$

### Gauge/gravity duality <u>conjecture</u> (Maldacena 1997)

- In a special limit, both SYM and weakly-coupled string pictures become valid.
- But they are two different descriptions of the same D-brane system. So...



## The dictionary



# SYM

STRING

large-N, strong coupling

large-N, finite coupling

finite-N, finite coupling SUGRA

tree-level string (SUGRA+α')



Quantum string (g<sub>string</sub>>0)

## **SYM**difficult

large-N, strong coupling

large-N, finite coupling

finite-N, finite coupling

# STRING

SUGRA easier

tree-level string (SUGRA+α') more difficult Quantum string (gstring>0) very difficult

## **SYM**difficult

# STRING

large-N, strong coupling

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finite-N, finite coupling SUGRA easier

tree-level string (SUGRA+α') more difficult Quantum string (gstring>0) very difficult



can be useful, if we use Monte Carlo !

Very important remark

 From the string theory point of view, SYM theories in less than four spacetime dimensions are as interesting as four dimensional theories!

(0+1)-d SYM  $\Leftrightarrow$  Black hole (0-brane) (1+1)-d SYM  $\Leftrightarrow$  Black 1-brane, black string (3+1)-d SYM  $\Leftrightarrow$  Black 3-brane (AdS<sub>5</sub>×S<sup>5</sup>)

# D0-brane quantum mechanics

$$S = \frac{N}{\lambda} \int dt \ Tr \Big\{ \frac{1}{2} (D_t X_i)^2 - \frac{1}{4} [X_i, X_j]^2 \\ \mathbf{0} \qquad + \frac{1}{2} \bar{\psi} D_t \psi - \frac{1}{2} \bar{\psi} \gamma^i [X_i, \psi] \Big\}$$

- Dimensional reduction of 4d N=4 (or 10d N=1)
- Strong coupling = low temperature and/or long distance

effective temperature (dimensionless) :  $\lambda^{-1/3}T$ 

• No UV divergence  $\Rightarrow$  no need for lattice

### Prediction from gravity side

$$\left(\frac{1}{N^2}E_{D0} \sim 7.4 \ T^{2.8} \quad (\lambda = 1)\right)$$

at large-N & low temperature (=strong coupling)  $\lambda^{-1/3}T \ll 1.$ 

Temperature in YM = Hawking temperature

SYM should give a microscopic understanding of the black hole thermodynamics

black hole mass vs energy density



Anagnostopoulos-M.H.-Nishimura-Takeuchi 2007, M.H.-Hyakutake-Nishimura-Takeuchi 2008



M.H.-Hyakutake-Nishimura-Takeuchi 2008

### Correlation functions (GKPW relation)



- Recipe to calculate the correlation function at <u>large-N</u> and <u>strong coupling</u> from supergravity (Gubser-Klebanov-Polyakov 1998, Witten1998)
- Similar relation holds also in D0-brane theory.

(Sekino-Yoneya 1999)

 $\langle \mathcal{O}(t)\mathcal{O}(0)\rangle \sim t^{\text{p}} \overset{\text{calculable}}{\underset{\text{via SUGRA}}{\text{via SUGRA}}}$ 



(M.H.-Nishimuea-Sekino-Yoneya 2009,2011)







(M.H.-Nishimura-Sekino-Yoneya 2011)

## (I+I)-d SYM and black hole/black string transition

Susskind, Barbon-Kogan-Rabinovici, Li-Martinec-Sahakian, Aharony-Marsano-Minwalla-Wiseman,...

SYM simulation : Catterall-Wiseman, 2010



Several lattice theories exist (from around 2002-2005)

- Cohen, Kaplan, Katz, Unsal
- Sugino
- Catterall



- D'Adda, Kanamori, Kawamoto, Nagata
- No fine tuning to all order in perturbation. (Kaplan-Katz-Unsal 2002)
- Numerical simulation supprt the absence of the fine tuning at nonperturbative level.

(Suzuki 2007, Kanamori-Suziki 2008, M.H.-Kanamori 2009, 2010)



- Consider 2d U(N) SYM on a spatial circle. It describes N DI-branes in R<sup>1,8</sup>×S<sup>1</sup>, winding on S<sup>1</sup>.
- T-dual picture : N D0-branes in R<sup>1,8</sup>×S<sup>1</sup>.



