Early CI results from Nanten2 towards Carina









Introduction

Carina

- What is it?
- Why are we observing it?
- Observations done towards the region
- Status of the 70m on-the-fly mapping project





The Carina Nebula







What is it?

- Also known as the Great Nebula in Carina, the Eta Carina Nebula, or NGC 3372, as well as "Grand Nebula"
- Extreme stellar population (distance ~2.3 kpc of the sun).
- Evaporating the gas and dust reservoirs which young stars are trying to accrete.
- Up until recently, mid-1990s, Carina was viewed as an evolved HII region devoid of active star formation.
- Formation occurring on the edges of the nebula.





Why study it?

- Carina Nebular (NGC 3372) ideal location to study:
 - Formation of stars (low and high mass)
 - Stars being formed in molecular clouds about 5-10pc away from massive stars (Eta Carinae and HD 93129A)
 - UV radiation and stellar winds
 - An environment that will change rapidly when massive stars go supernova (1-2Myr) and form a giant superbubble
- Different stages of star formation has been observed in this region.
- Only can be observed in the Southern Hemisphere.
- Ionised hydrogen
- Atomic Carbon



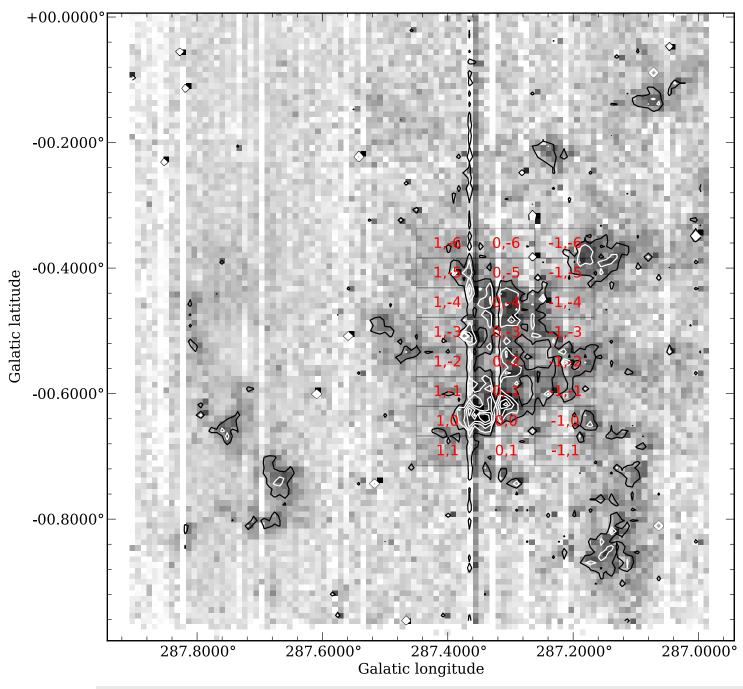


What lines do we observe?

CI 3P1-3P0
 492 GHz
 37 arcsec = 0.010 deg
 CI 3P2-3P1
 809 GHz
 22 arcsec
 CO J=4-3
 460 GHz
 39 arcsec = 0.011 deg
 CO J=7-6
 806 GHz
 22 arcsec = 0.006 deg

- Started observations in CI 3P1-3P0 (492 GHz)
- For more information: http://www.astro.uni-koeln.de/nanten2/node/173