• Wilson line phase = position of D0



localized distribution = 'black hole'

#### If the black hole becomes larger, then...



- Wilson line phase = position of D0  $W = diag(e^{i\theta_1, \cdots, e^{i\theta_N}})$
- U(I) center symmetry

$$\theta_i \to \theta_i + const.$$

Uniform = center unbroken

$$\left\langle \frac{1}{N} TrW \right\rangle = 0$$

**Non-uniform = center broken**  $\left\langle \frac{1}{N} TrW \right\rangle \neq 0$ 

#### Value of spatial Wilson loop



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# M-theory (conjecture!)

- Strong coupling limit of type IIA superstring.
- Theory of membrane in IId spacetime.
- string = membrane winding on II-th dimension.
- Low-energy effective theory is IId supergravity.







# AdS/CFT correspondence can hold also in **M-theory**.

D3-brane in IIB string  $\rightarrow AdS_5 \times S^5$ M5-brane in M-theory  $\rightarrow AdS_7 \times S^4$ M2-brane in M-theory  $\rightarrow AdS_4 \times S^7$ 



#### ABJM theory (Aharony-Bergman-Jafferis-Maldacena, 2008)



$$k Tr \Biggl\{ \frac{\epsilon^{\mu\nu\rho}}{2} \left( -A_{\mu}\partial_{\nu}A_{\rho} - \frac{2}{3}A_{\mu}A_{\nu}A_{\rho} + \tilde{A}_{\mu}\partial_{\nu}\tilde{A}_{\rho} + \frac{2}{3}\tilde{A}_{\mu}\tilde{A}_{\nu}\tilde{A}_{\rho} \right) \\ + \left( -D_{\mu}\bar{\Phi}^{\alpha}D^{\mu}\Phi_{\alpha} + i\bar{\Psi}^{\alpha}D\Psi_{\alpha} \right) - i\epsilon^{\alpha\beta\gamma\delta}\Phi_{\alpha}\bar{\Psi}_{\beta}\Phi_{\gamma}\bar{\Psi}_{\delta} + i\epsilon_{\alpha\beta\gamma\delta}\bar{\Phi}^{\alpha}\Psi^{\beta}\bar{\Phi}^{\gamma}\Psi_{\delta} \\ + i\left( -\bar{\Psi}_{\beta}\Phi_{\alpha}\bar{\Phi}^{\alpha}\Psi^{\beta} + \Psi_{\beta}\bar{\Phi}_{\alpha}\Phi^{\alpha}\bar{\Psi}^{\beta} + 2\bar{\Psi}_{\alpha}\Phi_{\beta}\bar{\Phi}^{\alpha}\Psi^{\beta} - 2\Psi^{\beta}\bar{\Phi}^{\alpha}\Phi_{\beta}\bar{\Psi}_{\alpha} \right) \\ + \frac{1}{3} \left( \Phi_{\alpha}\bar{\Phi}^{\beta}\Phi_{\beta}\bar{\Phi}^{\gamma}\Phi_{\gamma}\bar{\Phi}^{\alpha} + \Phi_{\alpha}\bar{\Phi}^{\alpha}\Phi_{\beta}\bar{\Phi}^{\beta}\Phi_{\gamma}\bar{\Phi}^{\gamma} + 4\Phi_{\beta}\bar{\Phi}^{\alpha}\Phi_{\gamma}\bar{\Phi}^{\beta}\Phi_{\alpha}\bar{\Phi}^{\gamma} - 6\Phi_{\gamma}\bar{\Phi}^{\gamma}\Phi_{\beta}\bar{\Phi}^{\alpha}\Phi_{\alpha}\bar{\Phi}^{\beta} \right) \Biggr$$

3d U(N)<sub>k</sub>×U(N)<sub>-k</sub> Superconformal Chern-Simons-Matter theory



λ=N/k

By solving ABJM we can understand string and M in a unified manner.

## localization









• Deform the theory, keeping the expectation value of a SUSY-invariant operator unchanged.

#### 'localization technique'

$$S \rightarrow S + tQV$$

$$\uparrow$$

$$QS = 0, \qquad Q\mathcal{O} = 0$$
supercharge
$$\frac{d}{dt} \int [dX]\mathcal{O}e^{-S-tQV} = -\int [dX]\mathcal{O} \cdot (QV)e^{-S-tQV}$$

$$= -\int [dX]Q \left(\mathcal{O} \cdot Ve^{-tQV}\right)e^{-S}$$

$$= Z_{t=0} \times \langle Q\text{-exact} \rangle = 0$$

• At  $t=\infty$ , the path integral 'localizes' to a matrix model.