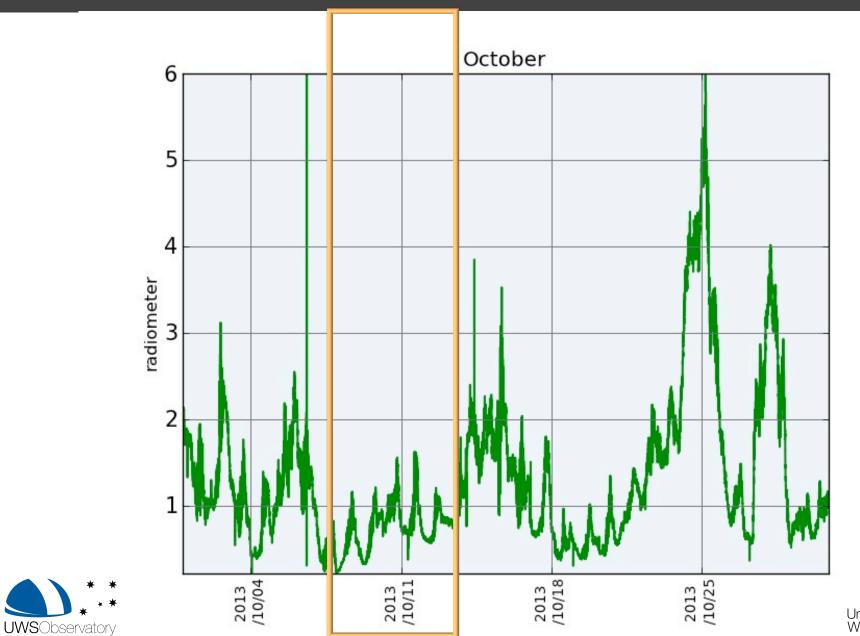




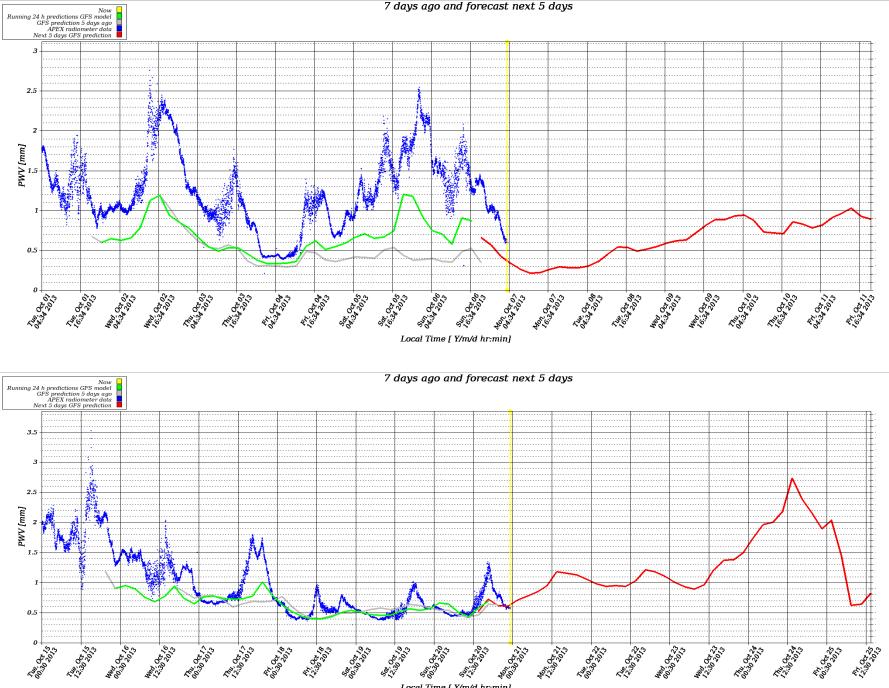
¹²CO 1-0 image taken from Mopra. Overlaid with footprints for observations with NANTEN2.



Water Vapor







Local Time [Y/m/d hr:min]

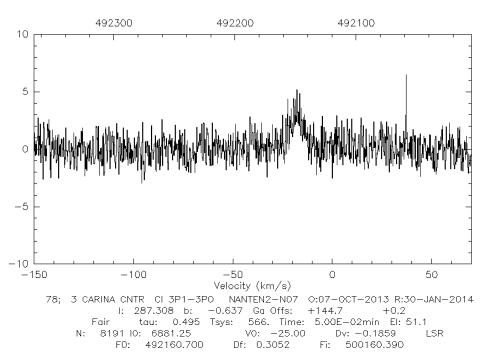
 79; 3 CARINA CNTR CI 3P1-3P0
 NANTEN2-N07
 0:07-0CT-2013
 R:30-JAN-2014

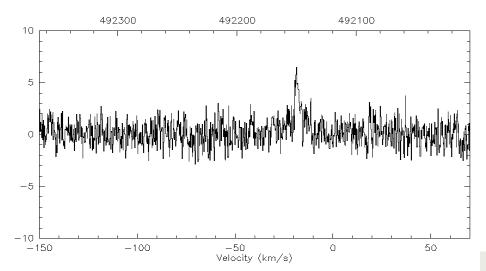
 I:
 287.308
 b:
 -0.637
 Ga
 Offs:
 +153.2
 +0.2

 Fair
 tau:
 0.495
 Tsys:
 566.
 Time:
 5.00E-02min
 El:
 51.1

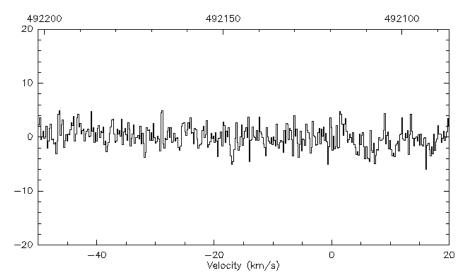
 N:
 8191
 IO:
 6881.25
 V0:
 -25.00
 Dv:
 -0.1859
 LSR

 FD:
 492160.700
 Df:
 0.3052
 Fi:
 500160.390

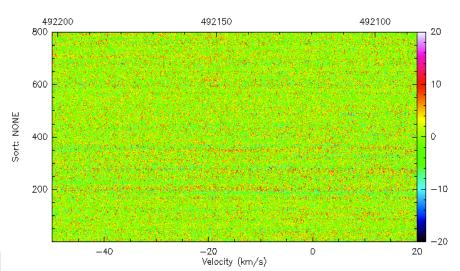




1; 3 CARINA CNTR CI 3P1-3P0 NANTEN2-N00 0:12-OCT-2013 R:30-JAN-2014 I: 287.308 b: -0.637 Ga Offs: -508.7 -256.9 Average tau: 0.686 Tsys: 886. Time: 5.00E-02min EI: 49.9 N: 8191 ID: 6881.25 V0: -25.00 Dv: -0.1859 LSR F0: 492160.700 Df: 0.3052 Fi: 500160.354

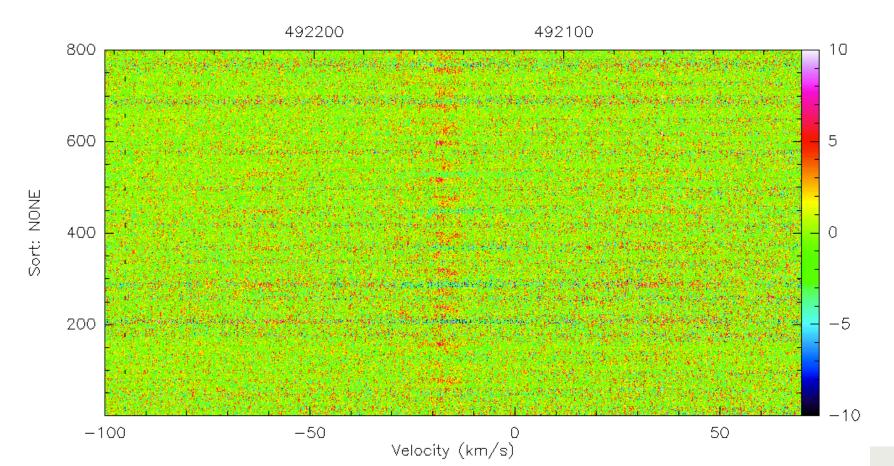


CARINA CNTR I: 287.308 b: -0.637 Ga Scan: 54974-54976 0: from 12-0CT-2013 to 12-0CT-2013 Nspectra: 800 Offset ranges: (-510.3:-176.1) (-256.9:-91.8) N: 32768 ID: 27525.0 V0: -25.0 Dv: -4.648E-02 LSR Cl 3P1-3P0 F0: 492160.700 Df: 7.63E-02 Bef: 1.00 Fef: 0.93 Fi: 500160.354 Gim: 0.500