# localization in ABJM

 The partition function of ABJM on S<sup>3</sup> reduces to 'ABJM matrix model'



(Kapustin-Willett-Yaakov 2009)

 $Z_{ABJM}$ 

$$= \frac{1}{N!^2} \int \frac{d^N \mu}{(2\pi)^N} \frac{d^N \nu}{(2\pi)^N} \frac{\prod_{i < j} \left[ 2\sinh\left(\frac{\mu_i - \mu_j}{2}\right) \right]^2 \left[ 2\sinh\left(\frac{\nu_i - \nu_j}{2}\right) \right]^2}{\prod_{i,j} \left[ 2\cosh\left(\frac{\mu_i - \nu_j}{2}\right) \right]^2} \exp\left[\frac{ik}{4\pi} \sum_{i=1}^N (\mu_i^2 - \nu_i^2)\right]^2$$

• Perturbative part (in string language) can be studied analytically even at strong coupling.

(Drukke-Marino-Putrov 2009-2011, Fuji-Hirano-Moriyama2011)

• Monte-Carlo gives the full answer at any N and any coupling. (M.H.-Honda-Honma-Nishimura -Shiba-Yoshida 2012)

## Planar limit ( $\lambda = N/k$ fixed)



### Large-N, fixed k

#### <u>IId</u> SUGRA has been reproduced!



strong evidence that ABJM gives the definition of M-theory

## Current status

 By carefully analyzing the Monte Carlo results, analytic solution at finite-N and finite coupling, which is applicable to both string and M-regions, have been obtained.

 The I/N correction seems to disagree with string prediction. →made a miscalculation in the string side? have to modify the dictionary? Or gauge/ gravity duality fails at quantum string level?

#### Dark Energy Accelerated Expansion



Development of Galaxies, Planets, etc.

Inflation

# Expanding universe from a matrix model?

Quantum Fluctuations

> 1st Stars about 400 million yrs.

> > **Big Bang Expansion**

13.7 billion years

## IIB matrix model

(Ishibashi-Kawai-Kitazawa-Tsuchiya 1996)

- (0+0)-d SYM
- Matrix regularization of IIB superstring
- Eigenvalue distribution = spacetime ?

$$S = Tr\left(-rac{1}{4}[X_{\mu},X_{
u}]^2 - rac{1}{2}ar{\psi}\gamma^{\mu}[X_{\mu},\psi]
ight)$$



## Expanding universe!

(Kim-Nishimura-Tsuchiya 2011)

• eigenvalue of X<sub>0</sub> = 'time'



• plot the extent of  $X_1, ..., X_9$  at each 'time'.



## Summary

- SUSY can be studied on computer, by combining lattice and non-lattice methods(e.g. Matrix model, localization,..).
- Simulation of the quantum gravity ⇒ black hole thermodymanics, inflation, birth of the universe, multiverse, ...
- Not string theorists, but lattice theorists, can study such exciting topics.

## THE END

How to put Super Yang-Mills on computer

## 'No-Go' for lattice SYM

- SUSY algebra contains infinitesimal translation.  $\{Q,\bar{Q}\}\sim\partial$
- Infinitesimal translation is broken on lattice by construction.
- So it is impossible to keep all supercharges exactly on lattice. Then SUSY breaking radiative corrections appear in general.
- Still it is possible to preserve a part of supercharges. (subalgebra which does not contain  $\partial$ )

## Basic ideas

(Kaplan-Katz-Unsal 2002)

- Keep a few supercharges exact on lattice.
- Use it (and other discrete symmetries) to forbid SUSY breaking radiative corrections.
- Only "extended" SUSY can be realized for a technical reason.
- In (0+1)-d and (1+1)-d, no fine tuning to all order in perturbation.







# (0+1)-d SYM

• Matrix quantum mechanics is UV finite.

#### No fine tuning!

(4d N=4 is also UV finite, but that relies on cancellations of the divergences...)

- We don't have to use lattice. Just fix the gauge & introduce momentum cutoff! (M.H.-Nishimura-Takeuchi, 2007)
- Lattice can also work, of course.