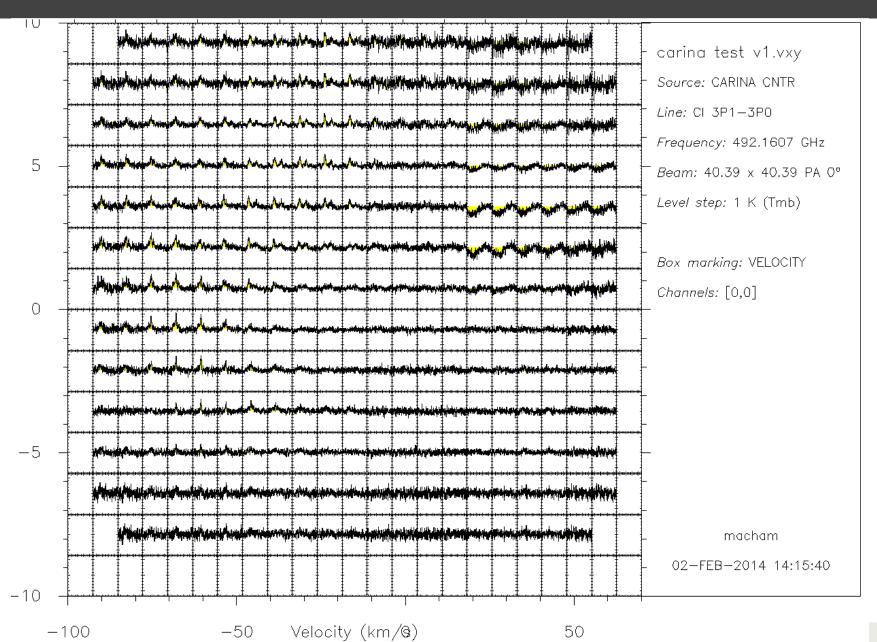


Sample spectra

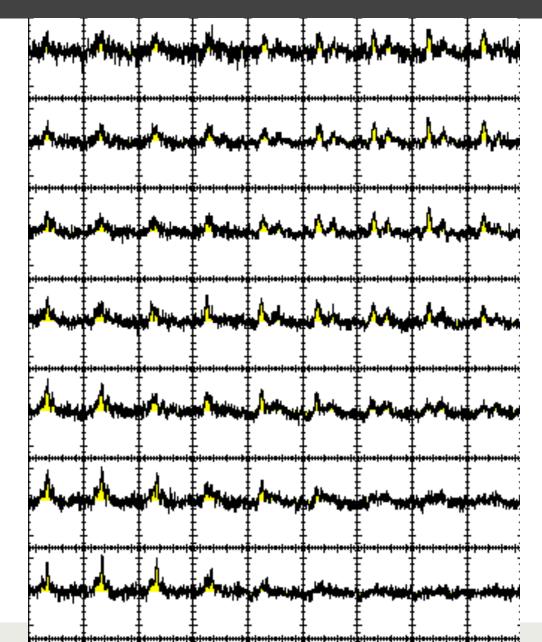
CARINA CNTR I: 287.308 b: -0.637 Ga Scan: 54375-54377 0: from 07-0CT-2013 to 07-0CT-2013 Nspectra: 800 Offset ranges: (-170.2:+163.4) (-85.3:+77.1) N: 32768 I0: 27525.0 V0: -25.0 Dv: -4.648E-02 LSR CI 3P1-3P0 F0: 492160.700 Df: 7.63E-02 Bef: 1.00 Fef: 0.93 Fi: 500160.390 Gim: 0.500



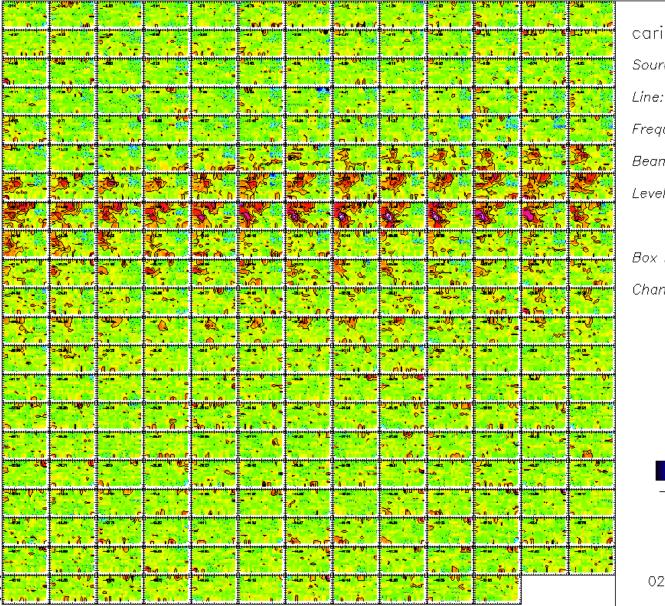
Hot off the press



Hot off the press

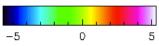


Hot off the press



carina test v1.lmv Source: CARINA CNTR Line: CI 3P1–3P0 Frequency: 492.1607 GHz Beam: 40.39 x 40.39 PA 0° Level step: 1 K (Tmb)

Box marking: VELOCITY Channels: [0,0]



macham 02-FEB-2014 14:14:24

DSS43 – Scientific Capabilities



Spectroscopy in a service observing mode and other single-dish observation

Photo: DSS43 (G. Wong)

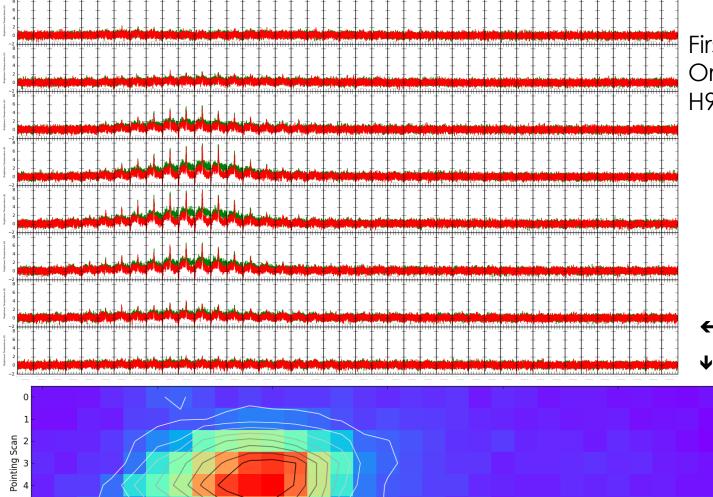
Frequency Range (GHz)	Band (cm)	Receiver	Beam FWHM (arcmin)	Tsys (K)
1.610 - 1.705	18	L-band	8	25
2.270 - 2.300	13	S-band maser	6.4	16
2.200 - 2.300	13	S-band hempt	6.4	25
8.183 - 8.633	3	X-band	1.8	25
18.0 – 26.5	1	K-band	0.8	40