## (2+1)-d maximal SYM

(Maldacena-Sheikh Jabbari-van Raamsdonk, 2003)

 Start with the Berenstein-Maldacena-Nastase Matrix model, which can be formulated without fine tuning.

$$S = \int dt \ Tr\left(\frac{1}{2}(D_t X_I)^2 - \frac{1}{4}[X_I, X_J]^2 + \frac{i\mu}{3}\epsilon^{abc}X_a X_b X_c + \frac{\mu^2}{18}X_a^2 + \frac{\mu^2}{72}X_i^2\right) + (\text{fermions})$$
  
$$I, J = I, ..., 9; a, b, c = I, 2, 3; i = 4, ..., 9$$

• BMN model has (modified) 16 SUSY

Fuzzy sphere  

$$-[X_{b}, [X_{a}, X_{b}]] + i\mu\epsilon^{abc}X_{b}X_{c} + \frac{\mu^{2}}{9}X_{a} = 0$$

$$\longrightarrow \quad X_{a} = \frac{\mu}{3}J_{a}, \quad [J_{a}, J_{b}] = i\epsilon_{abc}J_{c}$$

 preserves I6 SUSY. Around it one obtains (2+1)-d SYM on noncommutative space.

Fuzzy D2 brane out of N D0 branes (R. C. Myers 1999)

• 'Lattice' is embedded in matrices.

Large-N = continuum limit

 With maximal SUSY, commutative limit of the noncommutative space is smooth. (no UV/IR mixing) (Matusis-Susskind-Toumbas '00)





## (3+1)-d SYM

- 4d N=1 pure SYM : lattice chiral fermion assures SUSY (Kaplan 1984, Curci-Veneziano 1986)
- 4d N=4 :
  - again "Hybrid" formulation: Lattice + fuzzy sphere (M.H.-Matsuura-Sugino 2010, M.H. 2010)
  - •Large-N Eguchi-Kawai reduction(Ishii-Ishiki-Shimasaki-Tsuchiya, 2008)
  - •Another Matrix model approach(Heckmann-Verlinde, 2011)
  - recent analysis of 4d lattice:
  - Fine tuning is needed, but only for 3 bare lattice couplings.

(Catterall-Dzienkowski-Giedt-Joseph-Wells, 2011)

## SIGN PROBLEM?

The second se

### NO PROBLEM!

- SYM has the sign problem (execept for 4d N=I pure YM)
- Use the 'phase-quenched' effective action

$$S_{eff}[A] = S_B[A] - \log |\det D[A]|$$

 Phase can be taken into account by the 'phase reweighting' in principle, but usually it's hopelessly hard.

$$\begin{split} \langle \mathcal{O} \rangle &= \frac{\int [dA] \det D \cdot e^{-S_B} \cdot \mathcal{O}}{\int [dA] \det D \cdot e^{-S_B}} \\ &= \frac{\int [dA](phase) \cdot |\det D| \cdot e^{-S_B} \cdot \mathcal{O} / \int [dA] |\det D| \cdot e^{-S_B}}{\int [dA](phase) \cdot |\det D| \cdot e^{-S_B} / \int [dA] |\det D| \cdot e^{-S_B}} \\ &= \frac{\langle (phase) \cdot \mathcal{O} \rangle_{phase \ quench}}{\langle (phase) \rangle_{phase \ quench}} \sim 0/0 \end{split}$$

## Miracles happen in SYM!

- Almost no phase except for very low temperature.
- SU(2) is almost sign-free.

- (Anagnostopoulos-M.H.-Nishimura-Takeuchi 2007, Catterall-Wiseman 2008, Catterall et al 2011, Buchoff-M.H.-Matsuura, in progress.)
- Even when the phase fluctuates, phase quench (without reweight) gives right answer. ('right' in the sense it reproduces gravity prediction.)



- Matrix quantum mechanics is UV finite.
   No fine tuning!
- We don't have to use lattice. Just fix the gauge & introduce momentum cutoff! (M.H.-Nishimura-Takeuchi, 2007)

# (1+1)-d SYM

4 SUSY model (dimensional redcution of 4d N=1; sign-free) has been studied extensively.

- Conservation of the supercurrents. (Sugino model) (Suzuki 2007, Kanamori-Suzuki 2008, Kadoh-Suzuki 2009)
- Comparison with analytic results at small volume & large-N behaviors. (Sugino model) (M.H.-Kanamori 2009)
- Agreement between Sugino model and Cohen-Kaplan-Katz-Unsal model. (M.H.-Kanamori 2010)

All results supports the emergence of the correct continuum limit without fine tuning.

# Supercurrent conservation in the SU(2) Sugino model





Polyakov loop vs compactification radius SU(2), periodic b.c. (M.H.-Kanamori 2010)