For more information: http://www.atnf.csiro.au/observerstidbinbilla/

UWSObservatory



70m Tidbinbilla On-the-fly mapping



0.0

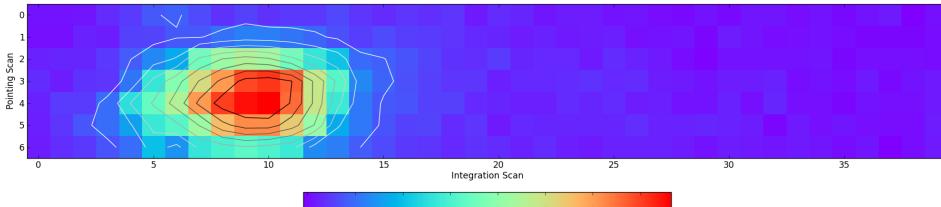
0.6

1.2

First OTF observation - Orion A 2013-07-24 H92 α at 8309.37MHz



←Tiled Spectra ◆Integrated emission map



1.8

2.4

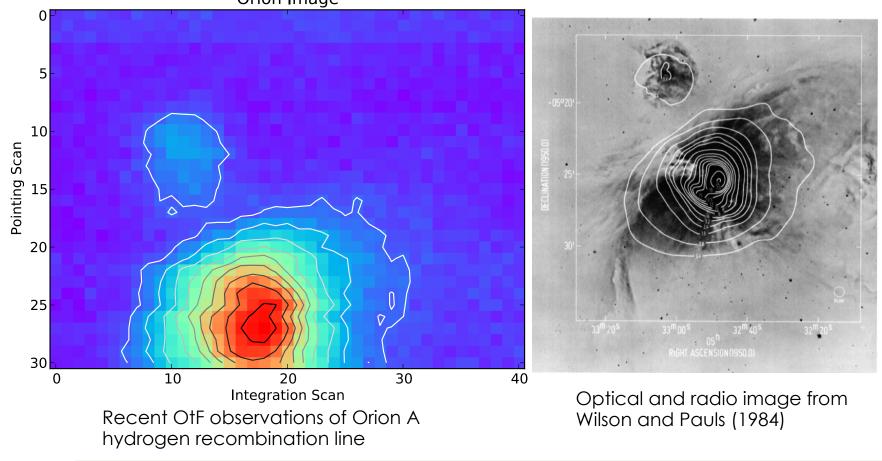
3.0

3.6

4.2

Observing Orion (Round 2)

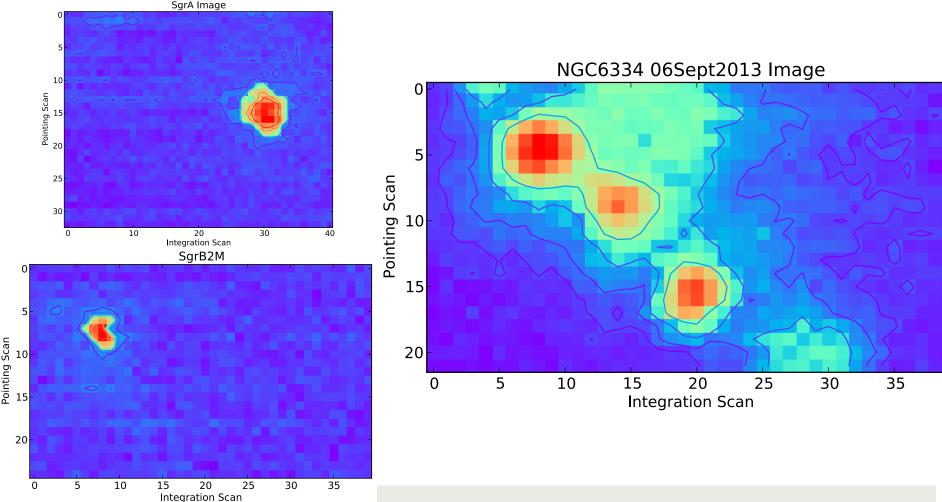
Mapping grid optimisation for 1.8' beam (2013-07-30) H92 α at 8309.37 MHz Orion Image







Other regions: <u>SgrA, SgrB2M & NGC6334</u>







Data Processing

- Using the ASAP component within CASA for data reduction
- Images created via ASCII then Python's Matplotlib package
- CASA for imaging (Thanks Kana-San!)

Imaging with CASA

